

Review 1

Response to the review on "Spectral signatures of the tropical Pacific dynamics from model and altimetry: A focus on the meso/submesoscale range" by M. Tchilibou et al.

We thank the reviewer for taking the time to review our manuscript so thoroughly. We very much appreciate the time, effort, and thought put in to do so. We have found the feedback extremely useful to correct some inconsistencies and imprecisions present in the first version. We are glad that the reviewer finds the paper fundamentally sound to be published after minor revisions. All the comments have been addressed in the last version.

Please find the reviewers' comments below in bold and our responses in non-bold.

Reviewer

The paper is concerned with understanding the dynamical and technical reasons why wavenumber spectra of SSH tend to have flatter spectral slopes in the tropical oceans. It uses primarily analysis of model simulations, with some supporting data analysis. The main conclusion is that the shallow (or flat) spectral slopes at wavelengths less than 200km is mostly due to internal tides and internal waves– this conclusion is supported by the fact that a realistic model simulation with high-frequency winds and tidal forcing has a shallow spectral slope at these large wavenumbers that resembles the observed spectrum from altimetry, whereas a simulation that is similarly realistic except for the inclusion of tidal forcing does not produce the shallow spectral slope (and instead has much less energy at wavelengths below 200km). The conclusion is also supported by various other lines of analysis.

The analysis is sound, and the study provides important information that will be of interest to other researchers. I enjoyed reading the paper and feel the authors did a careful job with the analysis and interpretation. My main criticism is that the "story" is a little complicated and hard to follow; I think this is partly just because it is just inevitably a complicated story, but it may be possible to make changes to the presentation to clarify the arguments. Beyond this, I have minor reservations concerning small details of terminology and interpretation. While I give many critical comments below, I think the paper is fundamentally sound, and that it should be straightforward to address my comments. I recommend that the paper be published after minor revisions.

Major comments:

(1) Lines 114-115: For testing the choices people make in data processing, it would be better to use a model with the most realistic signal possible. Using output from an unrealistic model (no tides) might give misleading guidance about how a particular analysis method should perform when using real data.

We agree that it would be better to make the tests in data processing by using a model with the most realistic signal possible. It corresponds with the R36Th model which is a regional model of the Solomon Sea. Because our study wants to deal with the entire tropical Pacific, we find more appropriate to use the G12d5 model. We may hope that this choice is not bad because the leakage effect concerns mainly the distortion at high wavenumbers of energetic low wavenumbers, and the G12d5 model is well suited

to get the large scale motions. Also, the comparison of spectra computed from G12d5 and R36 in Fig. 10 gives confidence on the choice we make.

(2) Line 228: This is the first occurrence of the term "isotropic spectrum". The term is not defined in the paper, but I think the term is being inappropriately used to mean "wavenumber magnitude spectrum". I think that what the authors mean is that it is the wavenumber spectrum after azimuthally integrating the 2D wavenumber spectrum. That is a wavenumber magnitude spectrum. Real data (or GCM data), in general, will not have an isotropic spectrum, which is a spectrum that is the same regardless of what direction a line of samples is taken (e.g., zonal versus meridional).

Yes, we agree. We suppress here the term isotropic, and when it appears again in the section 4.3, we use your terminology. We have used this term because it was been used in different papers instead of wavenumber magnitude spectrum as in Arbic et al. (2014):

The theory of spectral transfers $T(K, \omega)$ and spectral fluxes $\Pi(K, \omega)$ in the two-layer QG model, where $K = \sqrt{k^2 + l^2}$ is isotropic wavenumber and ω denotes frequency, is summarized here. We let \hat{A} denote the

(3) Line 232: I think this statement must be wrong. That is, I do not believe that the second part of the sentence logically follows from the first. It is true that using shorter segments reduces the maximum energy (because the spectrum is red and the segments are detrended). But, this should only affect the spectral slope via analysis artifacts. One analysis artifact that could affect the spectral slope as record length changes is spectral leakage or sidelobes— reducing the record length should increase the leakage from energetic low wavenumbers to weaker high wavenumbers, thus decreasing the spectral slope. The relationship between record length and the leakage is discussed in the book Random Data by Bendat and Piersol (2000 edition, p. 400).

Yes, we agree that this statement is wrong. It lacks the link with the leakage effect. We add the Bendat and Piersol reference. The sentence has been changed in such a way:

"Using shorter segments than this reduces the maximum energy and should increase the leakage from energetic low wavenumbers to weaker high wavenumbers, thus decreasing the spectral slope (Bendat and Piersol (2000))."

(4) Lines 267-281: I don't like this paragraph very much. It is a long and convoluted way of saying that the spectral leakage from low to high frequencies is worse with the tk01 taper and with shorter segments, which careful analysts will know or realize for themselves. There are also some statements that seem out of line with the conventional understanding of taper windows and spectral leakage: (a) I think the statement that the 10%-cosine taper causes artifacts at wavelengths around 10% of the record length shows a misunderstanding of the effects of taper windows. Try applying this taper to white noise, and then examine the spectrum. There will be no distortion at short wavelengths. The distortion arises from spectral leakage (side lobes) from more energetic frequencies, and it thus has as much to do with the spectrum of the untapered time series as the taper. The authors seem to believe that the distortion reliably occurs at wavelengths that are about 10% of the domain size for the 10%-cosine taper, but this is just a coincidence (and I do not think Fig 2 actually shows such behavior). The taper window affects all frequencies in the same way— by convolving the Fourier transform of the taper window with the Fourier transform of the input time series (e.g., appendix of Harris, 1978 paper mentioned above). (b) Again, I think the same misunderstanding is reflected in the statement that, "The advantage of Tk05 is in retrieving the large-scale peaks which are smoothed with the Hanning filter window". The Hanning taper window does not preferentially alter low frequencies. (c) The presence of the "unfiltered spectral fluctuations at small scale" in the doubly periodic spectra is a

separate issue, but I doubt that is an inherent problem with the technique— instead, I suspect some numerical issue may have arisen in the way it was implemented. In the above, my tone is critical, but I appreciate the point the authors are making about how the previous analyses with short segments and mild tapers (like 10%-cosine) are potentially (or likely) contaminated by spectral leakage from the energetic low frequencies. For a rectangular taper window (no taper) of length T , the sidelobes of the Fourier transform of the window function (i.e., of $\text{sinc}(fT)$) decay as $1/(fT)$. So, the sidelobes of a particular spectral peak will decay as $1/(fT)^2$, which means that longer records have less severe leakage effects. You can find something similar to this discussed in Bendat and Piersol, *Random Data*, 2000, p. 440. I think the message the authors are trying to convey is that, to safely avoid leakage in the tropics, it is best to use a long record and an effective taper window. It's a good point, and I commend the lead author for his attention to this detail. This was one of the things I liked most about the paper.

We greatly appreciate your fruitful comment. Yes, we agree that there is misinterpretation of the effect of the X% cosine taper on spectra, and some misleading interpretations of technical issue. We have made a lot of sensitivity tests and checked as most possible our results, but we got a little lost. We thank the reviewer of the reference (Harris, 1978), we have this paper. The text has been changed and we hope the story is now clearer and easier to follow.

"The particular sensitivity of spectra in the tropics to the spectral segment length and windowing is linked to energetic EKE and SSH signals extending out to longer wavelengths, and illustrates the ability to deal with spectral leakage from low to high wavenumbers. Tk01 is the worst tapering window, and the distortion of spectra is amplified for short data segments. Tk05, and Hann are a good compromise for preserving much of the original signal and reducing leakage, but need to be applied over larger segments.

So, to safely avoid leakage in the tropics, it is best to use a long record and an effective taper window. We do not advise to use the Tk01 filter window. The Tk05 or Hann filters give convincing results in the equatorial band, with a minimum of 15° to 20° needed in segment lengths. In the off-equator region, 10° data segments or 10°X10° boxes are sufficient. We choose to use the Tukey 0.5 filter in the paper."

(5) Lines 382-397: This whole paragraph strikes me as weak speculation, resorting to an exotic and potentially unnecessary mechanism. Energy at shorter periods and the nondispersive line can also occur as a simple result of westward mean flow, which seems more plausible to me. (Most of the region was mean westward flow, the SEC occupies most of the 10-20S region and the NEC covers a lot of the 10-20N region, and especially the part with intraseasonal SSH variance.) Farrar and Weller (2006, JGR) examined the effect of the NEC on the Rossby wave propagation and instabilities near 10-13N. Note that, while "linear Rossby wave theory" is often taken to mean the quasigeostrophic equations linearized about a state of rest, one could linearize about any background flow state and it would still be linear Rossby wave theory. Thus, a nondispersive ridge arising from the effect of a westward mean flow on the wave propagation would still be consistent with linear Rossby wave theory.

We understand the criticism of the reviewer about the interpretation of the non-dispersive line as the signature of coherent vortices. It was a subject of debate during several years. Because, these spectra look like spectra in mid latitudes, the discussion was to suggest the presence of eddies to discuss on a

possible inertial range. But we mention the weakly nonlinear regime of the region. Also, we didn't know your reference (Farrar and Weller, 2006) that demonstrate how the dispersion curve of Rossby waves follows similar non dispersive line. Therefore, this paragraph has been rewritten to take into account your comments.

"Although linear Rossby wave theory provides a first - order description of the EKE spectra, in both hemispheres energy extends to higher frequencies (Fig. 4a), and as the wavenumber and frequency increases, significant deviations from the baroclinic dispersion curves occur (Fig. 5a,c). Much of the energy lies approximately along a straight line called the 'non dispersive line' in wavenumber–frequency space as it implies non-dispersive motions. The wavenumber dependencies along the 'non dispersive line' could be the signature of non-linear eddies (Rhines, 1975). The westward propagation speed is estimated at ≈ 10 cm/s, close to the eddy propagation speed found in this latitudinal range by Fu (2009) and Chelton et al. (2007). But these regions are defined as a weakly nonlinear regime (Klocker and Abernathey, 2014). In this region of mean zonal currents the dispersion curves experience Doppler shifting by the zonal flow which makes the variability nearly non dispersive (Farrar and Weller,2006) . So, the non-dispersive line could account both for coherent vortices and more linear dynamics such as Rossby waves or meandering jets propagating westward (Morten et al., 2017). "

(6) Line 422, "[The fact that the meridional EKE spectrum has larger values than the zonal one] reveals the energetic meridional perturbations due to instabilities of the larger-scale zonal currents". This isn't obviously true, and I don't see any support given for this interpretation. It is widely appreciated that scales of variability near the equator tend to be larger in the zonal direction than in the meridional direction. This is true of many kinds of variability (mean currents, inertia-gravity waves, Kelvin waves, Yanai waves, TIWs). Instead, one might say that "[The fact that the meridional EKE spectrum has larger values than the zonal one] is consistent with the widely held notion that scales of variability near the equator tend to be larger in the zonal direction than in the meridional direction for many kinds of variability (mean currents, inertia-gravity waves, Kelvin waves, Yanai waves, TIWs)."

You are completely right, and it was also our interpretation. This sentence is a big mistake, and we correct it using your suggestions.

"...the meridional EKE spectrum has a higher level of energy than the zonal one (Fig. 7b). This reflects a shift of energy towards the smaller scales in the meridional direction that is consistent with the widely held notion that scales of variability near the equator tend to be larger in the zonal direction than in the meridional direction for many kinds of variability (mean currents, inertia-graviy waves, Kelvin waves, Yanai waves, TIWs)."

(7) Lines 437-438, "So, poleward of 10° the hypothesis of isotropy seems to be relevant for scales up to 250 km even if the flow is supposed to be weakly nonlinear". I don't understand the logic. I also really doubt this statement is true. The beta effect is still relatively strong on 10-20N, so I would expect the flow to still be sensitive to beta and to not be isotropic on at least 100-200km scales. Maybe I am misunderstanding what is meant by isotropic, in which case the term should be defined.

Our definition of an isotropic spectrum is the same than yours that you give in the remark 2: It is a spectrum that is the same regardless of what direction a line of samples is taken.

The spectra on Fig. 7ac are similar whatever the direction: zonal, meridional, and with the magnitude spectrum. So it argues for isotropy. We were a little bit surprised because isotropy is the signature of nonlinear eddies in mid latitudes, at first. So it is the reason why we have written this sentence: " So, poleward of 10° the hypothesis of isotropy seems to be relevant for scales up to 250 km even if the flow is supposed to be weakly nonlinear". We change a little this paragraph.

"In the 10°x10° off-equatorial boxes, the energy at long wavelengths is greatly reduced compared to the equatorial band. The peak of the EKE spectra corresponds to a wavelength of 300 km. Yet the zonal, meridional and magnitude EKE spectra are similar for wavelengths up to 250 km (Fig. 7a,c). So, poleward of 10° the hypothesis of isotropy seems to be relevant for scales up to 250 km even if the flow is supposed to be weakly nonlinear, and sensitive to beta effect (Klocker and Abernathey, 2014). The EKE slope over the redefined mesoscale range from 100 to 250 km is between -2 and -3 which lies between the prediction of SQG and QG turbulence."

Minor comments:

(1) Lines 59-60: It is not true that the deformation radius is theoretically infinite at the equator. (Or perhaps one could say that the deformation radius is finite in the most commonly used theoretical approaches to equatorial dynamics.) Look for the "equatorial Beta-plane approximation" or the "shallow water equations on the equatorial Beta-plane". It must be in most textbooks on oceanic or atmospheric dynamics.

Yes, we agree. We just referred to the definition for the off equatorial region but it is not correct. We suppress this part in the text.

(2) Line 70: The phrase "representativeness of SSH to infer the tropical dynamics" doesn't have any clear meaning to me. The SSH (pressure) field is a fundamental dynamical variable in itself. I think the point you are making is that, in the tropics, the assumption of geostrophic balance is much more questionable than in midlatitudes.

Yes, it is what we mean. The text has been changed for clarity.

"Also, the tropics are characterized by strong ageostrophic flow, and the representativeness of geostrophic balance from SSH to infer the tropical dynamics needs to be checked."

(3) Line 85: "Unresolved" has a couple of technical meanings (related to sampling and modeling), and I am not sure that either one of these is the intended meaning here.

Thank you for your checking. We have changed by "unsolved"

(4) Line 89: It is imprecise phrasing to say that "tidal and supertidal signals... greatly exceed the internal dynamics at scales less than 300 km wavelength". First of all, I do not understand why internal waves are not considered internal dynamics. Second, and less importantly, it would be clearer to say "supertidal SSH signals... greatly exceed the signals from internal dynamics".

We agree that "internal dynamics" is confusing here. It was a wrong way to separate low and high frequency motions. We change it by "subtidal dynamics", and the text has been changed to highlight the signature of supertidal signal in accordance with Savage et al. (2017).

"Recent results from a high-resolution 1/48° model highlight that the tidal and supertidal signals in one region of the equatorial Pacific greatly exceed the subtidal dynamics at scales less than 300 km wavelength, and supertidal phenomena are substantial at scales approximately 100 km and smaller (Savage et al. 2017)."

(5) Line 238: Isn't a 50% cosine taper (Tukey) a "full cosine taper" or a Hann window? The authors should be specific about whether they mean 50% at each end or 25% at each end when they say 50%. (If it is the latter, I agree the two are not the same.) Since the authors seem interested in taper windows, they may be interested in this paper: Harris, F.J, 1978. On the use of windows for harmonic analysis with the discrete Fourier transform. Proceedings of the IEEE., vol 66, p.51.

It is a 50% cosine taper. It means that there are 50% of the coefficients that are smaller than 1 (25% at each end). We thank the reviewer for the reference. We know it.

(6) Line 236: "Han" should be "Hann"... or is this just meant to be an abbreviation for "Hanning"?

Okay, we change Han by Hann

(7) Line 349: I think Lee et al. (2018) is a useful reference for this statement (perhaps better than Willet, 2006). By the way, it should be Willet et al. (2006). T. Lee, J.T. Farrar, S. Arnault, D. Meyssignac, W. Han, and T. Durland. Monitoring and interpreting the tropical oceans by satellite altimetry. In D. Stammer and A. Cazenave, editors, Satellite Altimetry Over Ocean and Land Surfaces. CRC Press, Taylor and Francis Group, 2018.

It is clearly a good reference. I have asked for the book and I have added this reference.

(8) Lines 351-352. Having equal amounts of energy propagating in opposite directions ('balanced northward and southward propagation') is a hallmark of standing modes—the TIWs largely take the form of standing meridional modes, as seen from other perspectives in Lyman et al, (2005) and Farrar (2008, 2011) and earlier work.

