



1                   **Numerical modeling of surface wave development under the action of wind**

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*DMITRY CHALIKOV*

4

*Shirshov Institute of Oceanology, Saint Petersburg 199053, Russia*

5

*Russian State Hydrometeorological University, Saint Petersburg 195196*

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*University of Melbourne, Victoria 3010, Australia*

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**Abstract**

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10 The numerical modeling of two-dimensional surface wave development under the action of wind  
11 is performed. The model is based on three-dimensional equations of potential motion with free  
12 surface written in a surface-following non-orthogonal curvilinear coordinate system where depth  
13 is counted from moving surface. Three-dimensional Poisson equation for velocity potential is  
14 solved iteratively. Fourier transform method, the second-order accuracy approximation of  
15 vertical derivatives on a stretched vertical grid and the fourth-order Runge-Kutta time stepping  
16 are used. Both the input energy to waves and dissipation of wave energy are calculated on the  
17 basis of the earlier developed and validated algorithms. A one-processor version of the model for  
18 PC allows us to simulate an evolution of wave field with thousands degrees of freedom over  
19 thousands of wave periods. A long-time evolution of two-dimensional wave structure is  
20 illustrated by the spectra of wave surface and input and output of energy.

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**1. Introduction**

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Development of waves under the action of wind is a process that is difficult to simulate since surface waves are very conservative and change their energy for hundreds and thousands periods. This is why the most popular method is spectral modeling. Waves as physical objects in this approach are actually absent, since evolution of spectral distribution of wave energy is simulated. The description of input and dissipation in this approach is not connected directly with the formulation of the problem, but it is rather adopted from other branches of wave theory where waves are the objects of investigation. However, the spectral approach turned out to be the only method capable to describe the space and time evolution of wave field in the ocean. The phase resolving models (or 'direct' models) designed for reproducing waves themselves cannot compete with spectral models since a typical size of domain in such models does not exceed several kilometers. Such domain includes just several thousands of large waves. Nevertheless, direct wave modeling plays an ever-increasing role in geophysical fluid dynamics, because it gives the possibility to investigate the processes which cannot be reproduced with spectral models. One of such problems is that of extreme wave generation. (Chalikov, 2009; Chalikov, Babanin, 2016a). Direct modeling is also a perfect instrument for development of parameterization of physical processes for spectral wave models. Besides, such models can be used for direct simulation of wave regimes of small water basins, for example, port harbors. Other approaches of direct modeling are discussed in (Chalikov et al. 2014; Chalikov, 2016)

Until recently, direct modeling was used for reproduction of quasi-stationary wave regime when wave spectrum essentially did not change. A unique example of direct numerical modeling of surface wave evolution is given in (Chalikov and Babanin, 2014) where development of wave field was calculated with use of a two-dimensional model based on full



46 potential equations written in the conformal coordinates. The model included algorithms for  
 47 parameterization of input and dissipation of energy (a description of similar algorithms is given  
 48 below). The model successfully reproduced an evolution of wave spectrum under the action of  
 49 wind. However, strictly one-dimensional (unidirectional) waves are not realistic; hence, the full  
 50 problem of wave evolution should be formulated on the basis of three-dimensional equations. An  
 51 example of such modeling is given in the current paper.

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## 53 2. Equations

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55 Let us introduce a *non-stationary surface-following non-orthogonal* coordinate system:

$$56 \quad \xi = x, \quad \vartheta = y, \quad \zeta = z - \eta(\xi, \vartheta, \tau), \quad \tau = t, \quad (1)$$

57 where  $\eta(x, y, t) = \eta(\xi, \vartheta, \tau)$  is a moving periodic wave surface given by the Fourier series

$$58 \quad \eta(\xi, \vartheta, \tau) = \sum_{-M_x < k < M_x} \sum_{-M_y < l < M_y} h_{k,l}(\tau) \Theta_{k,l}, \quad (2)$$

59 where  $M_x$  and  $M_y$  are the numbers of modes in directions  $\xi$  and  $\vartheta$ , respectively, while  $\Theta_{k,l}$  are  
 60 Fourier expansion basis functions. The 3-D equations of potential waves in the system of  
 61 coordinates (1) at  $\zeta \leq 0$  take the following form:

$$62 \quad \eta_\tau = -\eta_\xi \varphi_\xi - \eta_\vartheta \varphi_\vartheta + (1 + \eta_\xi^2 + \eta_\vartheta^2) \Phi_\zeta, \quad (3)$$

$$63 \quad \varphi_\tau = -\frac{1}{2} (\varphi_\xi^2 + \varphi_\vartheta^2 - (1 + \eta_\xi^2 + \eta_\vartheta^2) \Phi_\zeta^2) - \eta - p, \quad (4)$$

$$64 \quad \Phi_{\xi\xi} + \Phi_{\vartheta\vartheta} + \Phi_{\zeta\zeta} = \Upsilon(\Phi), \quad (5)$$

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66 where  $\Upsilon$  is the operator:

$$67 \quad \Upsilon(\Phi) = 2\eta_\xi(\Phi)_{\xi\xi} + 2\eta_\vartheta(\Phi)_{\vartheta\vartheta} + (\eta_{\xi\xi} + \eta_{\vartheta\vartheta})(\Phi)_\zeta - (\eta_\xi^2 + \eta_\vartheta^2)(\Phi)_{\zeta\zeta}, \quad (6)$$

68 capital fonts  $\Phi$  are used for domain  $\zeta < 0$  while the lower case  $\varphi$  refers to  $\zeta = 0$ .

69 It is suggested in (Chalikov et al., 2014) that it is convenient to represent velocity  
 70 potential  $\varphi$  as a sum of two components, i.e., an analytical ('linear') component

71  $\bar{\Phi}$ , ( $\bar{\varphi} = \bar{\Phi}(\xi, \vartheta, 0)$ ) and an arbitrary ('non-linear') component  $\tilde{F}$ , ( $\tilde{j} = \tilde{F}(x, y, 0)$ ):

$$72 \quad j = \bar{j} + \tilde{j}, \quad F = \bar{F} + \tilde{F}. \quad (7)$$

73 The analytical component  $\bar{\Phi}$  satisfies Laplace equation:

$$74 \quad \bar{\Phi}_{\xi\xi} + \bar{\Phi}_{\vartheta\vartheta} + \bar{\Phi}_{\zeta\zeta} = 0, \quad (8)$$

75 with known solution:

$$76 \quad \bar{\Phi}(\xi, \vartheta, \zeta) = \sum_{k,l} \bar{\varphi}_{k,l} \exp(|k|\zeta) \Theta_{k,l}, \quad (9)$$

77 ( $\bar{\varphi}_{k,l}$  are Fourier coefficients of surface analytical potential  $\bar{\varphi}$  at  $\zeta = 0$ ). The solution satisfies