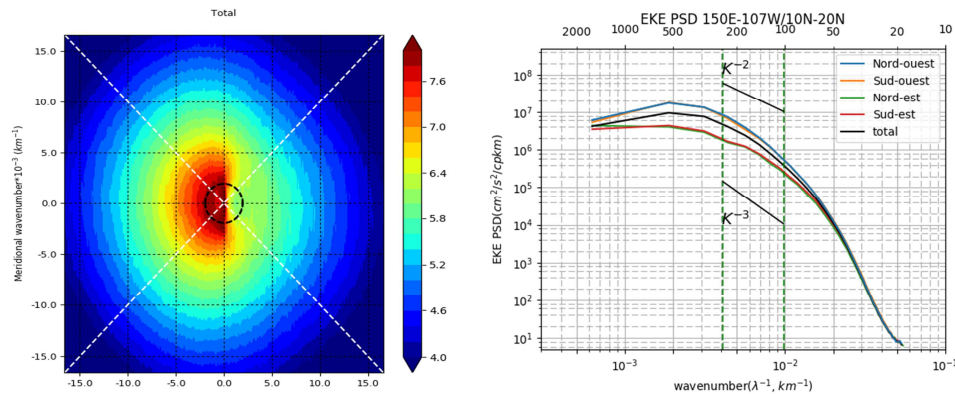
Yes, we thank the reviewer for their constructive comments. The text has been changed:

"have a meridional propagation with northward and southward motion roughly balanced that is a hallmark of standing meridional modes for TIWs as seen from others perspectives in Lyman et al. (2005) and Farrar (2008, 2011) and earlier work"

(9) Lines 400-401, about the steep spectral slopes being consistent with an inertial subrange. OK, but isn't there a difference between positive and negative values of zonal wavenumber? How does that fit with an inertial subrange? (I don't know the answer, but I suspect it isn't so simple.)

Yes, we agree that the sensitivity of the spectra to positive and negative values of zonal wavenumber is not discussed here, and that is a question of interest in our region. In this paper, we present only wavenumber spectra that are commonly discussed in most of the literature. This question is analysed as part of my Phd thesis where I give more details of the 3D spectra as the one below for the corresponding region showing meridional and zonal propagation. As you mentioned, there is a strong anisotropy in the zonal direction with dominant westward propagations for wavelengths higher than

200 km. It affects the horizontal wavenumber spectra as shown below. Considering the westward propagation only, steeper the slope that looks like a straight line with a -3 slope.



(10) Figure 7, use of the term "isotropic spectra"– see major comment 2.

We have changed the terminology in all the text according your major comment 2.

(11) Lines 443-453: I don't feel this paragraphs adds a lot, and I think it makes tenuous connections to midlatitude dynamics. (a) It is a trivial truth that the change in slope of the ridge is related to the change in wave speed (these are essentially the same thing if the ridge roughly makes a line through $k=0$, $\omega=0$). In addition, different equatorial wave modes and TIW modes, which have different meridional structures and extents, should feel the equatorial currents differently. The strong equatorial currents are almost surely an important factor influencing the propoagation of variability having phase speeds less than 1 m/s, so I do not think it is a good idea to try to explain the latitudinal change in the w - k spectrum as being due to changes in Beta and the deformation radius alone.

The comment here is close to the major remark 5 where the reviewer doubts on the relation with mid latitude dynamics. We agree that we need to be aware of the limit of our spectra and the text has been changed in consequence. This paragraph is a summary of the previous discussion in regard to the literatures.

“Our modeled zonal frequency-wavenumber spectra differ strongly across the equatorial and off equatorial regions. They show a good representation of the tropical wave and TIW/TIV dynamics. The slope of the ridge of westward variance in the zonal k - ω spectrum in Fig. 5 increases towards the equator. As the slope becomes steeper, more power is concentrated at lower wavenumbers. The change in slope of the ridge itself is mainly related to the change in deformation radius, and expresses linear or non-linear variability propagating non-dispersively (Wortham and Wunsch, 2014). The equatorial region differs from the off equatorial regions in having strong anisotropy with mainly zonally oriented structures (Fig. 7), higher energy at long wavelength due to the strong activity of long equatorial waves, and an overlap between geostrophic turbulence and Rossby wave time scales that produces long waves and slows down the energy cascade to eddies with scales consistent in the tropics with a generalized Rhines scale (L_r) (Theiss, 2004, Tulloch et al., 2009; Klocker et al., 2016; Eden, 2007).”

(12) Lines 454-462: After several readings of the paragraph, I think I understand the intended point: (i) Geostrophic balance is still an important factor near the equator, but the validity of geostrophic turbulence near the equator is questionable; (ii) The model spectra show contrasts between the equatorial and off-equatorial regions... Maybe this kind of rewording would help make it clearer.

We are sorry, part of this paragraph is not at the right place: Discussion about geostrophy/ageostrophy is in the next section 5.

The text has been corrected.

(13) Lines 466-468: I don't see the point of including this first sentence. The first part of the sentence is contradicted by the second part of the sentence. The SSH is a measure of the surface pressure field, an important dynamical variable, which may play a role in both geostrophic and ageostrophic motions.

Yes, we agree with your comment. This sentence is used to introduce the SSH variable that is the main purpose of this section. We change it by:

"The SSH is a measure of the surface pressure field, an important dynamical variable, which may be balanced in the tropics by both geostrophic and ageostrophic motions. The ocean circulation is classically inferred from altimetric SSH through the geostrophic equilibrium. Here, we consider how the wavenumber spectra of geostrophic currents (EKEg) differ from that of the total currents analyzed in section 4."

(14) Lines 471-472: I believe the so-called equatorial geostrophic approximation is of limited validity (only valid at low frequencies).

We think that it could be valid considering the G12d5 simulation. Picaut et al. (1989) have shown that the equatorial geostrophic approximation is not valid for periods shorter than 10 days, and most of these frequencies are filtered when using the G12d5 model.

(15) Lines 478-479: Something seems illogical about the statement that the total EKE is weaker than the geostrophic EKE. This must mean that the geostrophic EKE is not a useful concept in this case, ie, there must be a lot of variability that is not in geostrophic balance.

We were also surprised by this result but Ponte et al (2013) describe such behavior as an effect of wind-driven mixed-layer dynamics. We change the reference in the text that was not the good one.

"However, in all regions, the total EKE is steeper than the geostrophic EKE at scales from 250 km down to the 20 km resolved by the model. In mid latitude regions Ponte et al. (2013) also noted stronger geostrophic EKE at small wavelengths (and weaker spectral slopes) compared to upper ocean EKE spectra associated with wind-driven mixed layer dynamics."

(16) Lines 493-494, "Due to the strong ageostrophic component in the equatorial region, SSH spectra exhibit lower spectral power than in the off-equatorial region."->(a) I don't understand the link here. (b) Could it also have to do with the fact that f is small? (The authors don't need to respond to this.)

We want to say that although the EKE signal in the equatorial region is higher than in the off equatorial region, it is the opposite for the SSH signal. We agree that we have done a misinterpretation. The sentence has been changed.

"It is notable that although the level of energy is higher in the equatorial region than in the off-equatorial regions, the SSH variability is lower for wavelengths smaller than 500 km. The reduced SSH variability of the low frequency motions (> 10 days) in the G12d5 model is not in agreement with the higher small "scale" SSH levels in altimetry to be discussed in the next section (section 5.2)"

(17) Lines 551-552: Ok, but here "high frequency" means periods <48 hours. However, atmospherically forced internal waves in the equatorial region can have periods much longer than this (like 3, 4, 5, 7, and 14 days).

Yes, we agree

(18) Line 592: It is odd phrasing to say a "flat spectral peak". I assume the authors meant a "flat spectral slope".

Yes, corrected

(19) Line 608: El Nino or La Nina?

Yes!! El Nino!

(20) Line 612: delete "eddy"

Done

(21) Lines 612-613: Also, the inertia-gravity waves examined by Farrar and Durland had very large zonal wavelengths.

Yes, this reference is wrong, and has been deleted

(22) Line 636: It should say "structures tend"

Corrected

(23) Line 640: it should say "spectra are".

Corrected

(24) Line 648-649: How do we know there is a spectral cascade?

Yes, this term is not appropriate. The text has been changed

"In the equatorial band from 10°S-10°N, the total EKE is more energetic than the off-equatorial region, and the EKE spectral slope approaches k^{-3} over a large wavenumber range, from 100 to 600 km ..."

(25) Lines 662-664: I find this confusing. There is no way using geostrophic currents changes the SSH spectrum.

We agree that the last sentence is confusing. Here we want to make the link with SSH, the text has been changed.

"So using SSH and geostrophic currents slightly flattens the EKE wavenumber spectra, but the modeled SSH wavenumber spectra maintain a steep slope that doesn't match the observed altimetric SSH spectra"

(26) Line 678: I do not understand what a "turbulent spectral slope" is.

We agree, and delete this term

(27) Line 709: I think "predominate" may not be the right word choice here. (I'm not sure what is intended.)

We change the word by "major"

(28) Line 420: Just to be precise, this should say "wavelengths" instead of "scales".

Done

Review 2

Response to the review on "Spectral signatures of the tropical Pacific dynamics from model and altimetry: A focus on the meso/submesoscale range" by M. Tchilibou et al.

We thank the reviewer for taking the time to review our manuscript. Please find the reviewers' comments below in bold and our responses in non-bold.

Reviewer

The scientific goal of the paper is to explain the reason why the observed SSH wave number spectra exhibit flatter slopes in the tropics. The dynamical waves the authors address for generating these flat slopes are internal tides and waves. They use high resolution numerical models with and without internal tides and waves to infer that the latter waves are responsible for the flattening below the 200km length scales. They compare the slopes obtained with observed satellite SSH data. Conclusions are relevant, spectral slopes tend to match the satellite observed slopes (Figure 10).

With some modifications, I recommend the paper to be published in Ocean Science journal.

We are glad that the reviewer finds the paper fundamentally sound to be published. Reviewer 2 has made only a few remarks compared to reviewer 1, and a lot of modifications have been made to take into account the remarks of reviewer 1 that should satisfy reviewer 2.

Major comment: The paper deals with model descriptions and a large part deals with technical issues to access correct filtered SSH wave number spectra for model and satellite datas. Most of the paper deals with these technical issues and makes the paper hard to read. Simplifying those technical issues would enlighten the paper greatly. The discussion around all those different types of filtering should be transferred to an appendix.

These technical issues are an important piece of this paper. We agree that this part could be hard to read, especially since some misinterpretations were present. This discussion has been improved thanks to comments from reviewer 1. We think that this section is important for the paper, but have followed the reviewer's remarks and moved most of the details on the model configurations to Annexe 1, and the Spectral sensitivity tests to Annexe 2. The old Figure 2 has also been moved to the Annexe 2. We trust that the main text is a bit easier to follow now, and the important details are still included in the Annexe for the interested readers.

Suggestion: The dynamical discussions could be improved if scaled equations are added to the text to exhibit the importance of the different types of waves present and the linear or nonlinear behaviour of the processes acting at different length and time scale.

We understand the purpose of this suggestion, and we have tried to give in the text the necessary information to present the different dynamics that are discussed. We think that added equations and their comments will extend the paper too much, and we prefer to refer to the different references.

Minor comment: Figure 1 anotations are hard to read.

We will try to do our best to get the best quality for this figure. But these two plots come from published paper to introduce our discussion and it is difficult to change them

1 **Spectral signatures of the tropical Pacific dynamics from model and altimetry: A focus on the**
2 **meso/submesoscale range**

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27 **Abstract**

28 The processes that contribute to the flat Sea Surface Height (SSH) wavenumber spectral slopes
29 observed in the tropics by satellite altimetry are examined in the tropical Pacific. The tropical
30 dynamics are first investigated with a $1/12^\circ$ global model. The equatorial region from $10^\circ\text{N} - 10^\circ\text{S}$ is
31 dominated by Tropical Instability Waves with a peak of energy at 1000 km wavelength, strong
32 anisotropy, and a cascade of energy from 600 km down to smaller scales. The off-equatorial regions
33 from $10\text{-}20^\circ$ latitude are characterized by a narrower mesoscale range, typical of mid latitudes. In the
34 tropics, the spectral taper window and segment lengths need to be adjusted to include these larger
35 energetic scales. The equatorial and off-equatorial regions of the $1/12^\circ$ model have surface kinetic
36 energy spectra consistent with quasi-geostrophic turbulence. The balanced component of the
37 dynamics slightly flatten the EKE spectra, but modeled SSH wavenumber spectra maintain a steep
38 slope that does not match the observed altimetric spectra. A second analysis is based on $1/36^\circ$ high-
39 frequency regional simulations in the western tropical Pacific, with and without explicit tides, where
40 we find a strong signature of internal waves and internal tides that act to increase the smaller-scale
41 SSH spectral energy power and flattening the SSH wavenumber spectra, in agreement with the
42 altimetric spectra. The coherent M2 baroclinic tide is the dominant signal at ~ 140 km wavelength. At
43 short scales, wavenumber SSH spectra are dominated by incoherent internal tides and internal waves
44 which extend up to 200 km in wavelength. These incoherent internal waves impact on space scales
45 observed by today's alongtrack altimetric SSH, and also on the future SWOT 2D swath observations,
46 raising the question of altimetric observability of the shorter mesoscale structures in the tropics.

47

48

49 **1. Introduction**

50 Recent analyses of global sea surface height (SSH) wavenumber spectra from alongtrack altimetric
51 data (Xu and Fu, 2011, 2012; Zhou et al., 2015) have found that while the mid-latitude regions have
52 spectral slopes consistent with quasi-geostrophic (QG) theory or surface quasi-geostrophic (SQG)
53 theory, the tropics were noted as regions with very flat spectral slopes (Fig. 1a). The objective of this
54 paper is to better understand the processes specific to the tropics that contribute to the SSH
55 wavenumber spectral slopes observed by satellite altimetry, particularly in the “mesoscale” range at
56 scales < 600 km, 90 days (Tulloch et al., 2009).

57
58 Only a few studies have addressed the tropical dynamics at spatial scales smaller than this 600 km
59 cutoff wavelength. The tropics are characterized by a large latitude-dependent Rossby deformation
60 radius (L_d) varying from 80 km at 15° to 250 km in the equatorial band ~~and theoretically infinity at~~
61 ~~the equator~~. Different studies have clearly distinguished the tropical regions dominated by linear
62 planetary waves from the mid-latitudes dominated by non-linear regimes (Fu, 2004; Theiss, 2004;
63 Chelton, 2007). Close to the equator, baroclinic instability is inhibited while barotropic instability
64 becomes more important (Qiu and Chen 2004), and mesoscale structures arise from the baroclinic
65 and barotropic instabilities associated with the vertical and horizontal shears of the upper circulation
66 (Ubelmann and Fu, 2011; Marchesiello et al., 2011). This distinct regime in the tropics raises many
67 questions on the representation of the meso/submesoscale tropical dynamics in the global analyses
68 of alongtrack altimetric wavenumber spectra. How are these complex, f-variable zonal currents
69 folded into alongtrack wavenumber spectra, calculated in $10 \times 10^\circ$ bins with a dominant meridional
70 sampling in the tropics? Also, the tropics are characterized by strong ageostrophic flow, and the
71 representativeness of geostrophic balance from SSH to infer the tropical dynamics needs to be
72 checked.

73
74 Another dynamical contribution that could flatten the SSH wavenumber spectra in the tropics is
75 associated with high-frequency processes. In altimetric SSH data, the high-frequency barotropic tides
76 are corrected using global barotropic tidal models, and in the tropics away from coasts and islands,
77 these barotropic tide corrections are quite accurate (Stammer et al., 2014). Altimetric data are also
78 corrected for the large-scale rapid barotropic response to high-frequency atmospheric forcing (< 20
79 days), the so-called Dynamical Atmospheric Correction, using a 2D barotropic model forced by high-
80 frequency winds and atmospheric pressure (Carrere and Lyard, 2003). With only 10 to 35-day repeat
81 sampling, altimetry cannot track the evolution of these rapid barotropic processes, and a correction
82 is applied to prevent aliasing of their energy into lower frequencies. In addition to these large-scale
83 barotropic corrections which are removed from the altimetric data, there exist high-frequency SSH
84 signals from internal tides and internal waves that contribute energy at small scales < 300 km
85 wavelengths. Their impact on SSH wavenumber spectra has been predicted from model analyses in
86 different regions (Richman et al., 2012, Ray and Zaron, 2016), and show that they can dominate in
87 regions of low eddy energy. Dufau et al. (2016) demonstrated that ~~un~~resolved internal tides can
88 introduce spectral peaks in the altimetric wavenumber spectra from 100-300 km wavelength, ~~and~~
89 ~~can exceed the level of the internal dynamics~~—especially at low latitudes (Fig. 1b). Recent results from
90 a high-resolution $1/48^\circ$ model highlight that the tidal and supertidal signals in one region of the
91 equatorial Pacific greatly exceed the internal-subtidal dynamics at scales less than 300 km

92 | [wavelength, and supertidal phenomena are substantial at scales approximately 100 km and smaller](#)
93 (Savage et al. 2017).

94 A more technical contribution that can impact on the lower spectral slopes in the tropics concerns
95 the altimetric data processing, the spectral calculation and spectral slope estimation. Much attention
96 has been devoted to the effects of altimetric noise (Xu and Fu, 2012; Zhou et al., 2015, Biri et al.,
97 2016) which can flatten the spectral slope calculation if the noise is not removed correctly. Different
98 studies also use different tapering windows to reduce leakage of non-periodic signals in limited-
99 length data series, which can also modify the spectral slope. In global studies, a fixed wavelength
100 band from 70-250 km is often used for the spectral slope calculation (Xu and Fu, 2012; Dufau et al.,
101 2016), which is appropriate for estimating the spectral slope of the energy cascade at mid-latitudes,
102 but may not be well-adapted for the tropics where the maximum spectral slope extends at longer
103 wavelengths, due to the larger Rossby radius there (Fig. 1b).