78 boundary conditions:

$$79 \quad \begin{aligned} \zeta = 0: \quad & \bar{\Phi} = \bar{\varphi} \\ \zeta \rightarrow -\infty: \quad & \tilde{\Phi}_\zeta \rightarrow 0 \end{aligned} \quad (10)$$

80 The nonlinear component satisfies an equation:



$$81 \quad \tilde{\Phi}_{\xi\xi} + \tilde{\Phi}_{\eta\eta} + \tilde{\Phi}_{\zeta\zeta} = \Upsilon(\tilde{\Phi}) + \Upsilon(\bar{\Phi}), \quad (11)$$

82 Eq. (11) is solved with the boundary conditions:

$$83 \quad \begin{aligned} \zeta = 0: \quad \tilde{\Phi} &= 0 \\ \zeta \rightarrow -\infty: \quad \tilde{\Phi}_{\zeta} &\rightarrow 0 \end{aligned} \quad (12)$$

84 The derivatives of linear component  $\bar{\Phi}$  in (6) are calculated analytically. The scheme  
 85 combines 2-D Fourier transform method in the ‘horizontal surfaces’ and a second-order finite-  
 86 difference approximation on a stretched staggered grid defined by relation  $\Delta\zeta_{j+1} = \chi\Delta\zeta_j$  ( $\Delta\zeta$  is  
 87 a vertical step, while  $j = 1$  at the surface). The stretched grid provides increase of accuracy of  
 88 approximation for the exponentially decaying modes. The values of stretching coefficient  $\chi$  lie  
 89 within the interval 1.01-1.20. A finite-difference second-order approximation of Eq. (10) on a  
 90 non-uniform vertical grid is quite straightforward. Equation (11) is solved as Poisson equations  
 91 with iterations over the right-hand side. A detailed description of the scheme and its validation is  
 92 given in (Chalikov, 2016).

93 Equations (3) – (5) are written in a non-dimensional form by using the following scales:  
 94 length  $L$  where  $2\pi L$  is (dimensional) period in the horizontal direction; time  $L^{1/2}g^{-1/2}$  and velocity  
 95 potential  $L^{3/2}g^{1/2}$  ( $g$  is acceleration of gravity). The pressure is normalized by water density, so  
 96 that the pressure scale is  $Lg$ . Equations (3) – (5) are self-similar to the transformation with  
 97 respect to  $L$ . All the results presented in this paper are nondimensional.

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### 3. Energy input and dissipation

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101 Input energy to waves describes a pressure term  $p$  in a dynamic boundary condition (4).  
 102 Dissipation cannot be described with use of potential equations, but for realistic description of  
 103 wave dynamics, dissipation of wave energy should be taken into account, i.e., we should include  
 104 in equations (3) and (4) additional terms which, strictly speaking, contradict the assumption of  
 105 potentiality.

#### 3.1 Energy input from wind

107 According to the linear theory (Miles, 1957), the Fourier components of surface pressure  
 108  $p$  are connected with those of surface elevation through the following expression:

$$109 \quad p_{k,l} + ip_{-k,-l} = \frac{\rho_a}{\rho_w} (\beta_{k,l} + i\beta_{-k,-l}) (h_{k,l} + ih_{-k,-l}), \quad (13)$$

110 where  $h_{k,l}, h_{-k,-l}, \beta_{k,l}, \beta_{-k,-l}$ , are real and imaginary parts of elevation  $\eta$  and the so-called  $\beta$ -  
 111 function (i.e., Fourier coefficients at COS and SIN, respectively);  $\rho_a / \rho_w$  is a ratio of air and  
 112 water densities, respectively. Hence, for derivation of shape of beta-function it is necessary to  
 113 simultaneously measure wave surface elevation and non-static pressure on the surface.  
 114 Experimental measurement of surface pressure is a very difficult problem since the  
 115 measurements should be done very close to a moving surface, preferably, with a surface-  
 116 following sensor. Such measurements are done quite seldom, especially, in the field. The  
 117 measurements were carried out for the first time by a team of authors both in laboratory and field  
 118 (Snyder et al, 1981; Hsiao and Shemdin, 1983; Hasselmann and Bösenberg, 1991; Donelan et al.,  
 119 2005, 2006). The data obtained in this way allowed constructing an imaginary part of beta-



120 function used in some versions of wave forecasting models (Rogers et al. 2012). The second way  
121 of beta-function evaluation is based on the results of numerical investigations of statistical  
122 structure of the boundary layer above waves with use of Reynolds equations and an appropriate  
123 closure scheme. In general, this method works so well that many problems in the technical fluid  
124 mechanics are often solved using numerical models, not experimentally. This method was being  
125 developed beginning from (Chalikov, 1978, 1986), followed by (Chalikov and Makin, 1991;  
126 Chalikov and Belevich, 1992; Chalikov, 1995). The results were implemented in  
127 WAVEWATCH model, i.e., a third-generation wave forecast model (Tolman and Chalikov,  
128 1996) and thoroughly validated against the experimental data in the course of developing  
129 WAVEWATCH-III (Tolman et al., 2014). This method was later improved on the basis of more  
130 advanced coupled modeling of waves and boundary layer (Chalikov and Rainchik, 2010;  
131 hereafter CR), while the beta-function used in WAVEWATCH-III was corrected and extended  
132 up to high frequencies. Direct calculation of energy input to waves requires both real and  
133 imaginary parts of the beta-function. The total energy input to waves depends on imaginary part  
134 of  $\beta$ -function, while the moments of higher order depend both on imaginary and real parts of  $\beta$   
135 . This is why full approximation constructed in CR was used in the current work. Note that in the  
136 range of relatively low frequencies the new method is very close to the scheme implemented in  
137 WAVEWATCH-III.

138 It is a traditional suggestion that both coefficients are the functions of virtual  
139 nondimensional frequency  $\Omega = \omega_k U \cos \psi = U / c_k \cos \psi$  (where  $\omega_k$  and  $U$  are the  
140 nondimensional radian frequency and wind speed, respectively;  $c_k$  is a phase speed of the  $k^{\text{th}}$   
141 mode;  $\psi$  is an angle between wind and wave mode directions). Most of the schemes for  
142 calculations of  $\beta$ -function consider a relatively narrow interval of nondimensional frequencies  
143  $\Omega$ . In the current work, the range of frequencies covers an interval  $(0 < \Omega_p < 10)$ , and  
144 occasionally the values of  $\Omega > 10$  can appear. This is why the function derived in (Chalikov and  
145 Rainchik, 2010) through coupled simulations of waves and boundary layer, is used here. Wave  
146 model is based on potential equations for a flow with free surface, extended with an algorithm  
147 for breaking dissipation (see below description of the breaking dissipation parameterization).  
148 Wave boundary layer (WBL) model is based on Reynolds equations closed with  $K - \varepsilon$  scheme;  
149 solutions for air and water are matched through the interface. The  $\beta$ -function obtained in CR  
150 was used for evaluation of accuracy of the surface pressure  $p$  calculations. A shape of  $\beta$ -  
151 function connecting surface elevations and surface pressure, is studied up to high  
152 nondimensional wave frequencies both in positive and negative (i.e., for wind opposite to waves)  
153 domains. The data on  $\beta$ -function exhibit wide scatter, but since the volume of data was quite  
154 large (47 long-term numerical runs allowed us to generate about 1,400,000 values of  $\beta$ ), the  
155 shape of  $\beta$ -function was defined with satisfactory accuracy up to very high nondimensional  
156 frequencies  $(-50 < \Omega < 50)$ . As a result, the data on  $\beta$ -function in such a broad range, allow us  
157 to calculate wave drag up to very high frequencies and to explicitly divide the fluxes of energy  
158 and momentum transferred by the pressure and molecular viscosity. This method is free of  
159 arbitrary assumptions on the drag coefficient  $C_d$ ; and, on the contrary, such calculations allow  
160 investigating the nature of wave drag (see Ting et al., 2012)