104
105 Thus, the interpretation of altimetric tropical SSH spectra, at spatial scales smaller than 600 km,
106 remains a matter of debate in terms of ocean dynamics. This paper aims at filling this gap by studying
107 the dynamical processes contributing to the small-scale SSH spectra in the tropical Pacific using
108 modeling and observational data. Two different approaches are proposed to better understand the
109 contributions to the observed altimetric flatter spectral slopes. Firstly, we wish to explore the
110 spectral signatures in SSH and EKE of the tropical Pacific mesoscale dynamics (with periods greater
111 than 10 days and wavelengths down to 25 km) and we will concentrate particularly on the tropical
112 “mesoscale” band that varies with latitude. For this, we analyse the global 1/12° DRAKKAR model in
113 the tropical Pacific from 20°S to 20°N, using 5 day outputs covering the period 1987-2001. In
114 comparison to the altimetric analyses of Xu and Fu (2012) or Dufau et al. (2016), this model was
115 specifically chosen to have no high-frequency response to tides, internal waves or rapid tropical
116 waves, and is not limited at low wavelengths by the altimetric instrument noise, but rather by the
117 horizontal grid resolution. We will also use this model to explore the effects of using limited segment
118 lengths or specific windowing when calculating our wavenumber spectra.

119 In the second part of this paper, we will address the impact on SSH and EKE of the high-frequency
120 components using a unique modelling experiment: we will analyze a higher resolution and high-
121 frequency version of the model: a 1/36° regional model of the south west Pacific (Djath et al., 2014)
122 with and without tides. These two regional model runs have exactly the same configuration and high-
123 frequency atmospheric forcing, both versions include the atmospherically forced internal gravity
124 waves in the tropics. Careful filtering of the barotropic and coherent internal tides from the model
125 with tides also allows us to explore the relative impact of the incoherent tide-ocean circulation
126 interactions, and their signature on the alongtrack wavenumber spectra. This two-model
127 configuration allows us to make a brief investigation of the effects of high-frequency dynamics on the
128 wavenumber spectra, and to discuss the modeled spectra in comparison with altimetric wavenumber
129 spectra based on Topex/Poseidon, Jason and Saral/AltiKa altimeter data. These results will help to
130 better understand the physical content of altimetric observation today, as well as to explore the finer
131 scales that would be captured using future measurements of the SWOT satellite (Fu and Ubelmann,
132 2014).

133 In section 2, the different models and data used are presented. In section 3, we discuss processing
134 issues for the spectral calculation, particularly to reduce leakage effects in short tropical segments. In

135 section 4, we discuss the EKE spectral signature of the dynamics over the tropical Pacific as simulated
136 by the 1/12° resolution model. In section 5, results are discussed in term of balanced dynamics and
137 the 1/12° model's SSH spectra are compared to Jason and Saral AltiKa wavenumber spectra. Finally,
138 the contribution of the high-frequency motions to the SSH spectral signature are investigated using
139 the 1/36° regional resolution model with and without tides, to illustrate its close match with
140 altimetric data. Section 6 presents the conclusions of our study.

141

142 **2 Models, and altimetric data**

143 **2.1 Models**

144 To study mesoscale and submesoscales activity from an OGCM, the model has to properly resolve the
145 corresponding dynamical scales (i.e., be eddy-resolving). The effective resolution for numerical
146 models is that 6-8 grid points are needed to properly resolve dynamical features (Soufflet et al.,
147 2016). In mid latitudes numerical convergence requires ~km horizontal resolution, however in the
148 tropics, because of the larger Ld due the weaker Coriolis force, numerical convergence is obtained
149 from 1/12° horizontal resolution, and the increase of resolution to 1/36° only seems to displace the
150 dissipative range of the model toward smaller scale (Marchesiello et al., 2011).

151 In this paper, we first use a global model at 1/12° resolution [from the DRAKKAR consortium based on](#)
152 [the NEMO code \(Madec, 2008; Lecointre et al., 2011\) referenced as G12d5. This model has 46 levels,](#)
153 [and has been integrated from 1989 to 2007 using a 3-hourly ERA-interim reanalysis \(Dee et al., 2011\).](#)
154 [The 3D velocities and the 2D Sea Surface Height \(SSH\) are saved as 5-day means during the period of](#)
155 [integration. The domain considered in this study covers the tropical Pacific between 20°N -20°S. This](#)
156 [simulation has been used to document mesoscale variability in the South West Pacific Solomon Sea](#)
157 [\(Gourdeau et al., 2014; Gourdeau et al., 2017\).](#)

158 ~~In the~~ second part of the paper, we use a regional [DRAKKAR/NEMO model with 1/36° model](#)
159 [resolution and 75 levels, still with surface forcing from the 3h ERA Interim re-analysis. Two](#)
160 [simulations are performed: one without tidal forcing \(R36\) over the 1992-2012 period, and one with](#)
161 [tidal forcing \(R36T\) over the 1992-2009 period \(Tchilibou et al., 2018\). These different model](#)
162 [configurations are particularly important in this area where internal tides are active \(Niwa and](#)
163 [Hibiwa, 2011; Gourdeau, 1998\), and could modify accordingly the energy flux for the meso and](#)
164 [submesoscale bands \(Richman et al., 2012\). Daily mean model outputs are saved as \(R36\(T\)d\), as well](#)
165 [as instantaneous fields saved hourly \(R36\(T\)h\) during a 3 month period from January-March 1998.](#)
166 [We will use these different configurations](#) to investigate the impact of high frequency ageostrophic
167 motions such as baroclinic tides and internal waves.

168 [Further details on these different model configurations are given in Annexe 1.](#)

169 ~~Global Model at 1/12°~~

170 ~~The model used is the ORCA12-L46-MAL95 configuration of the global 1/12° OGCM developed and~~
171 ~~operated in the DRAKKAR consortium (www.ifremer.fr/lpo/drakkar) (Lecointre et al., 2011). The~~
172 ~~numerical code is based on the oceanic component of the NEMO (Nucleus for European Modelling of~~
173 ~~the Ocean) system (Madec, 2008). The model formulation is based on standard primitive equations.~~

Mis en forme : Police :Gras

174 ~~The equations are discretized on the classical isotropic Arakawa C grid using a Mercator projection.~~
175 ~~Geopotential vertical coordinates are used with 46 levels with a 6m resolution in the upper layers~~
176 ~~and up to 250 m in the deepest regions (5750 m). The “partial step” approach is used (Adcroft et al.,~~
177 ~~1997) to allow the bottom cells thickness to be modified to fit the local bathymetry. This approach~~
178 ~~clearly improves the representation of topography effects (Darnier et al. 2006; Penduff et al. 2007).~~
179 ~~The bathymetry was built from the GEBCO1 dataset~~
180 ~~(http://www.gebco.net/data_and_products/gebco) for regions shallower than 200m and from~~
181 ~~ETOPO2 (www.ngdc.noaa.gov/mgg/global/relief/ETOPO2) for regions deeper than 400m (with a~~
182 ~~combination of both datasets in the 200m-400m depth range). Lateral boundary conditions for~~
183 ~~coastal tangential velocity have a strong impact on the stability of boundary currents (Verron and~~
184 ~~Blayo, 1996). Based on sensitivity experiments, a “partial slip” condition is chosen, where the coastal~~
185 ~~vorticity is not set to 0 (“free slip” condition), but is weaker than in the “no slip” condition. The~~
186 ~~atmospheric forcing (both mechanical and thermodynamical) is applied to the model using the CORE~~
187 ~~bulk formulae approach (Large and Yeager, 2004, 2009).~~

188 ~~The simulation started from rest in 1978 with initial conditions for temperature and salinity provided~~
189 ~~by the 1998 World Ocean Atlas (Levitus, 1998). It was spun up for 11 years using the CORE II forcing~~
190 ~~dataset and then integrated from 1989 to 2007 using a 3 hourly ERA interim forcing (Dee et al.,~~
191 ~~2011). The 3D velocities, and the 2D Sea Surface Height (SSH) are saved as 5 day means during the~~
192 ~~period of integration. In the following, it is referenced as G12d5. The domain considered in this study~~
193 ~~covers the tropical Pacific between 20°N–20°S. This simulation has been used to document~~
194 ~~mesoscale variability in the South West Pacific Solomon Sea (Gourdeau et al., 2014; Gourdeau et al.,~~
195 ~~2017). The present study will analyse this simulation over the tropical Pacific.~~

196 ~~Regional Model at 1/36° with and without tides~~

197 ~~As part of the CLIVAR/SPICE program, regional simulations of the Solomon Sea in the South Western~~
198 ~~tropical Pacific have been performed (Ganachaud et al., 2014). The numerical model of the Solomon~~
199 ~~Sea used in this study has a 1/36° horizontal resolution, and 75 vertical levels. It is based on the same~~
200 ~~oceanic component as the NEMO system presented above. This 1/36° resolution model is embedded~~
201 ~~into the global 1/12° ocean model presented above and one way controlled using an open boundary~~
202 ~~strategy (Tréguier et al., 2001). Its horizontal domain is shown on Fig. 3b. The bathymetry of the high-~~
203 ~~resolution Solomon Sea model is based on the GEBCO08 dataset. Atmospheric boundary conditions,~~
204 ~~consisting in surface fluxes of momentum, heat and freshwater, are diagnosed through classical bulk~~
205 ~~formulas (Large and Yeager, 2009). Wind and atmospheric temperature and humidity are provided~~
206 ~~from the 3 hourly ERA Interim reanalysis (Dee et al., 2011). A first version of the regional model with~~
207 ~~45 vertical levels has been initialized with the climatological mass field of the World Ocean Atlas~~
208 ~~(Levitus et al., 1998) and was integrated from 1989 to 2007. More technical details on this~~
209 ~~configuration may be found in Djath et al. (2014). The new version used here is distinct from the~~
210 ~~former version by the number of vertical levels (75 levels in the new version) but above all by its~~
211 ~~ability to take account realistic tidal forcing (Tchilibou et al., 2018). It is of particular importance in~~
212 ~~this area where internal tides are particularly active (Niwa and Hibiwa, 2011; Gourdeau, 1998), and~~
213 ~~could modify accordingly the energy flux for the meso and submesoscale bands (Richman et al.,~~
214 ~~2012). The model is forced at the open boundary by prescribing the first 9 main tidal harmonics (M2,~~
215 ~~S2, N2, K2, K1, O1, P1, Q1, M4) as defined from the global tides atlas FES2014 (Lyard et al., 2019)~~
216 ~~through a forced gravity wave radiation condition. The model is initialized by the outputs from the~~

217 ~~ORCA 1/12° version. Two simulations are performed: one without the tidal forcing (R36) over the~~
218 ~~1992-2012 period, and one with the tidal forcing (R36T) over the 1992-2009 period. Model outputs~~
219 ~~are saved as daily mean (R36(T)d), and instantaneous fields are saved hourly (R36(T)h) during a 3~~
220 ~~month period from January-March 1998.~~

221 2.2 Altimetric data

222 Along-track SSH observations from TOPEX/Poseidon covering a period (January 1993 to December
223 2001) in common with the G12d5 simulations are analyzed over the tropical Pacific domain. The
224 most recent altimetric missions (Jason-2 and SARAL/Altika) are also analyzed over the January-2013
225 to December-2014 period to compare with the signature of the high frequency modelled SSH in
226 R36Th. These data are made available from the Copernicus Marine and Environment Monitoring
227 Service (CMEMS, <http://marine.copernicus.eu>). TOPEX/Poseidon and Jason-2 are conventional pulse-
228 width limited altimeters operating in the Ku-band (Lambin et al., 2010). SARAL/Altika with its 40 Hz
229 Ka-band emitting frequency, its wider bandwidth, lower orbit, increased Pulse Repetitivity Frequency
230 and reduced antenna beamwidth, provides a smaller footprint and lower noise than the Ku-band
231 altimeters (Verron et al., 2015). For the different missions we will analyze the 1 Hz data, extracted
232 over the same region as our model analysis.

233

234 3 Spectral methods

235 In the following sections we present spectral analyses of the modelled SSH or EKE fields, or the
236 altimetric SSH. The spectral analysis we use is based on Fast Fourier Transforms (FFT) of our signal,
237 which allows us to work with a limited sampled signal. Longer data records enable a better
238 decomposition of the variability at each frequency (wavenumber) and thus a better separation of
239 neighboring frequencies in the spectrum. However, for wavenumber spectra, long spatial data
240 records can mix information from different geographical regimes, especially in the tropics where
241 meridional sections cross the strong zonal currents, making their dynamical interpretation difficult.

242

243 Different studies performing spectral analysis of altimetric data or models over the global ocean use
244 very different data length segments to calculate the spectrum. Some altimetric studies use data
245 segment lengths of around 500 km (e.g. Dufau et al., 2016), or 1000 km length tracks averaged in 10°
246 or 20° square box, with or without overlapping (Xu and Fu, 2012). Model spectra (~~isotropic~~) are
247 mostly calculated in 10° or 20° square boxes (e.g., Sasaki and Klein, 2012; Biri et al., 2016; Chassignet
248 and Xu, 2017). These data segment lengths may be adequate for the mid-latitudes but are not
249 appropriate for the tropics, when the maximum energy can occur at 600-1000 km wavelengths.
250 Using shorter segments than this reduces the maximum energy and should increase the leakage from
251 energetic low wavenumbers to weaker high wavenumbers, thus decreasing the spectral slope
252 (Bendat and Piersol, 2000).

253

254 A wide variety of filter windows are applied in the different studies before calculating frequency
255 (wavenumber) spectra to reduce the leakage effect. These include the 10 % cosine taper window or
256 Tukey 0.1 window, referred hereafter as Tk01 (LeTraon et al., 2008; Richman et al., 2012; Dufau et
257 al., 2016); the Hanning window, referred as Hann (Capet et al., 2008; Rocha et al., 2016); or making
258 the signal double periodic instead of the tapering, referred as Dbp (Marchesiello et al. 2011; Sasaki

259 and Klein, 2012; Chassignet and Xu, 2017). In the following, we will also consider ~~the a 50% cosine~~
260 ~~taper Tukey-0.5~~ window (Tk05).

261
262 We tested the sensitivity of our G12d5 model's SSH wavenumber spectrum to the different tapering
263 windows and the double periodic method, using different data length sizes, and in one or two
264 dimensions. ~~The details are given in Annexe 2. The following steps were performed for these test~~
265 ~~spectra, evaluated within 10°S-10°N/160°W-120°W: the model data are extracted meridionally and~~
266 ~~zonally in fixed segment lengths of 5°, 10°, 20° and within a 20°X20° square box; the mean and linear~~
267 ~~trend (fitted plane for two-dimensional case) were removed from each data segment or box; the~~
268 ~~filter window (Tk01, Tk05, Hann) or Dbp are applied; temporal and spatial (longitude, latitude) series~~
269 ~~spectra are calculated and averaged in Fourier space. The results are shown in Fig. 2.~~

270 Tk01 meridional spectra in the tropics are the most perturbed by the short segment lengths (Fig. 2a).
271 In the 70-250 km range commonly used to define a global mesoscale band (delimited by the green
272 vertical lines), the spectral slope flattens as the data segment length decreases. 5° segment spectra
273 with a Tk01 window have a $k^{-1.3}$ slope, which explains the very shallow slope in the tropics observed
274 by Dufau et al. 2016 who applied this short data segment size and a Tk01 window. Meridional
275 spectra differ primarily at larger scales from 100-500 km, when short segment lengths are used (Fig.
276 2a). A comparison of the meridional spectrum using 20° segments and different windows (Tk01,
277 Tk05, Hann and Dbp) are shown in Figure 2b. Even with the 20° segments, Tk01 is distorted. On the
278 other hand, the Tk05, Hann and Dbp match well, with a near linear cascade of energy over the 30-
279 1000 km wavelength range, and are more adapted for the tropics since they capture the main range
280 of SSH mesoscale dynamics, particularly the spectral energy peaks around 1000 km wavelength.

281 Similar calculations were performed for the zonal spectra (not shown) and confirm that the Tk01
282 method deforms the zonal spectra and flattens the spectral slope within the 70-250 km wavelength
283 band as the data segment size decreases. Tk05, Hann and Dbp 20° segment spectra match, although
284 the Dbp has more noise at small scale.

285 We also conducted a sensitivity test in the off-equatorial region (not shown): Flattening and
286 deformation of the spectrum by Tk01 persist, but the 10° segments or 10° square box are long
287 enough to capture the off-equatorial dynamics.

288 ~~The particular sensitivity of spectra in the tropics to the spectral segment length and windowing is~~
289 ~~linked to energetic EKE and SSH signals extending out to longer wavelengths, and illustrates the~~
290 ~~ability to deal with spectral leakage from low to high wavenumbers. Tk01 is the worst tapering~~
291 ~~window, and the distortion of spectra. The discrepancies between the different window filters used~~
292 ~~result from the way they modify the original data. Tk01 preserves more of the long wavelength signal~~
293 ~~in the original data. However it creates unrealistic variability at shorter wavelengths because of its~~
294 ~~abrupt transition to zero at the edge of segments. This effect is amplified for short data segments,~~
295 ~~resulting in more important variability at small scale (for wavelengths near 10% of the data size).~~
296 ~~Thus there is more spurious energy near 55 km wavelength for the Tk01 with 5 degree (550 km)~~
297 ~~segments, near 110 km for the 10° segments, and near 220 km for the 20° segments (Fig. 2a). The~~
298 ~~Tk01 abrupt transition to zero combines with the low wavenumber leakage and distorts the spectra.~~
299 ~~There is a minor difference between Tk05 and Han, depending on data variability. The advantage of~~
300 ~~Tk05 is in retrieving the large-scale peaks which are smoothed with the Hanning filter window. Tk05,~~

301 ~~and Hann are~~ a good compromise for preserving much of the original signal and reducing leakage,
302 ~~but needs to be applied over larger segments. The Dbp technique compares well to Tk05 and Han for~~
303 ~~one-dimensional spectra. The only problem with this method is the difference at the large-scale and~~
304 ~~unfiltered spectral fluctuations at small scale.~~

305 ~~So, we~~ find that to safely avoid leakage in the tropics, it is best to use a long record and an effective
306 ~~taper window. We do not advise to use the Tk01 filter window. We conclude that the~~ Tk05 or
307 Hanning filters give ~~similar-convincing~~ results in the equatorial band, with a minimum of 15° to 20°
308 needed in segment lengths (Fig. A1). ~~We do not advise to use the Tk01 filter window. In the off-~~
309 ~~equator region, 10° data segments or 10°X10° boxes are sufficient.- We do not advise to use the tk01~~
310 ~~filter window, and we~~ We choose to use the Tukey 0.5 filter for our tropical spectral analyses in this
311 paper. ~~In the off-equator region, 10° data segments or 10°X10° boxes are sufficient.~~
312

313 4 Spectral representation of the tropical dynamics

314 In this section we analyze the spectral signatures of the tropical dynamics by first considering the
315 surface velocity fields of the G12d5 simulation over the open Pacific Ocean. Modeling studies mainly
316 analyze velocity or EKE fields, and we start our spectral analysis by checking that the model
317 represents well the main dynamical processes in the tropics. Surface velocity fields were averaged
318 over the first 40 m depth and include geostrophic and ageostrophic components. The model resolves
319 a domain of variability with periods greater than 10 days, and wavelengths exceeding 25 km, but
320 model dissipation may be active up to 70 km wavelength. Note that the resonant response to the
321 wind forcing through the 3-5 day period, large-scale equatorially trapped inertia-gravity waves, are
322 not represented in G12d5 because of the 5-day averaged model outputs.

323 The Tropical Pacific is characterized by a series of strong alternate zonal currents and a large range of
324 ocean variability, in response to the atmospheric forcing and to the intrinsic instability of the current
325 system. The main zonal currents spanning the tropical Pacific are shown in Fig. 23: North of 10°N is
326 the westward North Equatorial Current (NEC) and at its northern edge are the eastward SubTropical
327 CounterCurrent (STCC) and the Hawaiian Lee CounterCurrent (HLCC) (Kobashi and Kawamura, 2002;
328 Sasaki and Nonaka, 2006); between 3°-8°N is the eastward North Equatorial CounterCurrent (NECC);
329 South of 3°N, the westward South Equatorial Current (SEC) straddling the equator is divided in two
330 branches by the eastward Equatorial UnderCurrent (EUC) that reaches the surface to the east. The
331 eastward South Equatorial Counter Current (SECC) in the south western Pacific is between 6°-11°S.
332 Instabilities of these zonal currents result in meso and submesoscale activity illustrated by a snapshot
333 of vorticity (Fig. 32) that illustrates the description of vortices in Ubelmann and Fu (2011). It is
334 characterized by structures with a large range of scale and strong anisotropy in the equatorial band.
335 The largest structures (~500 km) correspond to the nonlinear Tropical Instability Vortices (TIVs), also
336 associated with the Tropical Instability Waves (TIWs), and occur north of the equator (Kennan and
337 Flament, 2000; Lyman et al., 2007). The off-equatorial regions (10-20° latitude) are characterized by
338 smaller-scale turbulent structures in Fig. 32.

339 In order to investigate how these well-known tropical dynamics project into frequency or
340 wavenumber spectra, we will analyze separately the equatorial band (10S-10°N) and the off-
341 equatorial band (10°N-20°N and 10°S-20°S) defined by the different boxes in Fig. 32. In the following,
342 the model's representation of the following diagnostics will be discussed together for the each zonal

343 | band : the EKE frequency spectra as a function of latitude and longitude (Fig. 43), the zonal EKE
344 | wavenumber-frequency (k - ω) spectra and meridional EKE wavenumber-frequency (l - ω) spectra (Fig.
345 | 54), and the 1D (zonal/meridional) EKE wavenumber spectra (Fig. 65).

346

347 | 4.1 Equatorial region

348 | The temporal variability of the tropical EKE signal is shown by EKE frequency spectra as a function of
349 | latitude and longitude in Fig. 34. In the equatorial band, most of the energy is concentrated within 5°
350 | of the equator (Fig. 34a). The highest EKE occurs in this band at annual to interannual scales, but
351 | there is still significant energy over all frequencies greater than the 10-days resolved by this model.
352 | EKE spectra averaged in latitude over 20°N-20°S are highly influenced by the energetic equatorial
353 | dynamics (Fig. 34b). This band includes the equatorial wave guide where waves tend to propagate
354 | zonally and are organized into a set of discrete meridional modes (Farrar, 2008). Since zonal
355 | wavenumber-frequency spectra are averaged from a number of latitudes within the equatorial band,
356 | contributions from the different modes may be seen at once (Fig. 45b). The eastward phase speed
357 | (positive wavenumber), due to fast moving Kelvin waves at the equator is visible even if the strong
358 | westward propagation (negative wavenumber) just off the equator overpowers the eastward
359 | propagation on the equator in the averaged spectrum. We have superimposed on the zonal
360 | wavenumber-frequency spectrum the theoretical dispersion curves of the first baroclinic-Rossby
361 | waves in a resting ocean. Values of wavenumber and frequency for which the EKE power spectrum is
362 | significantly above the background follow relatively well the variance-weighted mean location of
363 | dispersion curves for long equatorial waves. Meridional wavenumber-frequency (l - ω) EKE spectra
364 | were computed over the 20°N to 20°S section, in different longitude bands spanning the Pacific
365 | Ocean. Fig. 45d shows an example for the particularly energetic 120°W-150°W band. Other longitude
366 | bands across the Pacific show similar spectral energy patterns, but with lower energy levels. Figures
367 | 45b,d illustrate the strong anisotropy between the zonal (k , ω) and meridional (l , ω) spectra. The
368 | meridional structure of the dominant zonal equatorial waves is well known, with meridional
369 | amplitude decaying away from the equator over +/-5° or 550 km. This contributes in the meridional-
370 | frequency EKE spectrum to the fairly constant decrease in spectral energy from long wavelengths
371 | down to 100-250 km wavelength, in both north and south directions (Fig. 45d).

372

373 | The ridge of westward variance (Fig. 45b) is nearly vertical, with variance mainly restricted to large
374 | wavelengths but also extending to high frequencies in relation with TIW activity. In accordance with
375 | observations (Willet, 2006; Lee et al. 2018), the modeled TIWs are defined by periods and zonal
376 | wavelengths in the range of 15-40 days and 800-2000 km, respectively. ~~They, and~~ have a meridional
377 | propagation with northward and southward motion roughly balanced ~~that is a hallmark of standing~~
378 | ~~meridional modes for TIWs as seen from others perspectives in Lyman et al. (2005) and Farrar (2008,~~
379 | ~~2011) and earlier work~~ (Fig. 45d). The 33-day TIW variability is triggered by baroclinic instability of
380 | the SEC-NECC system, located between 3°N-5°N and 160°W-120°W (Fig. 34a,b). They have an
381 | asymmetric structure across the equator with larger energy north of the equator than south of it in
382 | accordance with the analysis of TOPEX/Poseidon sea level data by Farrar (2008). The 20-25 days
383 | variability, associated with another type of TIW triggers by barotropic instability of the EUC-SEC
384 | system (Masina et al., 1999), is centered at the equator, east of 140°W (Fig. 34a,b). Centered at the
385 | equator, from the background there is a 60-80 days variability extending from 150°E to 130°W (Fig.
386 | 34a,b) associated with intraseasonal Kelvin waves (Cravatte et al., 2003; Kessler et al., 1995) as

387 confirmed by eastward variance and energy centered at $l=0$ in the zonal and meridional-frequency
388 spectra, respectively (Fig. 45b,d).

389
390 The model represents these tropical signals well, and for wavelengths larger than 600 km the
391 equatorial waves are the dominant signal (Tulloch et al., 2009). For wavelengths smaller than 600
392 km, the variance no longer follows the Rossby wave dispersion curves, and exhibits a red noise
393 character in wavelength, and a nearly white noise in frequency. These rapid motions with 250-600
394 km wavelengths occur in response to wind forcing, wave interactions or current instability. The
395 corresponding zonal EKE wavenumber spectrum (Fig. 56) has a steep slope that continues rising to
396 long wavelengths with a k^3 relation reaching a peak at 1000 km, reflecting the zonal scales of the
397 TIWs, before flattening to a k^{-1} power law at larger scale. Below 70 km, EKE spectra drastically
398 steepen as an effect of model dissipation.

399 4.2 Off equatorial regions

400 Poleward of 10° the equatorial trapped waves become insignificant, and most of the energy is
401 concentrated at periods greater than 60 days (Fig. 34a). This corresponds with results by Fu (2004)
402 showing a decreasing frequency range with latitude, where the maximum frequency at each latitude
403 corresponds to the critical frequency of the first-mode baroclinic waves that varies from 60 days at
404 10°S to 110 days at 20°S (Lin et al., 2008). The zonal wavenumber-frequency spectrum strongly
405 differs from those in the equatorial belt (Fig. 45a,c), and is closer to the mid latitude spectra
406 (Wunsch, 2010; Wakata, 2007; Fu, 2004) with smaller energy in the south tropics than in the north as
407 also reported by Fu (2004). The theoretical dispersion curves for mid latitude first baroclinic Rossby
408 waves are shown for the case of meridional wavenumbers corresponding to infinite wavelengths. At
409 low wavenumbers (i.e., long wavelengths > 600 km) the motions follow the baroclinic dispersion
410 curves.

411 Although linear Rossby wave theory provides a first - order description of the EKE spectra, in both
412 hemispheres energy extends to higher frequencies (Fig. 34a), and as the wavenumber and frequency
413 increases, significant deviations from the baroclinic dispersion curves occur (Fig. 45a,c). Much of the
414 energy lies approximately along a straight line called the 'non dispersive line' in wavenumber-
415 frequency space as it implies non-dispersive motions. Barotropic and baroclinic instabilities, and
416 nonlinear processes might generate energy at periods down to 20-30 days, prohibited by the linear
417 dispersion relation (Hughes and Williams, 2010). Much of the energy lies approximately along a
418 straight line called the 'non dispersive line' in wavenumber frequency space as it implies non-
419 dispersive motions. The wavenumber dependencies along the 'non dispersive line' could be the
420 signature of non-linear eddies. In fact, for eddies that retain their shapes as they propagate, the
421 energy at every wavenumber must propagate non-dispersively (Rhines, 1975). The westward
422 propagation speed is estimated at ≈ 10 cm/s, close to the eddy propagation speed found in this
423 latitudinal range by Fu (2009) and Chelton et al. (2007). But these regions are defined as a weakly
424 nonlinear regime (Klocker and Abernathey, 2014). I, and n this region of mean zonal currents the
425 dispersion curves experience Doppler shifting by the zonal flow which makes the variability nearly
426 non dispersive (Farrar and Weller, 2006). So, the non-dispersive line may could account both for
427 coherent vortices and and more linear dynamics as Rossby waves or jet-meandering jets
428 propagating westward at an approximately uniform speed (Mortén et al., 2017). So our frequency-

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429 ~~wavenumber spectra are consistent both with linear dynamics at low wavenumbers but also with~~
430 ~~westward-propagating eddy-like coherent vortex structures or jet meanders at higher wavenumbers.~~

431 The zonal EKE wavenumber spectra (Fig. 56) in the off equatorial regions exhibit a standard shape
432 with a long-wavelength plateau and a spectral break at about 300-400 km, following by a drop in
433 energy close to a k^{-2}/k^{-3} relation (Stammer, 1997). These steep spectral slopes correspond with an
434 inertial range characteristic of mesoscale turbulence (Xu and Fu, 2011). These different spectra
435 confirm that the northern tropics are more energetic than the southern part with a mesoscale range
436 extending to larger scale. It quantifies the more active turbulence in the northern hemisphere, as
437 illustrated in Fig. 23b.

438 4.3 Anisotropic EKE spectra

439
440 Classically, wavenumber spectra are investigated throughout an oceanic basin by dividing the basin
441 in square boxes where spectra are calculated to take account of the regional diversity of QG
442 turbulence properties (Xu and Fu, 2011; Sasaki and Klein, 2012; Biri et al., 2016; Dufau et al., 2016).
443 Here, the spectra analysis of the equatorial and off-equatorial bands described above is revisited in
444 $10^\circ \times 10^\circ$ boxes for the off-equatorial region, and in $20^\circ \times 20^\circ$ boxes for the equatorial region that are
445 suited to recover the shape of the mesoscale range in the tropics (e.g. section 3). Within each
446 equatorial or off equatorial latitude band, spectra in the different boxes are similar (not shown).
447 Therefore spectra are averaged over all the boxes and we present one mean spectrum
448 representative of the square boxes for each equatorial, and off equatorial bands. In geostrophic
449 turbulence, which is nondivergent to leading-order, isotropy implies that 1D (zonal/meridional) and
450 2D ~~isotropic-azimuthally integrated~~ wavenumber spectra ~~azimuthally integrated (or wavenumber~~
451 ~~magnitude spectra)~~, are identical and follow the same power law. In the tropics there is a stronger
452 anisotropic component of the dynamics, which will be explored in Fig. 76.

453
454 When we concentrate on the $20^\circ \times 20^\circ$ equatorial box, we are limited to ~~scales-wavelengths~~ smaller
455 than 2000 km, and the meridional EKE spectrum has a higher level of energy than the zonal one (Fig.
456 76b). ~~This reflects a shift of energy towards the smaller scales in the meridional direction that it~~
457 ~~reveals the energetic meridional perturbations due to instabilities of the larger-scale zonal currents.~~
458 ~~is consistent with the widely held notion that scales of variability near the equator tend to be larger~~
459 ~~in the zonal direction than in the meridional direction for many kinds of variability (mean currents,~~
460 ~~inertia-gravity waves, Kelvin waves, Yanai waves, TIWs).~~ The ~~isotropic-magnitude~~ EKE spectrum is
461 mostly representative of the meridional one. Note that since alongtrack altimetry is mainly
462 orientated in the meridional direction in the tropics, altimetric SSH measurements are particularly
463 well suited to account for the dominant meridional variability, within the limit of the geostrophic
464 hypothesis. Despite the anisotropy at every scale, the different EKE spectral components have a
465 similar shape, with a continuous k^{-3} slope between 100 and 600 km wavelength. The peak of the EKE
466 spectra corresponds to a wavelength of 1000 km. These modeling results compare relatively well
467 with the analysis of the submesoscale dynamics associated with the TIWs by Marchesiello et al.
468 (2011). They observe a peak of energy around 1000 km corresponding to the TIW wavelength, and a
469 linearly decay of the spectrum with a slope shallower than -3. It is doubtful to define an inertial band
470 in the equatorial region, but we can say that at wavelengths from 100-600 km, the EKE spectral slope
471 of k^{-3} is consistent with a QG cascade of turbulence.

Commentaire [RM1]: There is no panel b in this Figure ?

473 In the $10^\circ \times 10^\circ$ off-equatorial boxes, the energy at long wavelengths is greatly reduced compared to
474 the equatorial band. The peak of the EKE spectra corresponds to a wavelength of 300 km. Yet the
475 zonal, meridional and ~~isotropic-magnitude~~ EKE spectra are similar for wavelengths up to 250 km (Fig.
476 ~~7a6a,c~~). ~~Their slopes are steeper than k^{-2} at scales smaller than 250 km.~~ So, poleward of 10° the
477 hypothesis of isotropy seems to be relevant for scales up to 250 km even if the flow is supposed to
478 be weakly nonlinear, and sensitive to beta effect (Klocker and Abernathey, 2014). ~~The peak of the~~
479 ~~EKE spectra corresponds to a wavelength of 300 km.~~ The EKE slope over the redefined mesoscale
480 range from 100 to 250 km is between -2 and -3 which lies between the prediction of SQG and QG
481 turbulence.

482
483 Our modeled zonal frequency-wavenumber spectra differ strongly across the equatorial and off
484 equatorial regions. ~~They, but~~ show a good representation of the tropical wave and TIW/TIV
485 dynamics. The slope of the ridge of westward variance in the zonal k - ω spectrum in Fig. ~~5-4~~ increases
486 towards the equator. As the slope becomes steeper, more power is concentrated at lower
487 wavenumbers. The change in slope of the ridge itself is mainly related to the change in deformation
488 radius, and expresses linear or non-linear variability propagating non dispersively ~~and wave~~
489 ~~propagation speed~~ (Wortham and Wunsch, 2014). The equatorial region differs from the off
490 equatorial regions in having strong anisotropy with mainly zonally oriented structures (Fig. ~~76~~),
491 higher energy at long wavelength due to the strong activity of long equatorial waves, and an overlap
492 between geostrophic turbulence and Rossby wave time scales that produces long waves and slows
493 down the energy cascade to eddies with scales consistent in the tropics with a generalized Rhines
494 scale (L_r) (Theiss, 2004, Tulloch et al., 2009; Klocker et al., 2016; Eden, 2007).