161 It was indicated above that an initial wave field is assigned as superposition of linear  
162 modes which amplitudes are calculated with JONSWAP spectrum with peak wave number



163  $k_p = 100$ . The initial value  $\Omega_0 = 6$  was chosen, i.e., a ratio of the nondimensional wind speed at  
 164 height  $\lambda_0 / 2 = 2\pi / 100$  and the phase speed  $c_0 = k_0^{-1/2}$  is equal to 6. Such a high ratio corresponds  
 165 to initial stages of wave development. The values of  $\Omega$  for other wave numbers are calculated  
 166 by assuming that wind profile is logarithmic:

$$167 \quad \Omega_k = \Omega_0 \frac{c_0}{c_k} \ln \frac{\lambda_k}{2z_0} \left( \ln \frac{\lambda_0}{2z_{00}} \right)^{-1} \cos \psi, \quad (14)$$

168 where  $z_{00}$  is effective nondimensional roughness for the initial wind profile, while  $z_0$  is the  
 169 actual roughness parameter that depends on the energy in a high-frequency part of spectrum and  
 170 on the wind profile. We call it ‘effective’, since very close to the surface the wind profile is not  
 171 logarithmic (Chalikov, 1995; Tolman, Chalikov, CR). The value of this parameter depends on  
 172 the wind velocity and energy in a high-wave number interval of wave spectrum, as well as on the  
 173 length scale of the problem. All these effects are possible to include by matching the wave model  
 174 with a one dimensional WBL model (Ting et al, 2012). Here, a simplified scheme for the  
 175 roughness parameter is chosen. It is well known that the roughness parameter (as well as a drag  
 176 coefficient) decreases with decrease of the inverse wave age. In our case wind speed is fixed, and  
 177 dependence for the nondimensional roughness parameter is constructed on the basis of the results  
 178 obtained in CR:

$$179 \quad z_0 = 15z_{00}\Omega, \quad (15)$$

180 where  $z_{00} = 10^{-3}$  is the initial value of the roughness parameter. Eq. (15) approximates  
 181 dependence of the effective roughness at the stage of wave development. Note that the results are  
 182 not sensitive to variation of the roughness parameter within reasonable limits.

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### 184 **3.2 High wave number energy dissipation**

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186 A nonlinear flux of energy directed to the small wave numbers produces downshifting of  
 187 spectrum, while an opposite flux forms a shape of spectral tail. The second process can produce  
 188 accumulation of energy near ‘cut’ wave number. Both processes become more intensive with  
 189 increase of energy input. Growth of amplitudes at high wave numbers is followed by the growth  
 190 of local steepness and numerical instability. This phenomenon well known in numerical fluid  
 191 mechanics is eliminated by use of a highly selective filter simulating nonlinear viscosity. To  
 192 support stability, additional terms are included into the right hand sides of equations (3) and (4):

$$193 \quad \frac{\partial \eta_{k,l}}{\partial \tau} = E_{k,l} - \mu_{k,l} \eta_{k,l}, \quad (16)$$

$$194 \quad \frac{\partial \varphi_{k,l}}{\partial \tau} = F_{k,l} - \mu_{k,l} \varphi_{k,l} \quad (17)$$

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196 ( $E_{k,l}$  and  $F_{k,l}$  are Fourier amplitudes of the right-hand sides of equations (3) and (4) while factor  
 197  $\mu_{k,l}$  is calculated using a formula:



$$\mu_{k,l} = \begin{cases} 0 & |k| < k_d \\ c_m k_0 \left( \frac{|k| - k_d}{(k_0 - k_d)} \right)^2 & k_d \leq |k| \leq k_0 \\ c_m k_0 & |k| > k_0 \end{cases} \quad (18)$$

where  $k$  and  $l$  are components of wave number  $|k|$ , while coefficients  $k_d$  and  $k_0$  are defined by the expressions:

$$k_d = d_m^2 M_x M_y \left( (l|k|^{-1} d_m M_x)^2 + (k|k|^{-1} d_m M_y)^2 \right)^{-1/2} \quad (19)$$

$$k_0 = M_x M_y \left( (l|k|^{-1} M_x)^2 + (k|k|^{-1} M_y)^2 \right)^{-1/2} \quad (20)$$

where  $c_m = 0.1$ ,  $d_m = 0.75$ . Expressions (18) - (20) can be interpreted in a straightforward way: the value of  $\mu_{k,l}$  is equal to zero inside the ellipse with semi-axes  $d_m M_x$  and  $d_m M_y$ ; then it grows linearly with  $|k|$  up to the value  $c_m$  and is equal to  $c_m$  outside the outer ellipse. This method of filtration that we call ‘tail dissipation’ was developed and validated with a conformal model by Chalikov and Sheinin (1998). The sensitivity of the results to the parameters in (18) - (20) is not high. The aim of the algorithm is support of smoothness and monotonicity of wave spectrum within a high wave number range. Since the algorithm affects amplitudes of small modes, it actually does not reduce the total energy, though it efficiently prevents development of numerical instability. Note that any long-term calculations cannot be performed without ‘tail dissipation’ eliminating development of the numerical instability at high wave numbers.

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### 3.3 Dissipation due to wave breaking

The main process of wave dissipation is wave breaking. This process is taken into account in all spectral wave forecasting models similar to WAVEWATCH (see Tolman and Chalikov, 1996). Since there are no waves in spectral models, no local criteria of wave breaking can be formulated. This is why breaking dissipation is represented in spectral models in a distorted form. Areal breaking occurs in relatively narrow areas of physical space; however, spectral image of such breaking is stretched over the entire wave spectrum, while in reality the breaking decreases height and energy of dominant waves. This contradiction occurs because waves in spectral models are assumed as linear ones, while in fact the breaking occurs in physical space with nonlinear sharp wave, usually composed of several modes.