495 ~~Although geostrophy is not valid at the equator, it is valid poleward from 2° latitude, and since our~~
496 ~~equatorial wavenumber spectral calculations are averaged or span over 20° in latitude, they include a~~
497 ~~strong geostrophic component. Even so, the validity of geostrophic turbulence theory is questionable~~
498 ~~in the equatorial band. Rather~~ Moreover, our modeled spectral analysis shows the contrasts between
499 the equatorial and off-equatorial regions for the wavenumber range where a steep slope is observed.
500 In the weakly nonlinear regime of the off-equatorial regions, we find spectral slopes of k^{-2} / k^{-3} over a
501 short 100-250 km wavenumber range. The ~~highly anisotropic, ageostrophic~~ equatorial dynamics are
502 characterized by a peak of energy at 1000 km due to TIWs, and a large "mesoscale" range over 100-
503 600 km wavelength with a k^{-3} spectral slope.

504

505 **5 Modeled and altimetric SSH wavenumber spectra**

506 **5.1 Contribution from low-frequency dynamics**

507 ~~The SSH is a measure of the surface pressure field, an important dynamical variable, which may be~~
508 ~~balanced in the tropics by both geostrophic and ageostrophic motions. The ocean circulation is~~
509 ~~classically inferred from altimetric SSH through the geostrophic equilibrium. Here, Altimetric SSH only~~
510 ~~provides access to the balanced, geostrophic component of the surface flow (in addition to signals~~
511 ~~from the high-frequency internal waves and tides, not included in the G12d5 version of the model~~
512 ~~used here). Since the tropical regions have a strong ageostrophic component of the flow,~~ we consider
513 how the wavenumber spectra of geostrophic currents (EKEg) differ from that of the total currents
514 analyzed in section 4. Close to the equator, as f approaches zero, the geostrophic current component

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515 | can still be calculated using the beta approximation, following Picaut et al. (1989). Fig. 7-6 shows the
516 | difference between the wavenumber spectra calculated from the total EKE averaged over the upper
517 | 40 m, and from the geostrophic component of EKE estimated at the surface.

518 | In the equatorial band at scales from 300 to 1000 km, the ageostrophic EKE is more energetic, with a
519 | stronger contribution to the total EKE than the geostrophic component (Fig. 6-7b). In the off-
520 | equatorial bands (Fig. 7a6a,c), the geostrophic and total EKE spectra are similar at larger wavelengths.
521 | However, in all regions, the total EKE is steeper than the geostrophic EKE at scales from 250 km down
522 | to the 20 km resolved by the model. In mid latitude regions Klein-Ponte et al. (201308) also noted
523 | stronger geostrophic EKE at small wavelengths (and weaker spectral slopes) compared to upper
524 | ocean EKE spectra in mid latitude regions which had QG slope characteristics associated with wind-
525 | driven mixed layer dynamics. In terms of spectral slope in the equatorial region, using the
526 | geostrophic EKE rather than the total EKE tends to flatten the spectra in the 600-110 km mesoscale
527 | range, and changes the spectral slope from k^3 to k^2 . In the off-equatorial regions, the geostrophic
528 | EKE has a slightly flatter spectral slope between -2 and -3 in the 100-250 km band.

529 |
530 | Since the altimetric groundtracks have a more meridional orientation in the tropics, the altimetric
531 | SSH spectra should be like the model's meridional SSH spectra that are shown on Figure 7-8. SSH
532 | meridional wavenumber spectra (Fig. 7-8) confirm that in the off-equatorial regions, the northern
533 | zone has higher spectral power over all wavelengths, as expected from the EKE spectra. Within the
534 | wavelength band from 100 to 250 km both off-equatorial regions have SSH spectral slopes between
535 | k^4 to k^5 (equivalent to k^2 - k^3 in EKE) similar to QG dynamics. The modelled SSH spectra show a
536 | similar anisotropy in the equatorial zone as the EKE spectra, with a more energetic meridional SSH
537 | spectrum than the zonal spectrum (not shown). It is notable that although the level of energy is
538 | higher in the equatorial region than in the off-equatorial regions, the SSH variability is lower for
539 | wavelengths smaller than 500 km. This reduced SSH variability of the G12d5 model is not in
540 | agreement with the higher small "scale" SSH levels altimetry to be discussed in the next section
541 | (section 5.2) Due to the strong ageostrophic component in the equatorial region, SSH spectra exhibit
542 | lower spectral power than in the off-equatorial region. From 100 to 600 km, the SSH spectral slopes
543 | in the equatorial region are close to k^4 , consistent with the k^2 spectral slopes in EKE. The fixed
544 | wavelength band used by previous studies [70-250 km] can be compared to this longer wavelength
545 | band. Using the fixed wavelength band leads to a slight reduction in the low-frequency SSH spectral
546 | slope estimate, but without a drastic modification. These results indicate that if the internal balanced
547 | dynamics of our $1/12^\circ$ model were the main contribution to the altimetric SSH, then we would
548 | expect a k^4 (sQG) slope in the equatorial band, and closer to k^5 (QG) in the off-equatorial band.

549 | Fig. 7-8 also shows the alongtrack Topex/Poseidon SSH spectra over the same region and period as
550 | the G12d5 simulation. The altimetric data are selected with the same segment lengths, and with the
551 | same pre-processing and spectral filtering as in the model. In the equatorial and off-equatorial zones,
552 | the altimetric SSH wavenumber spectra clearly exhibit the weaker k^2/k^1 spectral slopes in the 70-250
553 | km mesoscale range as described in previous studies (Xu and Fu, 2011,2012; Zhou et al., 2015). At
554 | scales larger than our spectral slope range (600 km in the equatorial region, 200 km in the off-
555 | equatorial zones), the model-altimeter spectra have similar shapes although the altimeter data has
556 | higher spectral power. Potentially, the high-frequency < 10d rapid equatorial waves with longer
557 | wavelengths are not included in the model, and may contribute to these differences. The spectral
558 | peaks in the altimetric data at 120-150 km wavelength are indicative of internal tides, as noted by

559 Dufau et al. (2016); Savage et al., (2017), and others. In addition to the internal tide peaks, the
560 general higher spectral energy in the altimetry data at wavelengths < 200 km has been proposed to
561 be due to high-frequency internal gravity waves (eg Richman et al., 2012, Savage et al., 2017), but
562 may also include altimetric errors from surface waves and instrument noise (Dibarboure et al., 2014).
563 We will investigate the high-frequency contribution to the altimetric SSH spectra in the next section.

564 **5.2 Contributions from high-frequency dynamics including internal tides**

565 To investigate the contribution of the high-frequency SSH variations, we include an analysis of the
566 meridional SSH spectra from a small region east of the Solomon Sea in the SW Pacific. This spectral
567 analysis is derived from the 1/36° model with high-frequency atmospheric forcing and instantaneous
568 snapshots saved once per hour during a 3 month period, and run in the two configurations, with and
569 without tides (see section 2). The model has been validated and analysed (Djath et al., 2014), and a
570 companion paper will address the model with tides more fully (Tchilibou et al., 2018). Here we
571 consider specifically the impact of the different high-frequency tides and non-tidal signals on the
572 meridional SSH spectra.

573 The internal tide can be broken down into a coherent component that is predictable and can be
574 separated with harmonic and modal analysis, and an incoherent component that varies over time,
575 due to changing stratification (Zaron, 2017) or interaction with the mesoscale ocean circulation
576 (Ponte and Klein, 2015). The coherent baroclinic (internal) tide and the barotropic tide are calculated
577 in our study using a harmonic and modal decomposition (Nugroho, 2017) which separates the
578 barotropic mode and 9 internal tide modes, and provides a more stable energy repartition between
579 the baroclinic and barotropic components (F. Lyard, Personal Communication). Previous studies have
580 addressed the internal tide and high-frequency components in the tropics by careful filtering of a
581 model with tides (e.g. Richman et al., 2012; Savage et al., 2017). Aside from the issues of artifacts
582 introduced by the tidal filtering, it is often tricky to cleanly separate the spectral contributions
583 coming from the mesoscale ocean circulation and the incoherent component of the internal tides.
584 The advantage of using our two-model configuration is that we can specifically calculate the high-
585 frequency non-tidal components of the SSH spectra from the first model, and the component due to
586 the interaction of the internal tide and the model's eddy-current turbulence with the second model.

587 | Fig. 89 shows the geographical distribution of the standard deviation of SSH for the model including
588 the tidal forcing, for the low frequency (> 48 hr) component of the ocean (mesoscale) dynamics and
589 for the high frequency component (< 48 hr) due mainly to internal waves and internal tides. The large
590 | mesoscale variability (up to 6 cm) east of the Solomon Sea in Fig. 89a is similar to the model without
591 | tides (not shown), and well documented as current instability from the SECC-SEC current system (Qiu
592 | and Chen, 2004). It is notable that the high frequency variability from the model with tides in Fig. 89b
593 | is as high as the mesoscale variability, especially in the Solomon Sea, and comes mainly from the M2
594 | baroclinic tide. We note that the M2 barotropic tide amplitude within the Solomon Sea is relatively
595 | weak (not shown), and the largest internal tide amplitudes are close to their generation sites,
596 | particularly where the barotropic tide interacts with the northern and southern Solomon Islands and
597 | the southeastern Papua New Guinea (PNG) extremities (Tchilibou et al., 2018). For the model
598 | without tides, the high frequency variability due to the atmospherically forced internal gravity waves
599 | is very low (~1 cm) compared to the model with tides, and shows a relatively uniform distribution
600 | (not shown).

601 | The region used for our spectral analysis [2-13°S; 163-165°E; Fig. 89b] is outside the Solomon Sea,
602 | with its strong regional circulation delimited by the islands and bathymetric gradients, and is more
603 | representative of the open Pacific Ocean conditions analysed in the previous sections. The latitude
604 | band from 2°S-13°S lies mostly the equatorial band defined in our previous analyses, and it is mainly
605 | representative of the SECC region (Fig. 23).

606 | The meridional SSH spectra from the 1/36° model run with no tides (R36h) with hourly outputs is
607 | shown in Fig. 10-9 (in green). The SSH from this version with no tides but averaged over 5 days is also
608 | shown (in orange), i.e., with equivalent temporal sampling to our 1/12° model analysis. The
609 | difference between these curves represents the non-tidal high-frequency component of the
610 | circulation (< 10 days) due to rapid tropical waves and internal gravity waves forced by the
611 | atmospheric forcing and current-bathymetric interactions. Also shown is the spectrum calculated at
612 | the same location from our open-ocean G12d5 1/12° model (in cyan) with similar spectral slope to
613 | the 5-day averaged version of our regional R36h 1/36° model, though with slightly lower energy at
614 | scales less than 70 km wavelength as expected, but also in the 180 to 600 km wavelength band. So
615 | the 1/36° model with no tides, when filtered to remove the high-frequency forcing, is quite close to
616 | the 1/12° model in this equatorial band. The main point is that the additional high-frequency
617 | dynamics in R36h increase the spectral SSH power from 300 km down to the smallest scales from 0.4
618 | cm² to 0.5 cm², and reduce the spectral slope calculated in the fixed 70-250 km range from k⁻⁵ with
619 | the 5-day average (in orange), to k⁻⁴ for the full model with no tides (in green).

620 | The 1/36° model with tides (R36Th) is also shown in blue, but with the barotropic tide removed. The
621 | additional meridional SSH spectral power is due both to the coherent and incoherent internal tides,
622 | with a large increase in variance up to 300 km wavelength from 0.5 cm² for R36h to 2.8 cm² for
623 | R36Th. So, the main contributors to the high wavenumber SSH spectral power are from the baroclinic
624 | tides compared to atmospherically-forced high frequency dynamics (green curve). To illustrate the
625 | respective part of coherent and incoherent baroclinic tides, the coherent baroclinic tide signature
626 | based on the nine tidal constituents summed over the first 9 internal modes is calculated, and this
627 | signal is added to the model without tides (purple curve). The coherent baroclinic tides explain most
628 | of the tidal signature in the 300-30 km wavelength range, and the difference with the raw signal
629 | (blue curve) exhibits the signature of incoherent tides. The contribution of the incoherent
630 | component increases significantly at scales smaller than 30 km and explain 30 % of the SSH variance.
631 | The most energetic coherent internal tide component comes from the M2 tide, and the large
632 | increase in amplitude centered around 120-140 km wavelength corresponds with the first baroclinic
633 | mode (not shown). The other peaks around 70 km, and 40 km could be due to higher modes, and
634 | similar peaks are found in the tidal analysis of MITGCM model data by Savage et al (2017) in the
635 | central equatorial Pacific. At the main M2 internal tide wavelengths, the incoherent internal tide has
636 | 1.6 times the SSH energy of the coherent tide, indicating that even at the main internal tide
637 | wavelengths, the incoherent internal tide is energetic.

638 | We note that at wavelengths from 70-250 km used in the global altimetry spectral analysis, this 1/36°
639 | model with the full tidal and high-frequency forcing has a flat spectral ~~slope~~peak of around k^{-1.5}, quite
640 | similar to the analysis of alongtrack spectral from Jason-2 (in dashed black) and Saral (in solid black),
641 | in the same region but over the longer 2013-2014 period. We note that the barotropic tide has also
642 | been removed from the altimetric data, using the same global tide atlas applied at the open
643 | boundary conditions for our regional model (FES2014, Lyard et al., 2018). If we use the “mesoscale”

644 range defined for the global model analysis in the equatorial band over 100-600 km wavelength, we
645 still have a weak spectral slope of k^{-2} for both the model with tides and altimetry. Jason-2 has a
646 higher noise level than Saral at scales less than 30 km wavelength (Dufau et al., 2016); the small
647 differences in spectral energy between Jason-2 and Saral over wavelengths from 150 to 450 km may
648 be influenced by the different repetitive cycles of the very few tracks available (1 track for Jason-2
649 and 3 tracks for SARAL/AltiKA) between both missions, and their slightly different track positions.

650 This regional analysis provides a number of key results. The high-frequency, high-resolution regional
651 model confirms our open ocean $1/12^\circ$ analysis. The dynamics at scales > 10 days, with no tidal
652 forcing, give rise to SSH spectral slopes from 70-250 km of around k^{-5} in this equatorial band in
653 accordance with the G12d5 simulation. Note that it differs from the k^{-4} slope typical of the equatorial
654 region discussed above. It reflects modulation associated with low frequency variability. This 3 month
655 period corresponds with an El Niño event characterized by ~~intense~~ relatively low mesoscale activity
656 in this region of the South West Pacific (Gourdeau et al., 2014). Including the high-frequency but non-
657 tidal forcing increases the smaller-scale energy, and flattens the SSH spectra with slopes of around k^{-4} .
658 This non-tidal high-frequency (< 10 -day) component increases the SSH spectral energy out to scales
659 of 200 km wavelength, suggesting a dominance of rapid small-scale ~~eddy~~ variability of internal gravity
660 waves (~~Farrar and Durland, 2012~~; Garrett and Munk, 1975). But the higher frequency atmospheric
661 forcing and ocean instabilities alone cannot explain the very flat altimetric spectral slopes in this
662 equatorial region.

663 When coherent and incoherent internal tides are included, the spectral slope in the 70-250 km
664 wavelength band becomes very close to that observed with altimetric spectra. This confirms the
665 recent results presented by Savage et al. (2017) for a small box in the eastern tropics, and previously
666 proposed by Richman et al. (2012) and Dufau et al. (2016). The separation of the coherent M2
667 internal tide demonstrates that it clearly contributes SSH energy in the 50-300 km wavelength band,
668 but the incoherent tide, and its cascade of energy into the supertidal frequencies, is the dominant
669 signal at scales less than 50 km. The incoherent and coherent internal tides have similar energy
670 partitioning within the 50-300 km wavelength band.

671

672 **6. Discussion and Conclusion**

673 The processes that could contribute to the flat Sea Surface Height (SSH) wavenumber spectral slopes
674 observed in the tropics by satellite altimetry have been examined in the tropical Pacific. This study
675 has used two complementary approaches to better understand how the equatorial and off-
676 equatorial dynamics impact on the SSH wavenumber spectra. In the first part of this study, we have
677 concentrated on the low-frequency (> 10 days) internal tropical dynamics to better understand how
678 the complex zonal current system and dominant linear tropical waves affect the mainly meridional
679 altimetric SSH wavenumber spectra. In the second part of the study, we have used a high-frequency,
680 high-resolution regional modeling configuration, with and without tides, to explore the high-
681 frequency contributions to the meridional SSH wavenumber spectra.

682 Our $1/12^\circ$, 5-day averaged model confirms the results from previous modeling studies that at
683 seasonal to interannual time scales the most energetic large-scale structures tend to be anisotropic
684 and governed by linear dynamics. At intraseasonal frequencies and in the tropical “mesoscale” band
685 at scales less than 600 km wavelength, one major question was how the cascade of energy is affected

686 by the expected high level of anisotropy and the weak non-linear regimes. Within the “mesoscale”
687 range, the EKE wavenumber spectra ~~is-are~~ isotropic in the off-equatorial regions between 10° and
688 20°, and it is more anisotropic in the equatorial region between 10°N-10°S, with higher level of
689 energy for the meridional EKE spectrum than for the zonal one that reveals larger scales of variability
690 in the zonal direction than in the meridional direction, as expected. ~~the energetic meridional~~
691 ~~perturbations due to instabilities of the larger-scale zonal currents.~~

692 In the off-equatorial range, EKE peaks at around 300 km wavelength, and the steep EKE decrease at
693 smaller wavelength is characterized by spectral slopes between k^{-2} and k^{-3} , which lie between the
694 regimes of SQG and QG turbulence. These weakly nonlinear off-equatorial regions thus have a similar
695 structure to the non-linear mid-latitudes within the range from 100-250 km. In the equatorial band
696 from 10°S-10°N, the total EKE is more energetic than the off-equatorial region, and ~~the spectral~~
697 ~~cascade extends over a large wavenumber range, from 100 to 600 km. Again,~~ the EKE spectral slope
698 ~~within this range~~ approaches k^{-3} over a large wavenumber range, from 100 to 600 km, consistent
699 with QG dynamics, even though there is a strong ageostrophic component here. Using the fixed
700 wavelength (70-250 km) band to estimate “mesoscale” spectral slope leads to a slight reduction in
701 the low-frequency spectral slope estimate, but without a drastic modification. When geostrophic
702 velocities are used to calculate EKE rather than the total surface flow, there is similar spectral energy
703 in the off-equatorial regions at longer wavelengths. In the equatorial band 10°N-10°S, the
704 ageostrophy is more evident with a more marked change in spectral slope based on geostrophic
705 velocities and the beta-approximation at the equator. At large scales in the equatorial band, the
706 ageostrophic equatorial currents are more active, related to the energetic zonal currents. In all
707 regions, at wavelengths shorter than 200 km, the geostrophic spectra become more energetic and
708 the small-scale ageostrophic components are counteracting the balanced geostrophic flow, as found
709 at mid-latitudes (Klein et al., 2008, Ponte et al., 2015). This gives a slightly flatter spectral slope over
710 the 70-250 km wavelength, but the regime remains between k^{-2} and k^{-3} in the off-equatorial region,
711 approaching k^{-2} (and k^{-4} in SSH) in the equatorial band. So using SSH and geostrophic currents slightly
712 flattens the EKE wavenumber spectra, but the modeled SSH wavenumber spectra maintain a steep
713 slope that doesn't match the observed ~~cannot explain alone the very flat~~ altimetric SSH spectra.

714 The choice of regional box size and filtering options also impacts on the spectra. Previous global
715 altimetric studies have calculated alongtrack SSH wavenumber spectra in $10 \times 10^\circ$ boxes, and with
716 varying segment lengths (512 km for Dufau et al., 2016; around 1000 km for Xu and Fu, 2011,
717 Chassignet et al., 2017, etc), and with different tapering or filtering applied (see section 3). In the
718 equatorial band where the EKE peak extends out to 600 km wavelength, it is important to have
719 segment sizes and filtering that preserve this peak and shorter scales. The combined effects of a 10%
720 cosine taper and the short segment lengths leads to a much flatter altimetric SSH spectra, reaching k^{-1}
721 in the Dufau et al (2016) study. We find that the double periodic spectra, the hanning and tukey
722 50% taper filter all give similar results in the tropics, but it is necessary to extend the box size to a
723 minimum of 15° to 20° in segment length or box size in the equatorial band. In the off-equatorial
724 band, these filtering options with a 10° segment length or box size are sufficient. Even with the
725 preferred pre-processing for the altimetric data, and larger segment lengths in our analyses, the
726 altimetric SSH spectra remain quite flat (k^{-2} in the off-equatorial zone, $k^{-1.3}$ in the equatorial band),
727 and do not reflect the steeper, ~~more turbulent~~ spectral slopes predicted by the model.

728 The regional high-resolution models with both high-frequency atmospheric and tidal forcing and
729 high-frequency hourly outputs provide the last pieces of the puzzle. In contrast to previous results
730 based on global ocean models with tidal forcing (Richman et al., 2012; Savage et al., 2017), this 2-
731 model configuration with and without tides, has the same atmospheric and boundary forcing, which
732 allows us to clearly separate the internal tide signals from the high frequency dynamical component.
733 Even though only a small region of the tropical Pacific is available for this analysis, the regional model
734 and the global 1/12° model show similar QG spectral slopes when they are compared over the same
735 domain and with 5-day averaged data. Using hourly data and no tides increases the SSH spectral
736 power at scales smaller than 200 km, possibly due to internal gravity waves in the tropics (Farrar and
737 Durland, 2012; Garrett and Munk, 1975). We note that Rocha et al. (2016) found a similar increase in
738 their detided alongtrack model runs in Drake Passage, but at scales less than 40 km wavelength, far
739 below the noise level of our present altimeter constellation. In the tropics, this contribution of high-
740 frequency non-tidal SSH signals out to 200 km wavelength will also impact on today's alongtrack
741 altimeter constellation, whose noise levels block ocean signals at scales less than 70 km for Jason
742 class satellites, and 30-50 km for Saral and Sentinel-3 SAR altimeters (Dufau et al., 2016). So non-tidal
743 internal gravity waves will partially contribute to the higher small-scale SSH variance and flatter
744 spectral slopes in today's altimetric SSH data.

745 The regional model with tides shows the very important contribution of internal tides to the flat SSH
746 slopes in the tropics. We have separated out the predictive part of the barotropic tide and internal
747 tides, since open ocean barotropic tides are well corrected for in altimetric data today (Lyard et al.,
748 2018; Stammer et al., 2014), and corrections are becoming available for the coherent part of the
749 internal tide (Ray and Zaron, 2016). In this open ocean tropical region east of the Solomon Sea, when
750 coherent and incoherent internal tides are included, the spectral slope in the 70-250 km wavelength
751 band becomes very close to that observed with altimetric spectra. This confirms the recent results
752 presented by Savage et al. (2017) for a small box in the eastern tropics, and previously proposed by
753 Richman et al. (2012) and Dufau et al. (2016). The separation of the coherent M2 internal tide
754 demonstrates that it clearly contributes significant SSH energy in the 50-300 km wavelength band,
755 but around the main internal tide wavelengths, there is a strong signature of M2 incoherent internal
756 tide. The incoherent tide, and its cascade of energy into the supertidal frequencies, is the dominant
757 signal at scales less than 50 km. This strong incoherent internal tide is consistent with recent studies
758 that suggest that internal tides interacting with energetic zonal jets can generate a **predominate**
759 **major** incoherent internal tide (Ponte and Klein, 2015), and may explain the reduction of the
760 coherent internal tides in the equatorial band in global models (Shriver et al., 2014) and altimetric
761 analyses (Ray and Zaron, 2016). Our model highlights that the internal tide signal is strong in this
762 equatorial region, and the incoherent tide accounts for 35% of the SSH spectral power in the 50-300
763 km wavelength band, and is not predictable.

764 These results have important consequences for the analyses of alongtrack altimetric data today, and
765 for the future high-resolution swath missions such as SWOT. Today's constellation of satellite
766 altimeters have their alongtrack data filtered to remove noise at scales less than 70 km for all
767 missions (Dibarboure et al., 2014; Dufau et al., 2016), and these data are now being used with no
768 internal tide correction in the global gridded altimetry maps of SSH and geostrophic currents. The
769 imprint of these internal tides is evident in the alongtrack data (see Fig. 1b from Dufau et al., 2016)
770 but is also present in the gridded maps (R. Ray, personal communication). In the future, a coherent
771 internal tide correction may be applied to the alongtrack data based on Ray and Zaron (2016), to

772 reduce some of this non-balanced signal. It is particularly important to remove the unbalanced
773 internal wave signals from SSH before calculating geostrophic currents. But it is clear that the
774 incoherent internal tide and internal gravity waves reach scales of 200 km in the tropics, and their
775 signature in SSH remains a big issue for detecting balanced internal ocean currents from alongtrack
776 altimetry and the future SWOT wide-swath altimeter mission. Removing this signal to detect purely
777 balanced motions will be challenging, since filtering over 200 km removes much of the small-scale
778 ocean dynamics of interest in the tropics. On the other hand, there will also be a great opportunity to
779 investigate the interaction of the internal tide and ocean dynamics in the tropics in the future, with
780 both models and fine-scale altimetric observations.

781

782 Acknowledgments

783 The authors wish to acknowledge Ssalto/Duacs AVISO who produced the altimeter products, with
784 support from CNES (<http://www.avisio.altimetry.fr/duacs/>). The authors would like to thank the
785 DRAKKAR team for providing them with the high resolution global ocean simulation, and especially
786 J.M. Molines for his support. This work benefited from discussions with J. Jouanno, F. Marin, Y. Morel
787 from LEGOS. We particularly thank [Tom Farrar \(WHOI\) and an anonymous reviewer for their](#)
788 [constructive comments, and](#) J. Verron (IGE), C. Menesguen (LOPS), and X. Capet (LOCEAN) for their
789 time, and their fruitful comments. M. Tchilibou is funded by Université de Toulouse 3. L. Gourdeau,
790 G. Sérazin are funded by IRD; R. Morrow is funded by CNAP, and B. Djath was funded by CNES. This
791 work is a contribution to the joint CNES/NASA SWOT project "SWOT in the tropics" and is supported
792 by the French TOSCA programme.

793

794 [Annexe 1 Model configurations used in this study](#)

795 [- Global Model at 1/12°](#)

796 [The model used is the ORCA12.L46-MAL95 configuration of the global 1/12° OGCM developed and](#)
797 [operated in the DRAKKAR consortium \(\[www.ifremer.fr/lpo/drakkar\]\(http://www.ifremer.fr/lpo/drakkar\)\) \(Lecointre et al, 2011\). The](#)
798 [numerical code is based on the oceanic component of the NEMO \(Nucleus for European Modelling of](#)
799 [the Ocean\) system \(Madec, 2008\). The model formulation is based on standard primitive equations.](#)
800 [The equations are discretized on the classical isotropic Arakawa C grid using a Mercator projection.](#)
801 [Geopotential vertical coordinates are used with 46 levels with a 6m resolution in the upper layers](#)
802 [and up to 250 m in the deepest regions \(5750 m\). The "partial step" approach is used \(Adcroft et al,](#)
803 [1997\) to allow the bottom cells thickness to be modified to fit the local bathymetry. This approach](#)
804 [clearly improves the representation of topography effects \(Barnier et al. 2006; Penduff et al. 2007\).](#)
805 [The bathymetry was built from the GEBCO1 dataset](#)
806 [\(\[http://www.gebco.net/data_and_products/gebco\]\(http://www.gebco.net/data_and_products/gebco\)\) for regions shallower than 200m and from](#)
807 [ETOPO2 \(\[www.ngdc.noaa.gov/mgg/global/relief/ETOPO2\]\(http://www.ngdc.noaa.gov/mgg/global/relief/ETOPO2\)\) for regions deeper than 400m \(with a](#)
808 [combination of both datasets in the 200m-400m depth range\). Lateral boundary conditions for](#)
809 [coastal tangential velocity have a strong impact on the stability of boundary currents \(Verron and](#)
810 [Blayo, 1996\). Based on sensitivity experiments, a "partial-slip" condition is chosen, where the coastal](#)
811 [vorticity is not set to 0 \("free slip" condition\), but is weaker than in the "no-slip" condition. The](#)

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812 atmospheric forcing (both mechanical and thermodynamical) is applied to the model using the CORE
813 bulk-formulae approach (Large and Yeager, 2004, 2009).

814 The simulation started from rest in 1978 with initial conditions for temperature and salinity provided
815 by the 1998 World Ocean Atlas (Levitus, 1998). It was spun up for 11 years using the CORE-II forcing
816 dataset and then integrated from 1989 to 2007 using a 3-hourly ERA-interim forcing (Dee et al.,
817 2011). The 3D velocities, and the 2D Sea Surface Height (SSH) are saved as 5-day means during the
818 period of integration. In the following, it is referenced as G12d5. The domain considered in this study
819 covers the tropical Pacific between 20°N -20°S. This simulation has been used to document
820 mesoscale variability in the South West Pacific Solomon Sea (Gourdeau et al., 2014; Gourdeau et al.,
821 2017). The present study will analyse this simulation over the tropical Pacific.

822 - Regional Model at 1/36° with and without tides

823 As part of the CLIVAR/SPICE program, regional simulations of the Solomon Sea in the South Western
824 tropical Pacific have been performed (Ganachaud et al., 2014). The numerical model of the Solomon
825 Sea used in this study has a 1/36° horizontal resolution, and 75 vertical levels. It is based on the same
826 oceanic component as the NEMO system presented above. This 1/36° resolution model is embedded
827 into the global 1/12° ocean model presented above and one-way controlled using an open boundary
828 strategy (Tréguier et al., 2001). Its horizontal domain is shown on Fig. 3b. The bathymetry of the high-
829 resolution Solomon Sea model is based on the GEBCO08 dataset. Atmospheric boundary conditions,
830 consisting in surface fluxes of momentum, heat and freshwater, are diagnosed through classical bulk
831 formulas (Large and Yeager, 2009). Wind and atmospheric temperature and humidity are provided
832 from the 3-hourly ERA Interim reanalysis (Dee et al., 2011). A first version of the regional model with
833 45 vertical levels has been initialized with the climatological mass field of the World Ocean Atlas
834 (Levitus et al., 1998) and was integrated from 1989 to 2007. More technical details on this
835 configuration may be found in Djath et al. (2014). The new version used here is distinct from the
836 former version by the number of vertical levels (75 levels in the new version) but above all by its
837 ability to take account realistic tidal forcing (Tchilibou et al., 2018). It is of particular importance in
838 this area where internal tides are particularly active (Niwa and Hibiwa, 2011; Gourdeau, 1998), and
839 could modify accordingly the energy flux for the meso and submesoscale bands (Richman et al.,
840 2012). The model is forced at the open boundary by prescribing the first 9 main tidal harmonics (M2,
841 S2, N2, K2, K1, O1, P1, Q1, M4) as defined from the global tides atlas FES2014 (Lyard et al., 2018)
842 through a forced gravity wave radiation condition. The model is initialized by the outputs from the
843 ORCA 1/12° version. Two simulations are performed: one without the tidal forcing (R36) over the
844 1992-2012 period, and one with the tidal forcing (R36T) over the 1992-2009 period. Model outputs
845 are saved as daily mean (R36(T)d), and instantaneous fields are saved hourly (R36(T)h) during a 3
846 month period from January-March 1998.

847 Annexe 2 : Spectral sensitivity tests

848 We tested the sensitivity of our G12d5 model's SSH wavenumber spectrum to different tapering
849 windows and the double periodic method, using different data length sizes, and in one or two
850 dimensions. The following steps were performed for these test spectra, evaluated within 10°S-
851 10°N/160°W-120°W: the model data are extracted meridionally and zonally in fixed segment lengths
852 of 5°, 10°, 20° and within a 20°X20° square box; the mean and linear trend (fitted plane for two-
853 dimensional case) were removed from each data segment or box; the filter window (Tk01, Tk05,

Commentaire [RM2]: There is no Fig 3b ?

854 [Hann](#) or [Dbp](#) are applied ; temporal and spatial (longitude, latitude) series spectra are calculated and
855 [averaged in Fourier space. The results are shown in Fig. A1.](#)

856 [Tk01 meridional spectra in the tropics are the most perturbed by the short segment lengths \(Fig.](#)
857 [A1a\). In the 70-250 km range commonly used to define a global mesocale band \(delimited by the](#)
858 [green vertical lines\), the spectral slope flattens as the data segment length decreases. 5° segment](#)
859 [spectra with a Tk01 window have a \$k^{-1.3}\$ slope, which explains the very shallow slope in the tropics](#)
860 [observed by Dufau et al. 2016 who applied this short data segment size and a Tk01 window.](#)
861 [Meridional spectra differ primarily at larger scales from 100-500 km, when short segment lengths are](#)
862 [used \(Fig. A1a\). A comparison of the meridional spectrum using 20° segments and different windows](#)
863 [\(Tk01, Tk05, Hann and Dbp\) are shown in Figure A1b. Even with the 20° segments, Tk01 is distorted.](#)
864 [On the other hand, the Tk05, Hann and Dbp match well, with a near linear cascade of energy over](#)
865 [the 30-1000 km wavelength range, and are more adapted for the tropics since they capture the main](#)
866 [range of SSH mesoscale dynamics, particularly the spectral energy peaks around 1000 km](#)
867 [wavelength.](#)

868 [Similar calculations were performed for the zonal spectra \(not shown\) and confirm that the Tk01](#)
869 [method deforms the zonal spectra and flattens the spectral slope within the 70-250 km wavelength](#)
870 [band as the data segment size decreases. Tk05, Hann and Dbp 20° segment spectra match, although](#)
871 [the Dbp has more noise at small scale.](#)

872 [We also conducted a sensitivity test in the off-equatorial region \(not shown\): Flattening and](#)
873 [deformation of the spectrum by Tk01 persist, but the 10° segments or 10° square box are long](#)
874 [enough to capture the off-equatorial dynamics.](#)

875 [The particular sensitivity of spectra in the tropics to the choice of spectral segment length and](#)
876 [windowing is linked to energetic EKE and SSH signals extending out to longer wavelengths, and their](#)
877 [spectral leakage from low to high wavenumbers. Tk01 gives the worst performance in the tropics,](#)
878 [and the distortion of spectra is amplified for short data segments. Both the Tk05 and the Hann](#)
879 [windowing are a good compromise for preserving much of the original signal and reducing leakage,](#)
880 [but they need to be applied over larger segments..](#)

881 [So, to safely avoid leakage in the tropics, it is best to use a long record and an effective taper](#)
882 [window. We do not advise to use the Tk01 filter window. The Tk05 or Hann filters give convincing](#)
883 [results in the equatorial band, with a minimum of 15° to 20° needed in segment lengths. In the off-](#)
884 [equator region, 10° data segments or 10°X10° boxes are sufficient. We choose to use the Tukey 0.5](#)
885 [filter in the paper.](#)

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1098

1099 **Figure captions**

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1101 Figure 1: a) Spatial distribution of altimetric alongtrack SSH wavenumber spectral slope calculated in
1102 the fixed 70-250 km mesoscale range (from Xu and Fu, 2011; their Fig. 2). b) Latitudinal dependence
1103 of the altimetric SSH alongtrack wavenumber spectra in the Atlantic Ocean (from Dufau et al., 2016;
1104 their Fig. 3). The colors of the spectra refer to the geographical boxes where alongtrack data were
1105 averaged on the right.