The mechanics of wave breaking at developed wave spectrum differs from that in a wave field represented by few modes, normally considered in many theoretical and laboratory investigations. Since the breaking in laboratory conditions is initiated by special assigning of amplitudes and phases, it cannot be similar to the breaking in natural conditions. To some degree, the wave breaking is similar to development of extreme wave that appears suddenly with no pronounced prehistory (Chalikov and Babanin, 2016a, 2016b). There are no signs of modulational instability in both phenomena, which suggests a process of taking energy from other modes. The evolution leading to breaking or ‘freaking’ seems just opposite: full energy of main wave remains nearly constant while the columnar energy is focusing around the crest of this wave which becomes sharper and unstable. Probably, even more frequent cases of wave

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235 breaking and extreme wave appearance can be explained by local superposition of several  
236 modes.

237 The instability of interface leading to breaking is an important and poorly developed  
238 problem of fluid mechanics. In general, this essentially nonlinear process should be investigated  
239 for a two-phase flow. Such approach was demonstrated, for example, by Iafrati (2001).  
240 However, the progress in solving this highly complicated problem is not too fast.

241 The problem of breaking parameterization includes two points: (1) establishing of a  
242 criterion of breaking onset and (2) developing of an algorithm of breaking parameterization. The  
243 problem of breaking is discussed in details in Babanin (2011). Chalikov and Babanin (2012)  
244 performed numerical investigation of the processes leading to breaking. It was found that a clear  
245 predictor of breaking, formulated in dynamical and geometrical terms, probably does not exist.  
246 The most evident criterion of breaking is the breaking itself, i.e., the process when some part of  
247 upper portion of sharp wave crest is falling down. This process is usually followed by separation  
248 of detached volume of liquid into water and air phases. Unfortunately, there is no possibility to  
249 describe this process within the scope of potential theory.

250 Some investigators suggest using the physical velocity approaching the rate of surface  
251 movement in the same direction as a criterion of breaking onset. This is incorrect, since the  
252 kinematic boundary condition suggests that these quantities are exactly equal to each other. It is  
253 quite clear that the onset of breaking can be characterized by appearance of non-single-value  
254 piece of surface. This stage can be investigated with two-dimensional model which due to a high  
255 flexibility of the conformal coordinates allows us to reproduce a surface with the inclination in  
256 the Cartesian coordinates larger than 90 degrees. (In the conformal coordinates the dependence  
257 of elevation on curvilinear coordinate is always single-value). The duration of this stage is  
258 extremely short, the calculations being always interrupted by the numerical instability with sharp  
259 violation of conservations laws (constant integral invariants, i.e., full energy and volume) and  
260 strong distortion of the local structure of flow. Numerous numerical experiments with conformal  
261 model showed that after appearance of non-single value, the model never returns to stability.  
262 However, introducing of appearance of the non-single-surface as a criterion of breaking  
263 instability even in conformal model is impossible, since a behavior of model at a critical point is  
264 unpredictable, and the run is most likely to be terminated, no matter what kind of  
265 parameterization of breaking is introduced. It means that even in a very precise conformal  
266 model, stabilization of solution should be initiated prior to breaking.

267 Consideration of exact criterion for breaking onset for the models using transformation of  
268 the coordinate type (1) is useless, since the numerical instability in such models arises not  
269 because of the breaking approaching but because of appearance of large local steepness. Multiple  
270 experiments with direct 3-D wave model show that appearance of local steepness

271  $\max\left(\frac{\partial\eta}{\partial x}, \frac{\partial\eta}{\partial y}\right)$  exceeding  $\approx 2$  (that corresponds to a slope of about 60 degrees) is always

272 followed by numerical instability. Decrease of time step does not make any effect. As seen, a  
273 surface with such slope is very far from being a vertical 'wall', when real breaking starts.  
274 However, an algorithm for breaking parameterization must prevent appearance of large local  
275 steepness. The situation is similar to the numerical modeling of turbulence (LES technique),  
276 where the local highly selective viscosity is used to prevent appearance of too large local  
277 gradients of velocity. The description of breaking in direct wave modeling should satisfy the  
278 following conditions. (1) It should prevent large local gradients of elevation; in our case the



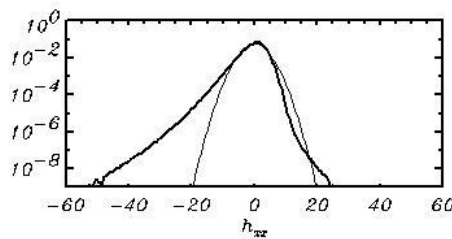
279 breaking algorithm should prevent the onset of instability at each point of half million of grid  
 280 points over more than 100 thousand of time steps.(2) It should describe in a more or less realistic  
 281 way the loss of kinetic and potential energies with preservation of balance between them. (3) It  
 282 should preserve the volume. It was suggested in (Chalikov, 2005) that an acceptable scheme can  
 283 be based on the local highly selective diffusion operator with special diffusion coefficient.  
 284 Several schemes of such type were validated, and finally the following scheme was chosen:

$$285 \quad \eta_\tau = E_\eta + J^{-1} \left( \frac{\partial}{\partial \xi} B_\xi \frac{\partial \eta}{\partial \xi} + \frac{\partial}{\partial \vartheta} B_\vartheta \frac{\partial \eta}{\partial \vartheta} \right), \quad (21)$$

$$286 \quad \varphi_\tau = F_\varphi + J^{-1} \left( \frac{\partial}{\partial \xi} B_\xi \frac{\partial \varphi}{\partial \xi} + \frac{\partial}{\partial \vartheta} B_\vartheta \frac{\partial \varphi}{\partial \vartheta} \right), \quad (22)$$

287 where  $F_\eta$  and  $F_\varphi$  are the right-hand sides of equations (3) and (4) including the terms introduced  
 288 by (16) – (20). It was suggested in the first versions of the scheme that diffusion coefficient  
 289 depends on a local slope, however, such scheme did not prove to be very reliable since it did not  
 290 prevent all of the events of numerical instability. A scheme based on the calculation of the local  
 291 curvilinearity  $\eta_{\xi\xi}$  and  $\eta_{\vartheta\vartheta}$  turned out to be a lot more reliable. The calculations of 75 different  
 292 runs were performed with full 3-D model in (Chalikov et al, 2014) over period of  $t = 350$  (70,000  
 293 time steps). The total number of values used for the calculations of dependence in Fig. 1 (thick  
 294 curve) is about 6 billion. The normal probability calculated with the same dispersion is shown by  
 295 thin curve.

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**Figure 1.** Probability of curvilinearity  $\eta_{\xi\xi}$ .  
 Thick curve calculated with full 3-D model;  
 thin curve is a probability calculated over  
 ensemble of linear modes with the same  
 spectrum.

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305 It is seen that the probability of large negative values of curvilinearity is by orders larger than the  
 306 probability calculated over ensemble of linear modes with the same spectrum.

307 The curvilinearity turned out to be very sensitive to the shape of surface. This is why it  
 308 was chosen as a criterion of breaking approach. Coefficients  $B_\xi$  and  $B_\vartheta$  depend nonlinearly on  
 309 the curvilinearity

$$310 \quad B_\xi = \begin{cases} \Delta \xi C_B \eta_{\xi\xi}^2 & \eta_{\xi\xi} < \eta_{\xi\xi}^{cr} \\ 0 & \eta_{\xi\xi} \geq \eta_{\xi\xi}^{cr} \end{cases} \quad (23)$$

$$311 \quad B_\vartheta = \begin{cases} \Delta \vartheta C_B \eta_{\vartheta\vartheta}^2 & \eta_{\vartheta\vartheta} < \eta_{\vartheta\vartheta}^{cr} \\ 0 & \eta_{\vartheta\vartheta} \geq \eta_{\vartheta\vartheta}^{cr} \end{cases} \quad (24)$$

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313 where  $\Delta \xi$  and  $\Delta \vartheta$  are horizontal steps in  $x$  and  $y$  direction in grid space, and coefficients are  
 314  $C_B = 2.0$ ,  $\eta_{\xi\xi}^{cr} = \eta_{\vartheta\vartheta}^{cr} = -50$ . Algorithm (21) - (24) does not change the volume and decreases the  
 315 local potential and kinetic energy. It is assumed that the lost momentum and energy are





316 transferred to current and turbulence (see Chalikov and Belevich, 1992). Besides, the energy also  
 317 goes to other wave modes. The choice of parameters in (21) - (24) is based on simple  
 318 considerations: local piece of surface can closely approach the critical curvilinearity but not  
 319 exceed it. The values of the coefficients are picked with reserve to provide stability of long runs.

320 We do not think that the suggested breaking parameterization is a final solution of the  
 321 problem. Other schemes will be tried in the next version of the model. However, the results  
 322 presented below show that the scheme is reliable and provides a realistic energy dissipation rate.

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#### 324 4. Calculations and results

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326 The elevation and surface velocity potential fields are approximated in the current  
 327 calculations by  $M_x = 256$  and  $M_y = 128$  modes in directions  $x$  and  $y$ . The corresponding grid  
 328 includes  $N_x \times N_y = (1024 \times 512)$  knots. The vertical derivatives are approximated at vertical  
 329 stretched grid  $d\zeta_{j+1} = \nu d\zeta_j$ , ( $j = 1, 2, 3, \dots, L_w$ ) where  $\nu = 1.2$  and  $L_w = 10$ . The small number of  
 330 levels used for solution of the equation for nonlinear component of the velocity potential is  
 331 possible because just a surface vertical derivative for the velocity potential  $\partial\Phi / \partial\zeta (\zeta = 0)$  is  
 332 required. The velocity potential mainly consists of an analytical component  $\bar{\varphi}$ , while a nonlinear  
 333 component provides but small correction. To reach an accuracy of solution  $\varepsilon = 10^{-6}$  for equation  
 334 (11), no more than two iterations were usually sufficient.

335 The parameters chosen were used for solution of the problem of wave field evolution  
 336 over acceptable time (of the order of 10 days). The initial conditions were assigned on the basis  
 337 of empirical spectrum JONSWAP (Hasselmann et al, 1973) with a maximum placed at wave  
 338 number  $k_p = 100$  with angle spreading  $(\cosh \psi)^{256}$ . Details of initial conditions are of no  
 339 importance because an initial energy level is quite low.

340 The total energy of wave motion  $E = E_p + E_k$  ( $E_p$  - is potential energy, while  $E_k$  is  
 341 kinetic energy) is calculated with the following formulas:

342

$$343 \quad E_p = 0.25 \overline{\eta^2}, \quad E_k = 0.5 \overline{(\varphi_x^2 + \varphi_y^2 + \varphi_z^2)}, \quad (25)$$

344

345 where single bar denotes averaging over the  $\xi$  and  $\vartheta$  coordinates, while double bar denotes  
 346 averaging over entire volume. The derivatives in (25) are calculated according to transformation  
 347 (1). An equation of integral energy  $E = E_p + E_k$  evolution can be represented in the following  
 348 form:

$$349 \quad \frac{dE}{dt} = \overline{I} + \overline{D_b} + \overline{D_t} + \overline{N}, \quad (26)$$

350 where  $\overline{I}$  is the integral input of energy from wind (Eqs. (13) - (15));  $\overline{D_b}$  is a rate of energy  
 351 dissipation due to the wave breaking (Eqs. (21) - (24));  $\overline{D_t}$  is a rate of energy dissipation due  
 352 to33filtration of high-wave number modes ('tail dissipation', Eqs. (16) - (20));  $\overline{N}$  is an integral  
 353 effect of the nonlinear interactions described by the right-hand side of the equations when  
 354 surface pressure  $p$  is equal to zero. The differential form for calculation of the energy



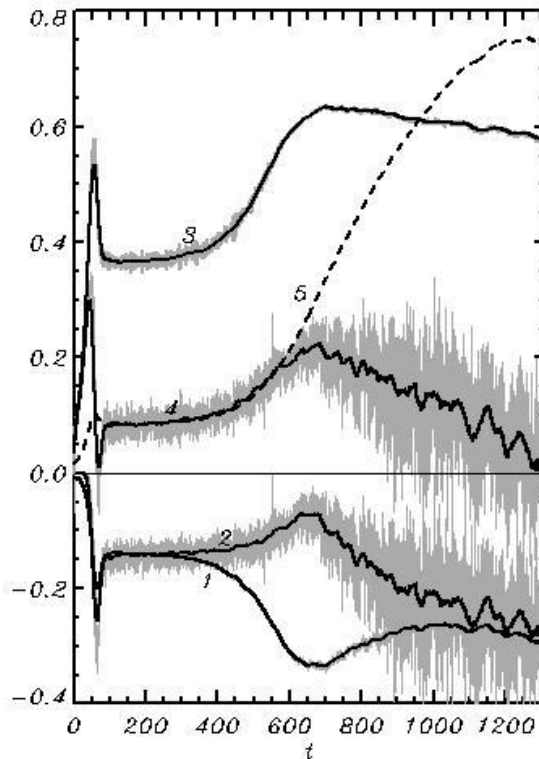
355 transformation can be, in principle, derived from Eqs. (3) – (5), but here a more convenient and  
 356 simple method was applied. Different rates of integral energy transformations can be calculated  
 357 with help of fictitious time steps (i.e., apart from the basic calculations). For example, the value  
 358 of  $\overline{I}$  is calculated by the following relation:

$$360 \quad \overline{I} = \frac{1}{\Delta t} (\overline{E^{t+\Delta t}} - \overline{E^t}), \quad (27)$$

361  
 362 where  $\overline{E^{t+\Delta t}}$  is the integral energy of wave field obtained after one time step with the right side of  
 363 equation (4) containing only the surface pressure calculated with Eqs. (13) – (15). For  
 364 calculation of the dissipation rate due to filtration, the right-hand side of the equations contains  
 365 just the terms introduced in Eqs. (16), (17), while for calculation of the effects of breaking, only  
 366 the terms introduced in (21) – (22) are in use.