1106 ~~Figure 2: Sensitivity experiments for different spectral processing techniques applied to meridional~~
1107 ~~SSH wavenumber spectra representative of the equatorial region. a) SSH wavenumber spectra using~~
1108 ~~a Tukey 0.1 window (blue) and a Tukey 0.5 window (red) depending on segment lengths: 5° (dots),~~
1109 ~~10° (dash), 20° (line). b) SSH wavenumber spectra using different windowing over a 20° segment~~
1110 ~~length: Tukey 0.1 window (Tk01, blue), Tukey 0.5 window (Tk05, red), Hanning window (Han, green).~~
1111 ~~The double periodic method (Dbp, black) is also tested. For reference k^{-4} and k^{-5} curves are plotted.~~
1112 ~~Units: cm^2/cpkm~~

1113 Figure 32: Snapshot of relative vorticity of the 1/12° G12d5 simulation. unit in $1.E5 \text{ s}^{-1}$. The yellow
1114 lines delineate the equatorial and off-equatorial regions. The dashed lines delineate square boxes for
1115 the different regions to compute wavenumber spectra. The black arrows illustrate the main zonal
1116 tropical currents (SEC: South Equatorial Current, SECC: South Equatorial CounterCurrent, NECC:
1117 North Equatorial CounterCurrent, NEC: North Equatorial Current, STCC: SubTropical CounterCurrent,
1118 HLCC: Hawaiiin Lee Counter Current).

1119 Figure 43: a) latitudinal distribution of the EKE Frequency power spectra computed at each model
1120 grid point of the G12d5 simulation, and averaged in longitude. The black line is the critical period
1121 from Lin et al. (2008). b) longitudinal distribution of the EKE Frequency power spectra computed at
1122 each model grid point of the G12d5 simulation, and averaged between 20°S-20°N. Units are in \log_{10}
1123 of $\text{cm}^2/\text{s}^2/\text{cpday}$.

1124 Figure 54: Zonal wavenumber-frequency EKE spectra averaged over a) 10°N-20°N region, b) 10°S-
1125 10°N region, and c) 10°S-20°S region. d) Meridional wavenumber-frequency EKE spectra covering
1126 20°S-20°N averaged over the 120°W-150°W region. Superimposed on a) and c) are the theoretical
1127 dispersion curves for the first mode baroclinic waves. Superimposed on b) are the theoretical
1128 dispersion curves for the first 3 baroclinic wave modes, and the Kelvin wave mode. Units are in \log_{10}
1129 of $\text{cm}^2/\text{s}^2/\text{cpday}/\text{cpkm}$.

1130 Figure 65: Zonal wavenumber EKE spectra averaged over the equatorial (orange line), and off-
1131 equatorial latitude bands (north: green; south: blue). Units are in $\text{cm}^2/\text{s}^2/\text{cpkm}$.

1132 Figure 76: ~~1D~~-zonal (orange), meridional (green); and ~~2D isotropiemagnitude~~ (blue) EKE wavenumber
1133 spectra averaged over a) 10°N-20°N, b) 10°S-10°N, and c) 10°S-20°S regions. The ~~isotropic magnitude~~
1134 geostrophic EKE wavenumber spectra- is also shown (EKEg, blue dash line). The vertical green dash
1135 lines delineate the fixed 70-250 km mesoscale range. For reference, k^{-2} and k^{-3} curves are plotted
1136 (black lines). Units are in $\text{cm}^2/\text{s}^2/\text{cpkm}$.

1137 | Figure 87: Meridional SSH wavenumber spectra averaged over the equatorial (orange), and off-
1138 | equatorial latitude bands (north: green, south:blue) for the G12d5 simulation (line). Topex-Poseidon
1139 | along track altimetric SSH wavenumber spectra are averaged over the same latitude bands (dash).
1140 | Units are in cm^2/cpkm .

1141 | Figure 98: SSH variability of the $1/36^\circ$ regional model with explicit tides (R36Th) over the 3 month
1142 | simulation for a) the mesoscale signal, and b) the internal waves and internal tides defined by a 48 hr
1143 | cutoff period. Units in cm^2 . The SARAL/AltiKA (black line) and Jason-2 (dash line) tracks used to
1144 | compute the altimetric spectra in Fig. 10 are superimposed.

1145 | Figure 109: Meridional SSH wavenumber spectra averaged over 163°E - 165°E for the hourly outputs
1146 | of the $1/36^\circ$ resolution regional model without tides (R36h, green), and 5 day averaged outputs
1147 | (R36d5, orange). Meridional SSH spectra of the G12d5 simulation is in cyan. SSH meridional
1148 | wavenumber spectra for the hourly outputs of the $1/36^\circ$ regional model with explicit tides once the
1149 | barotropic tides has been removed (R36Th-BT, in blue). The spectrum of the coherent baroclinic tides
1150 | has been added to the spectrum of the model without tides (R36h+BC, purple), the contribution of
1151 | the only M2 coherent baroclinic tide is in red (R36h+M2BC). The difference between the blue and
1152 | purple curves corresponds with the incoherent internal tides. The corresponding along track SSH
1153 | altimetric spectra for SARAL/AltiKa (line) and Jason-2 (dash) are in black. Units are in cm^2/cpkm .

1154 |
1155 | Figure 2A1: Sensitivity experiments for different spectral processing techniques applied to meridional
1156 | SSH wavenumber spectra representative of the equatorial region. a) SSH wavenumber spectra using
1157 | a Tukey 0.1 window (blue) and a Tukey 0.5 window (red) depending on segment lengths: 5° (dots),
1158 | 10° (dash), 20° (line). b) SSH wavenumber spectra using different windowing over a 20° segment
1159 | length: Tukey 0.1 window (Tk01, blue), Tukey 0.5 window (Tk05, red), Hanning window (Han, green).
1160 | The double periodic method (Dbp, black) is also tested. For reference k^{-1} and k^{-5} curves are plotted.
1161 | Units : cm^2/cpkm

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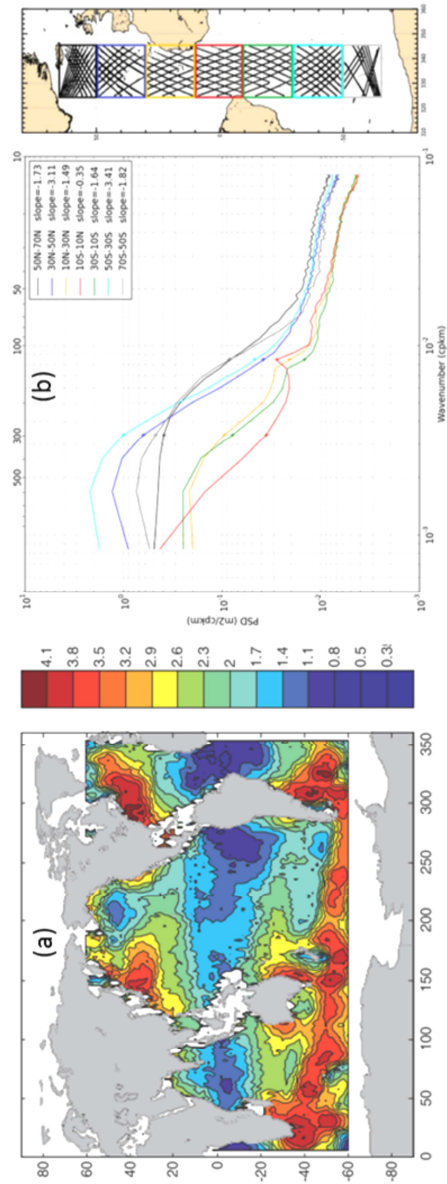
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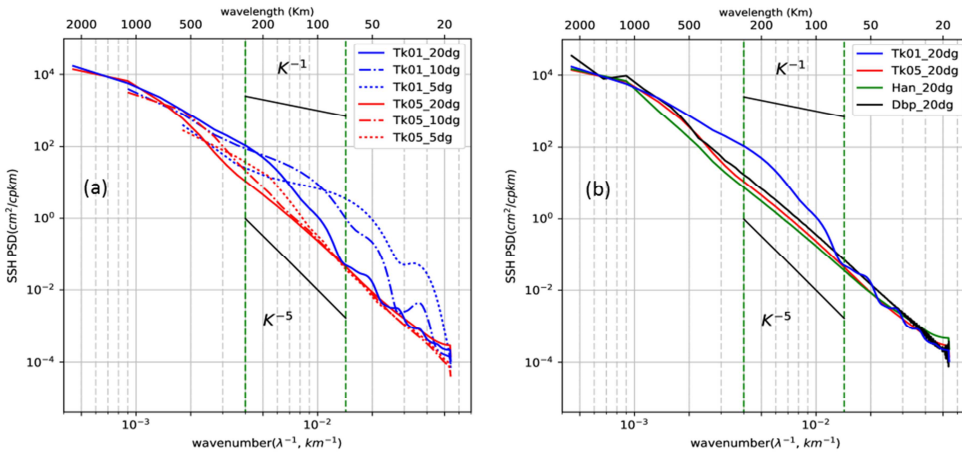
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1175 Figure 1: a) Spatial distribution of altimetric alongtrack SSH wavenumber spectral slope calculated in
1176 the fixed 70-250 km mesoscale range (from Xu and Fu, 2011; their Fig. 2). b) Latitudinal dependence
1177 of the altimetric SSH alongtrack wavenumber spectra in the Atlantic Ocean (from Dufau et al., 2016;
1178 their Fig. 3). The colors of the spectra refer to the geographical boxes where alongtrack data were
1179 averaged on the right.

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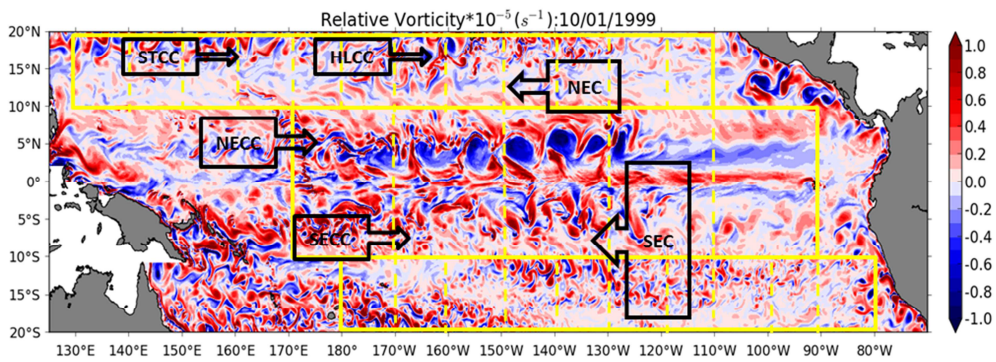


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1184 Figure 2: Sensitivity experiments for different spectral processing techniques applied to meridional
 1185 SSH wavenumber spectra representative of the equatorial region. a) SSH wavenumber spectra using
 1186 a Tukey 0.1 window (blue) and a Tukey 0.5 window (red) depending on segment lengths: 5° (dots),
 1187 10° (dash), 20° (line). b) SSH wavenumber spectra using different windowing over a 20° segment
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 1189 The double periodic method (Dbp, black) is also tested. For reference k^{-1} and k^{-5} curves are plotted.
 1190 Units: cm^2/cpkm

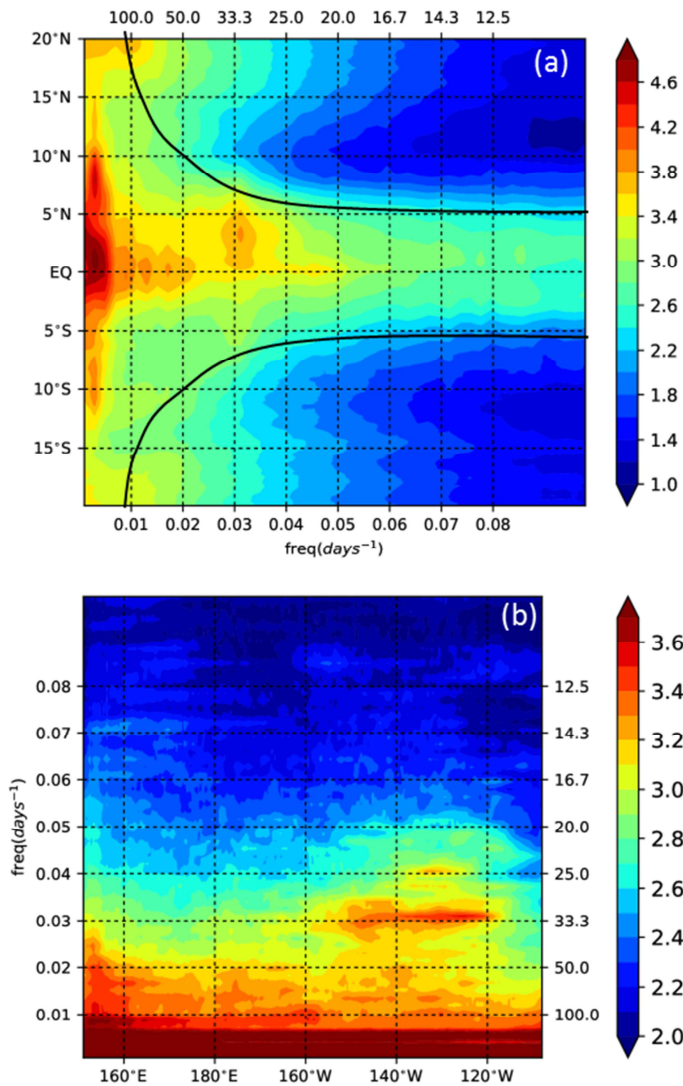
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1193 | Figure 23: Snapshot of relative vorticity of the 1/12° G12d5 simulation. unit in $1.E5 s^{-1}$. The yellow
 1194 | lines delineate the equatorial and off-equatorial regions. The dashed lines delineate square boxes for
 1195 | the different regions to compute wavenumber spectra. The black arrows illustrate the main zonal
 1196 | tropical currents (SEC: South Equatorial Current, SECC: South Equatorial CounterCurrent, NECC:
 1197 | North Equatorial CounterCurrent, NEC: North Equatorial Current, STCC: SubTropical CounterCurrent,
 1198 | HLCC: [Hawaiian Lee Counter Current](#)).

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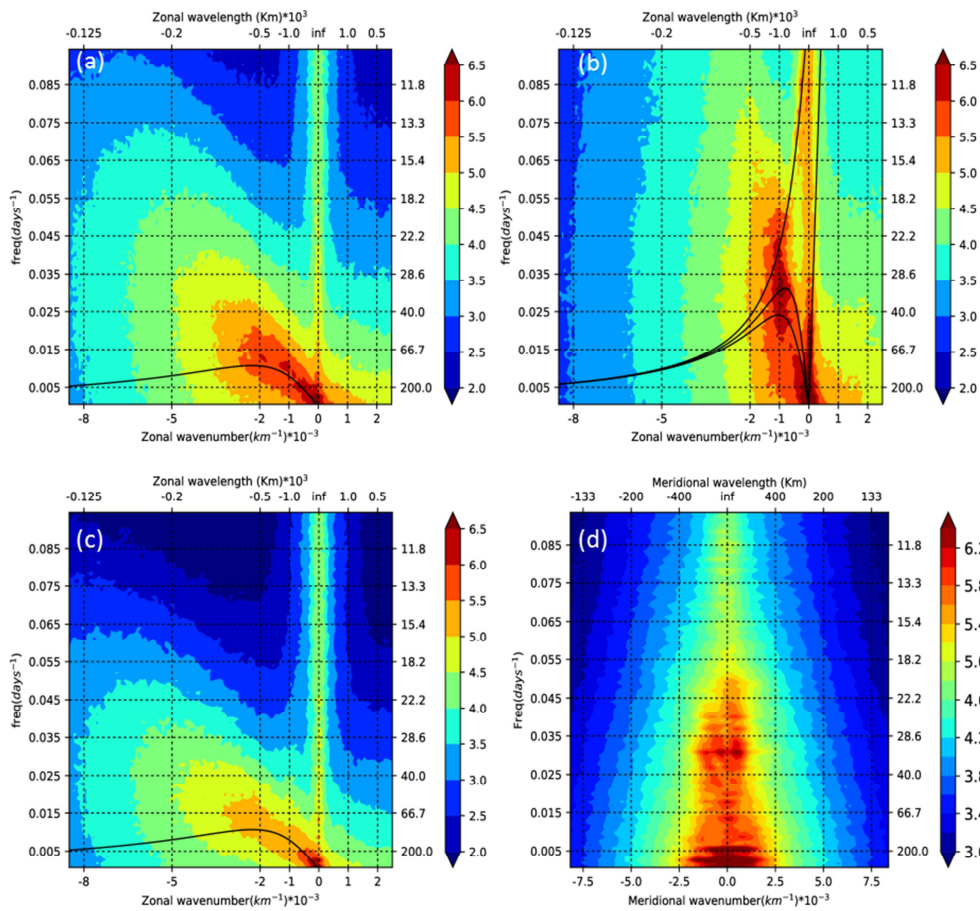


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1202 | Figure-43: a) latitudinal distribution of the EKE Frequency power spectra computed at each model
 1203 grid point of the G12d5 simulation, and averaged in longitude. The black line is the critical period
 1204 from Lin et al. (2008). b) longitudinal distribution of the EKE Frequency power spectra computed at
 1205 each model grid point of the G12d5 simulation, and averaged between 20°S-20°N. Units are in log₁₀
 1206 of cm²/s²/cpday.