367 An evolution of the characteristics calculated by formula (27) is shown in Fig. 2.

368



**Figure 2.** Evolution of integral characteristics of solution, rate of evolution of integral energy multiplied by  $10^7$  due to: 1 – tail dissipation  $D_t$  (Eqs. 16-20); 2 – breaking dissipation  $D_b$  (Eqs. 21-24); 3 – input of energy from wind  $I$  (Eqs. 13-15); 4 – balance of energy  $I + D_t + D_b$ . Curve 5 shows the evolution of wave energy  $10^5 E$ . Vertical bars of grey color show the instantaneous values; thick curve shows the smoothed behavior.

393

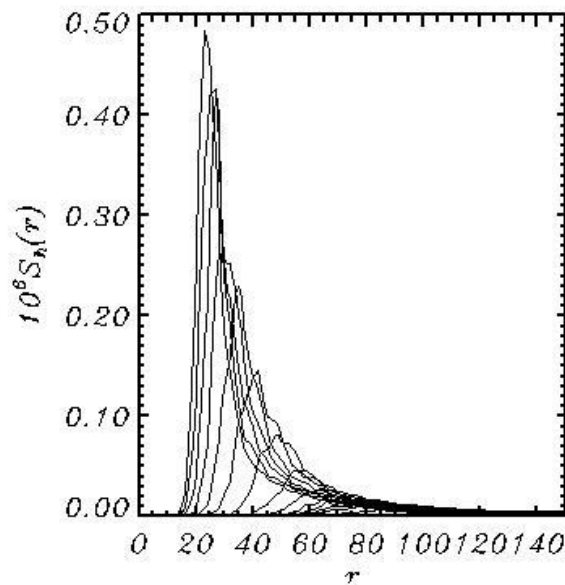
394 Up to the end of integration, the sum of all energy transition terms (tail dissipation  $\overline{D_t}$ , breaking  
 395 dissipation  $\overline{D_b}$  and energy input  $\overline{I}$ ) is approaching zero (curve 4), and the energy growth  $E$   
 396 (Curve 5) stops. Then the energy tends to decrease, but we are not sure about the nature of this  
 397 effect. Such behavior can be explained by a fluctuating character of mutual adjustment of input



398 and dissipation or simply by worsening of the approximation because of the downshifting  
 399 process. Note that opposite to a more or less monotonic behavior of tail dissipation (Curve 1),  
 400 the breaking dissipation is highly intermittent, which is consistent with common views on the  
 401 nature of wave breaking.

402 The data on evolution of wave spectrum are shown in Fig. 3.

403



**Figure 3.** The wave spectra  $S_h(r)$  integrated over angle  $\psi$  in the polar coordinates and averaged over consequent intervals of length about 100 units of nondimensional time  $t$ . The spectra are growing and shifting from right to left.

423

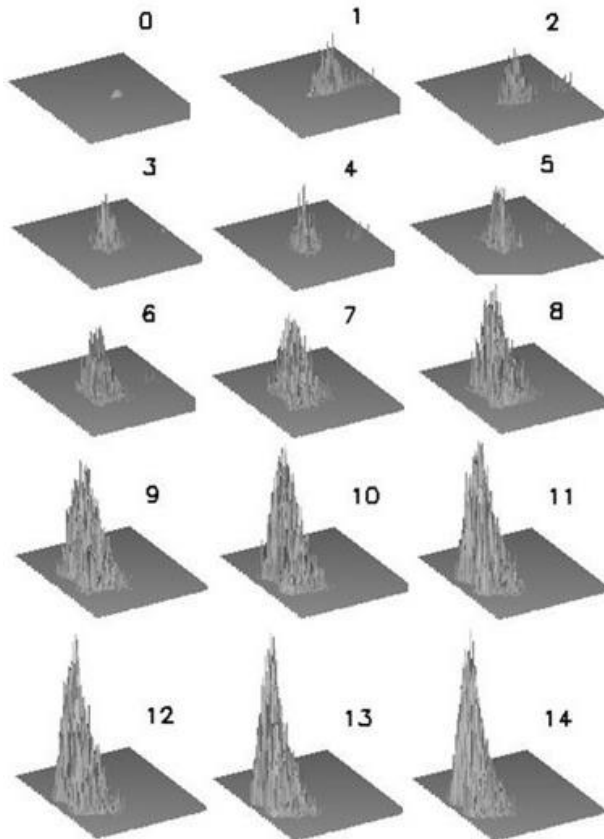
424 The 2-D wave spectrum  $S(k,l)$  ( $0 \leq k \leq M_x, -M_y \leq l \leq M_y$ ) averaged over 13 time  
 425 intervals of length equal to  $\Delta t \approx 100$ , was transferred to the polar coordinates  $S_p(\psi, r)$   
 426 ( $-\pi/2 \leq \psi \leq \pi/2, 0 \leq r \leq M_x$ ) and then averaged over angle  $\psi$  to obtain 1-D spectrum  $S_h(r)$ :

$$427 \quad S_h(r) = \sum S_p(\psi, r) r \Delta\psi. \quad (28)$$

428 An angle  $\psi = 0$  coincides with the direction of wind  $U$ ,  $\Delta\psi = \pi/180$ .

429 The wave spectra  $S_h(r)$  calculated by averaging over angle  $\psi$  in the polar coordinates  
 430 and averaged over consequent intervals of length about 100 units of nondimensional time  $t$  are  
 431 presented in Fig. 3. The spectra increase and move from high to low wave numbers, i.e., they  
 432 undergo downshifting. A maximum value of  $S_h(r)$  increases as much as 152 times. According  
 433 to the data in Fig. 2, the total energy increases 44 times. This difference is explained by the  
 434 spectrum narrowing and by the overlapping effect (i.e., decrease of high-frequency spectrum for  
 435 long fetches). The 3-D images of wave spectrum  $\log_{10}(S(k,l))$  are shown in Fig. 4.

436



**Figure 4** Sequence of 3-D images of  $\lg_{10}(S(k,l))$  where each panel corresponds to single curve in Fig. 3. The left side refers to wave number  $l(-M_y \leq l \leq M_y)$  and front side – to  $k(-M \leq k \leq M)$ .

464 As seen, each spectrum consists of separated peaks and holes<sup>1</sup>. This phenomenon was first  
 465 observed and discussed by Chalikov et al (2014). The repeated calculations with different  
 466 resolution showed that such structure of 2-D spectrum is typical. It cannot be explained by fixed  
 467 combination of interacting modes, since in different runs (with the same initial conditions but  
 468 different set of phases for the modes) peaks are located in different locations in Fourier space.

469 Another presentation is given in Fig 5 where the  $\log_{10}(S(\psi, r))$ , averaged over the  
 470 successive seven period length  $\Delta t = 200$ , is given. The first panel with a mark 0 refers to initial  
 471 conditions. Disturbances within the range  $(125 < k < 150)$  reflect initial adjustment of the input  
 472 and dissipation at high wave number slope of spectrum. The pictures characterize well the  
 473 downshifting and angle spreading of spectrum due to nonlinear interactions.

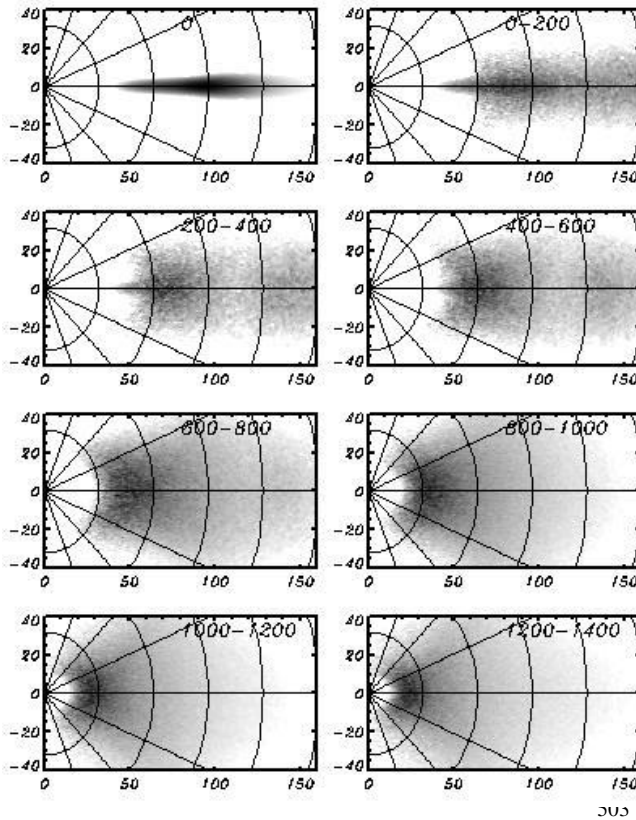
474 Evolution of the integrated over angles  $\psi$  wave spectrum  $S_h(r)$  can be described with  
 475 the equation  
 476

<sup>1</sup>The wave spectrum looks rather like La Sagrada Familia (Gaudi) in Barcelona than the St. Mary Axe ('Pickle') in London.



477 
$$\frac{dS_h(r)}{dt} = I(r) + D_t(r) + D_b(r) + N(r), \quad (29)$$

478



**Figure 5.** Sequence of 2-D images of  $\lg_{10}(S(k,l))$  averaged over consequent seven periods length  $\Delta t = 200$ . Numbers indicate the period of averaging (first panel marked 0, refers to initial conditions). Horizontal and vertical axes correspond to wave numbers  $k$  and  $l$  correspondingly

504 where  $I(r), D_t(r), D_b(r)$  and  $N(r)$  are the spectra of the input energy, tail dissipation, breaking  
 505 dissipation and a rate of nonlinear interactions, all obtained by integration over angles  $\psi$ . All of  
 506 the spectra shown below were obtained by transformation of 2-D spectra into the polar  
 507 coordinate  $(\psi, r)$  and then integrated over angles  $\psi$  within the interval  $(-\pi/2, \pi/2)$ . The  
 508 spectra can be calculated using an algorithm similar to the algorithm (27) for integral  
 509 characteristics. For example, the spectrum of energy input  $I(k, l)$  is calculated as follows:

510 
$$I(k, l) = (S_c^{t+\Delta t}(k, l) - S_c^t(k, l)) / \Delta t, \quad (30)$$

511 where  $S_c(k_x, k_y)$  is a spectrum of columnar energy calculated by relation

512 
$$S_c(k, l) = \frac{1}{2} \left( h_{k,l}^2 + h_{-k,-l}^2 + \int_{-H}^0 (u_{k,l}^2 + u_{-k,-l}^2 + v_{k,l}^2 + v_{-k,-l}^2 + w_{k,l}^2 + w_{-k,-l}^2) d\zeta \right) \quad (31)$$

513 where grid values of velocity components  $u, v, w$  are calculated by relations:

514 
$$u = \varphi_\xi + \varphi_\zeta \eta_\xi, \quad v = \varphi_\eta + \varphi_\zeta \eta_\eta, \quad w = \varphi_\zeta, \quad (32)$$

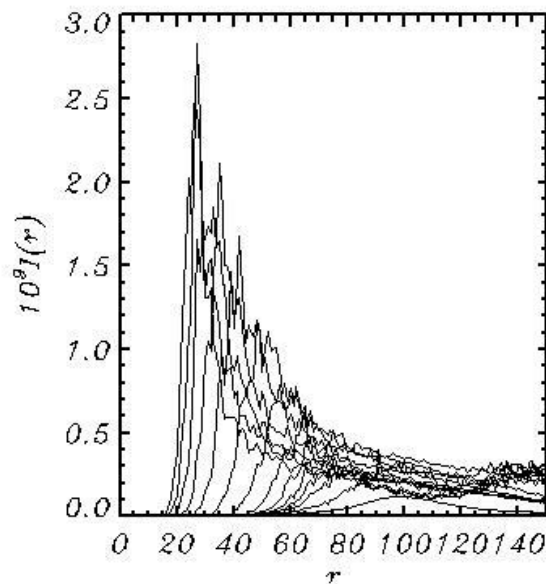
515 and  $u_{k,l}, v_{k,l}$  and  $w_{k,l}$  are their Fourier coefficients.



516 For calculation of  $I(k_x, k_y)$  the fictitious time steps  $\Delta t$  are made only with a term  
517 responsible for the energy input, i.e., surface pressure  $p$ . Spectrum  $I(k, l)$  was averaged over the  
518 periods  $\Delta t \approx 100$ , then transformed into a polar coordinate system and integrated in Fourier  
519 space over angles  $\psi$  within the interval  $(-\pi/2, \pi/2)$ .

520 The evolution of input spectra (Fig. 6) is, in general, similar to that of wave spectra  
521 shown in Fig. 3. Note that a maximum of spectra is located at the maximum of wave spectra  
522 since the input depends mainly on spectral density, while the dependence on frequency is less  
523 important.

524



**Figure 6.** The spectrum of energy input  $I(r)$  integrated over angle  $\psi$  in the polar coordinates and averaged over consequent intervals of length about 100 units of nondimensional time  $t$ .

544

545 Algorithm (29) – (32) was applied for calculation of the dissipation spectra due to  
546 dumping of a high-wave number part of spectrum (tail dissipation) and for calculation of the  
547 spectrum of breaking dissipation. In the first case, the fictitious time step was made taking into  
548 account the terms described by Eqs (16) – (20), while in the second case the time step was made  
549 using the terms described by Eqs (21) – (24).