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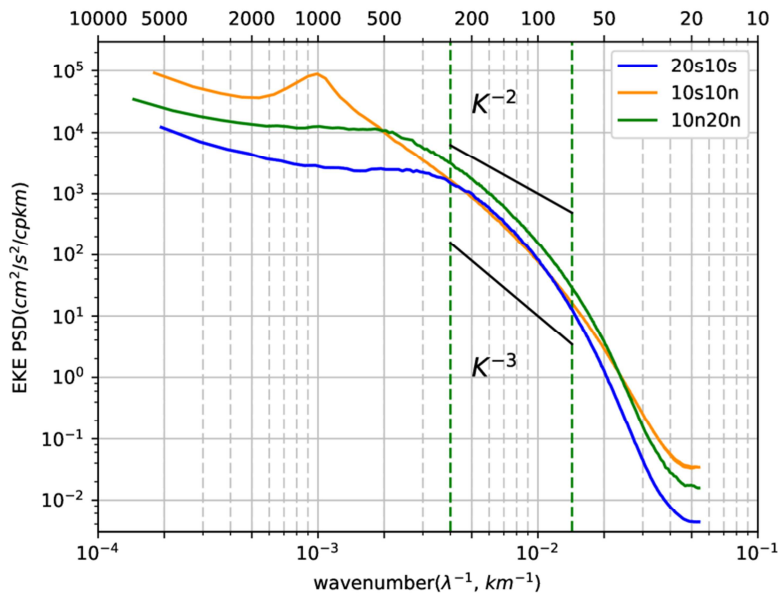


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1210 | Figure 54: Zonal wavenumber-frequency EKE spectra averaged over a) 10°N-20°N region, b) 10°S-
 1211 10°N region, and c) 10°S-20°S region. d) Meridional wavenumber-frequency EKE spectra covering
 1212 20°S-20°N averaged over the 120°W-150°W region. Superimposed on a) and c) are the theoretical
 1213 dispersion curves for the first mode baroclinic waves. Superimposed on b) are the theoretical
 1214 dispersion curves for the first 3 baroclinic wave modes, and the Kelvin wave mode. Units are in log₁₀
 1215 of cm²/s²/c/day/c/pkm.

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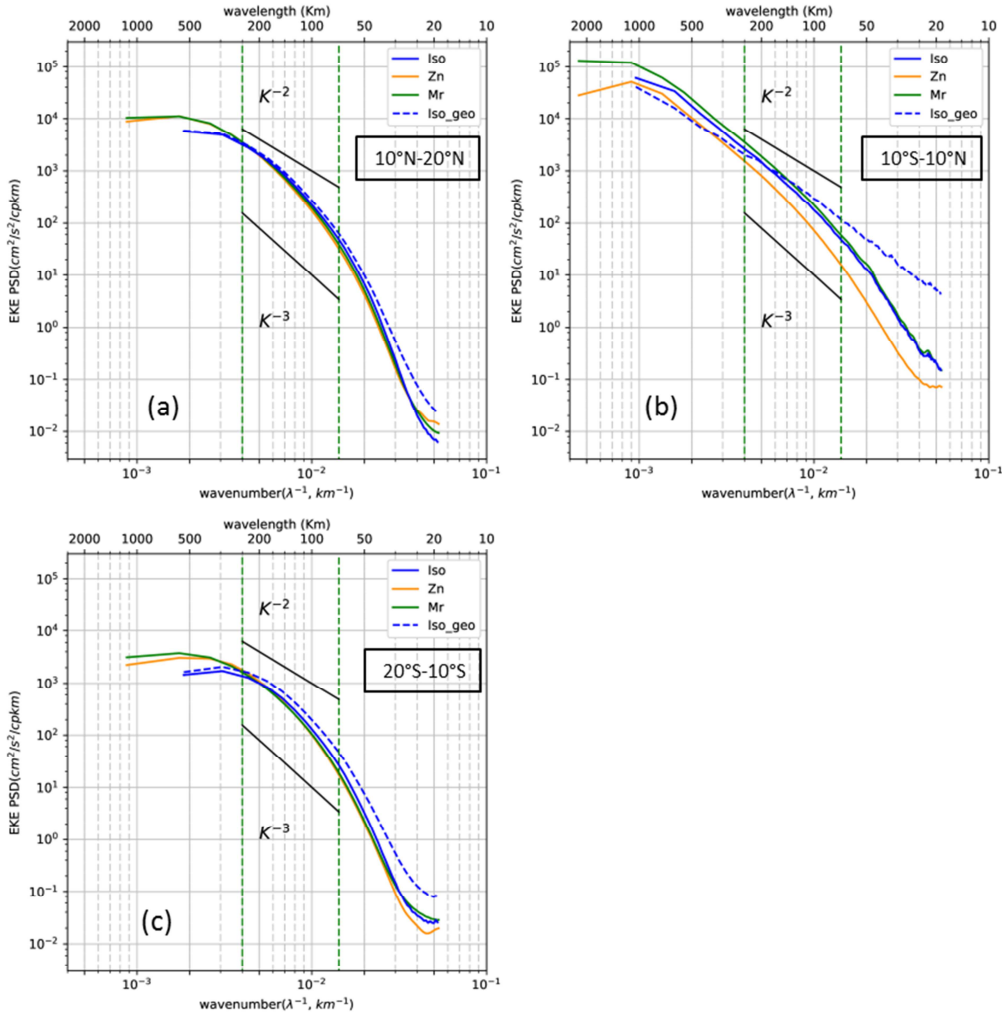
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1219 | Figure 65: Zonal wavenumber EKE spectra averaged over the equatorial (orange line), and off-
 1220 equatorial latitude bands (north: green; south: blue). Units are in cm²/s²/cpkm.

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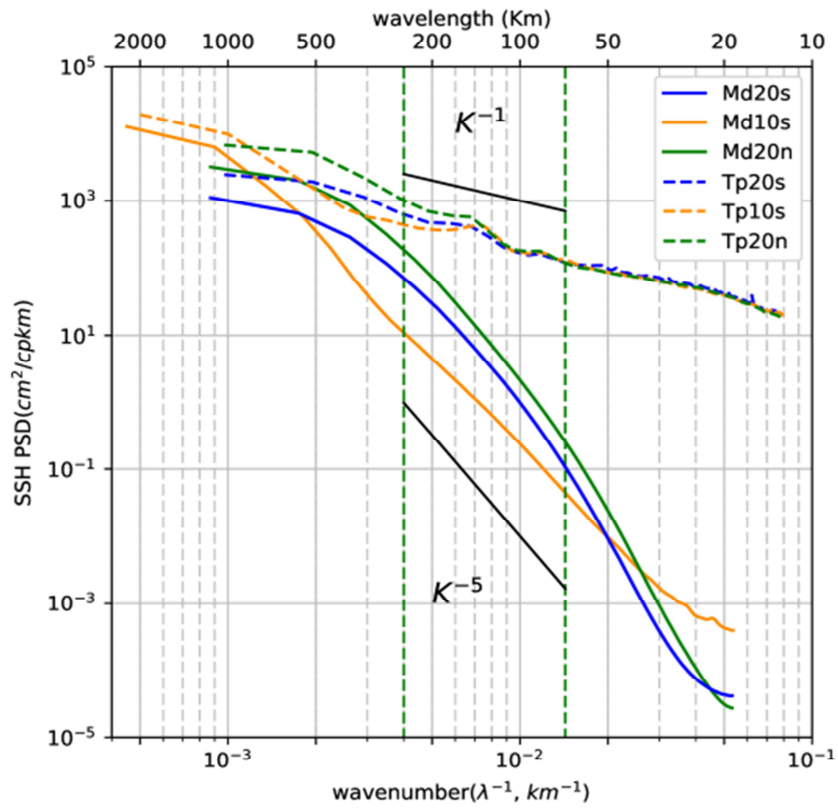


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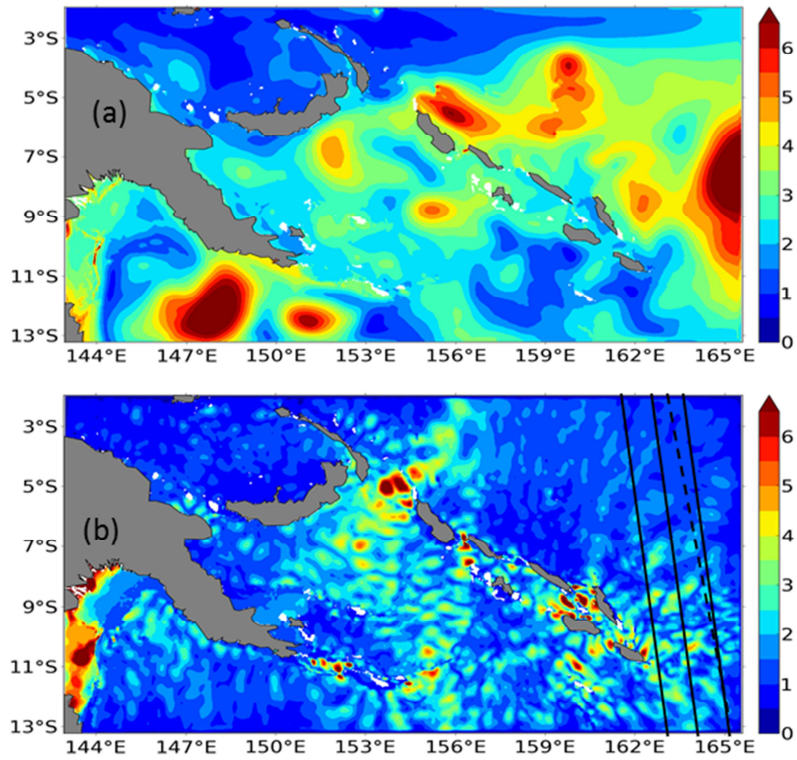
1225 Figure 76: 1D-zonal (orange), meridional (green); and 2D-isotropic magnitude (blue) EKE wavenumber
1226 spectra averaged over a) 10°N-20°N, b) 10°S-10°N, and c) 10°S-20°S regions. The isotropic magnitude
1227 geostrophic EKE wavenumber spectra is also shown (EKEg, blue dash line). The vertical green dash
1228 lines delineate the fixed 70-250 km mesoscale range. For reference, k^{-2} and k^{-3} curves are plotted
1229 (black lines). Units are in $\text{cm}^2/\text{s}^2/\text{cpkm}$.

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1232 | Figure 87: Meridional SSH wavenumber spectra averaged over the equatorial (orange), and off-
 1233 equatorial latitude bands (north: green, south:blue) for the G12d5 simulation (line). Topex-Poseidon
 1234 along track altimetric SSH wavenumber spectra are averaged over the same latitude bands (dash).
 1235 Units are in cm^2/cpkm .
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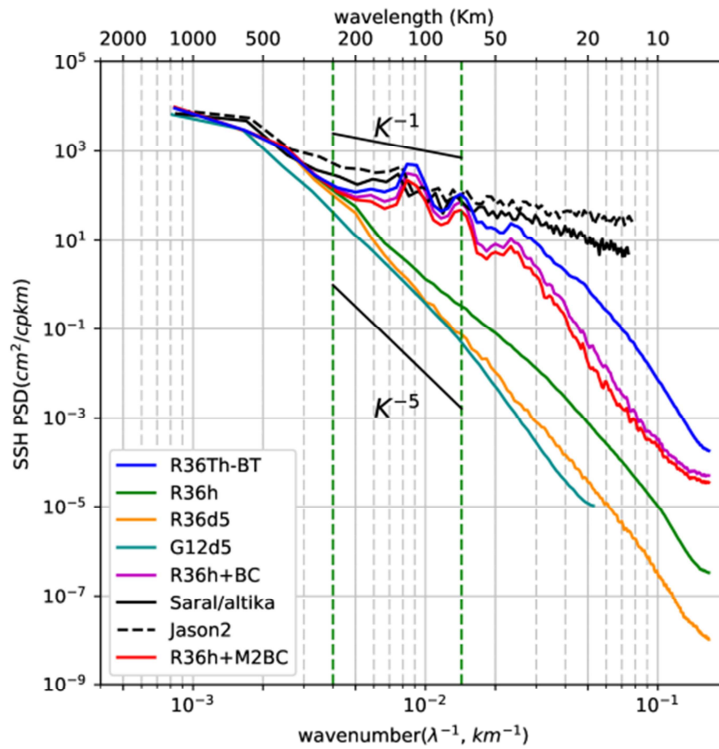


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1239 Figure 98: SSH variability of the 1/36° regional model with explicit tides (R36Th) over the 3 month
1240 simulation for a) the mesoscale signal, and b) the internal waves and internal tides defined by a 48 hr
1241 cutoff period. Units in cm^2 . The SARAL/AltiKA (black line) and Jason-2 (dash line) tracks used to
1242 compute the altimetric spectra in Fig. 10 are superimposed.

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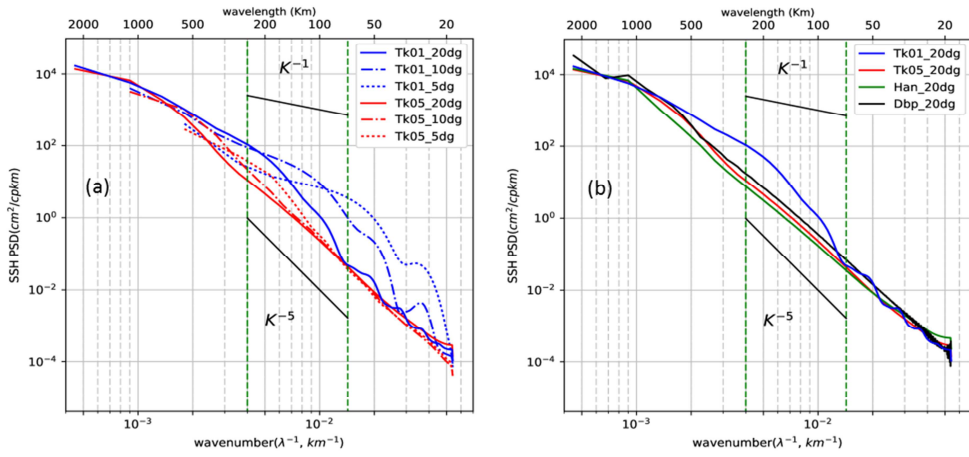
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1247 | Figure 409: Meridional SSH wavenumber spectra averaged over 163°E-165°E for the hourly outputs
1248 | of the 1/36° resolution regional model without tides (R36h, green), and 5 day averaged outputs
1249 | (R36d5, orange). Meridional SSH spectrum of the G12d5 simulation is in cyan. SSH meridional
1250 | wavenumber spectra for the hourly outputs of the 1/36° regional model with explicit tides once the
1251 | barotropic tides has been removed (R36Th-BT, in blue). The spectrum of the coherent baroclinic tides
1252 | has been added to the spectrum of the model without tides (R36h+BC, purple), the main
1253 | contribution of the onlys from the M2 coherent baroclinic tide, shown is in red (R36h+M2BC). The
1254 | difference between the blue and purple curves corresponds with the incoherent internal tides. The
1255 | corresponding along track SSH altimetric spectra for SARAL/AltiKa (black line) and Jason-2
1256 | (black dashed) are in black shown. Units are in cm²/cpkm.
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[Figure A1: Sensitivity experiments for different spectral processing techniques applied to meridional SSH wavenumber spectra representative of the equatorial region. a\) SSH wavenumber spectra using a Tukey 0.1 window \(blue\) and a Tukey 0.5 window \(red\) depending on segment lengths: 5° \(dots\), 10° \(dash\), 20° \(line\). b\) SSH wavenumber spectra using different windowing over a 20° segment length: Tukey 0.1 window \(Tk01, blue\), Tukey 0.5 window \(Tk05, red\), Hanning window \(Han, green\). The double periodic method \(Dbp, black\) is also tested. For reference \$k^{-1}\$ and \$k^{-5}\$ curves are plotted. Units : cm²/cpkm](#)