550 The spectra of tail dissipation calculated similar to spectra  $I(r)$  are shown in Fig.  
551 7. Dissipation occurs at the periphery of spectrum, outside the ellipse with semi-axes  $d_m M_x$  and  
552  $d_m M_y$ <sup>2</sup>. This is why such dissipation, averaged over angles, seems to affect a middle part of 1-D  
553 spectrum. The tail dissipation effectively stabilizes the solution.

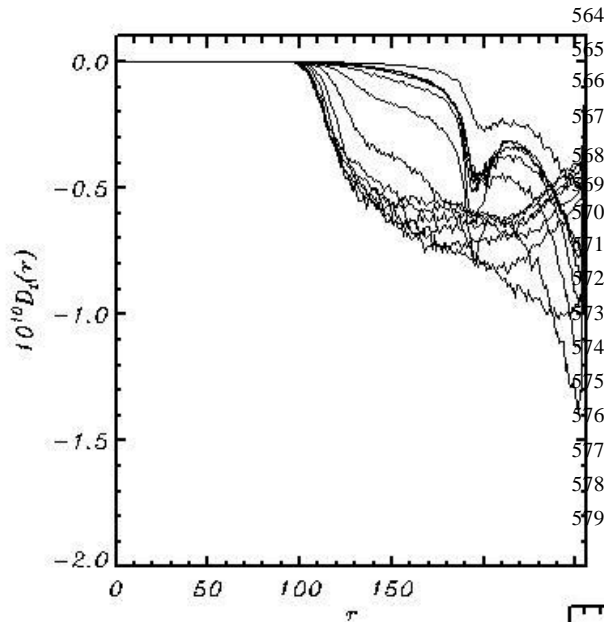
554 The breaking dissipation averaged over angles is presented in Fig. 8. As seen, the  
555 breaking dissipation has a maximum at spectral peak. It does not mean that in the vicinity of  
556 wave peak the probability of large curvilinearity is quite high. The high rate of breaking  
557 dissipation can be explained by high wave energy in the vicinity of wave peak. The energy lost

<sup>2</sup>The 2-D Fourier spectral ‘tail’ looks like ‘peacock’ tail.



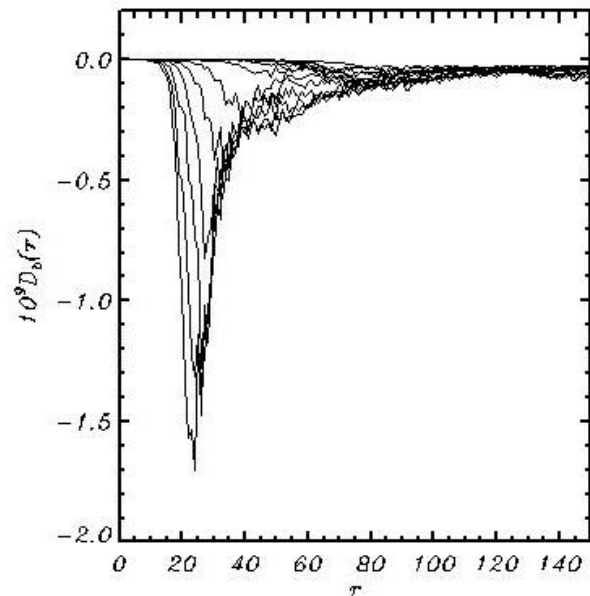
558 through breaking, described by the diffusion mechanism, correlates with the energy of breaking  
 559 waves. Opposite to high wave number dissipation which regulates shape of spectral tail, the  
 560 breaking dissipation forms the main energy-containing part of spectrum.

561 The diffusion mechanism suggested in (21), (22) modifies an elevation and surface  
 562 stream function in a close vicinity of breaking point. The amplitudes of side perturbation are  
 563 small and decrease very quickly over the distance from a breaking point.



564  
 565 **Figure 7.** Tail dissipation spectra  
 566  $D_t(r)$  integrated over angle  $\psi$  in  
 567 the polar coordinates and averaged over  
 568 consequent intervals of length about  
 569 100 units of nondimensional time  $t$ .  
 570  
 571  
 572  
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580  
 581  
 582 **Figure 8,** Breaking dissipation spectra  $D_b(r)$   
 583 integrated over angle  $\psi$  in the polar coordinates  
 584 and averaged over consequent intervals of length  
 585 about 100 units of nondimensional time  $t$ .  
 586  
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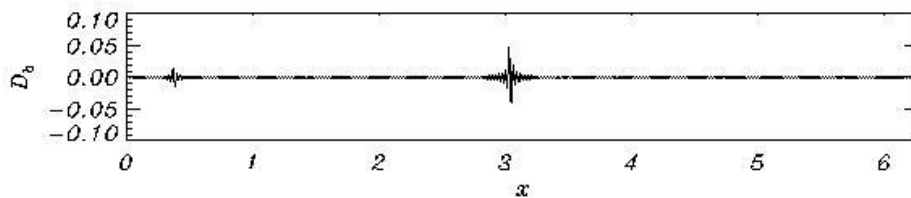


600 An example of profile of the energy input due to breaking  $D_b(x)$  is given in Fig. 9. As  
 601 seen, energy input is fluctuating around the breaking point. A diffusion operator chosen for  
 602 breaking parameterization not only decreases total energy but also redistributes the energy  
 603 between Fourier modes in Fourier space.

604 In general, for the specific conditions considered in the paper, the breaking is an  
 605 occasional process taking place in a small part of domain. The kurtosis of input energy due to the  
 606 breaking  $D_b(\xi, \mathcal{A})$ , i.e., the value

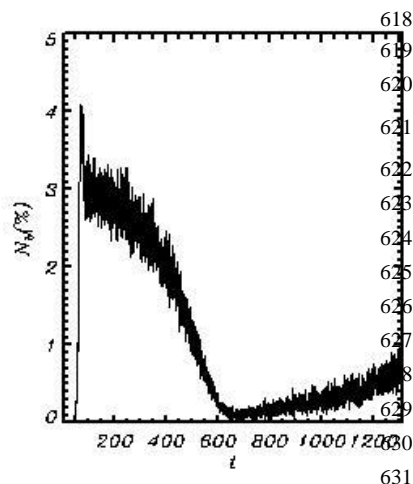
$$607 \quad Ku = \overline{D_b^4} \left( \overline{D_b^2} \right)^{-2} - 3 \quad (33)$$

608 is of the order of  $10^3$ , which corresponds to plain function with occasional separated peaks.  
 609



610  
 611 **Figure 9.** Example of energy input due to breaking  $D_b(x)$ .

612  
 613 The number of breaking points in terms of percentage of the total number of points is given in  
 614 Fig. 10. As seen, the number of breaking events is going down to  $t = 600$  and then is growing up  
 615 to the end of the calculations. The number of breaking events is not directly connected with  
 616 intensity of breaking, which is seen when comparing Fig. 10 and curve 2 in Fig.1.  
 617

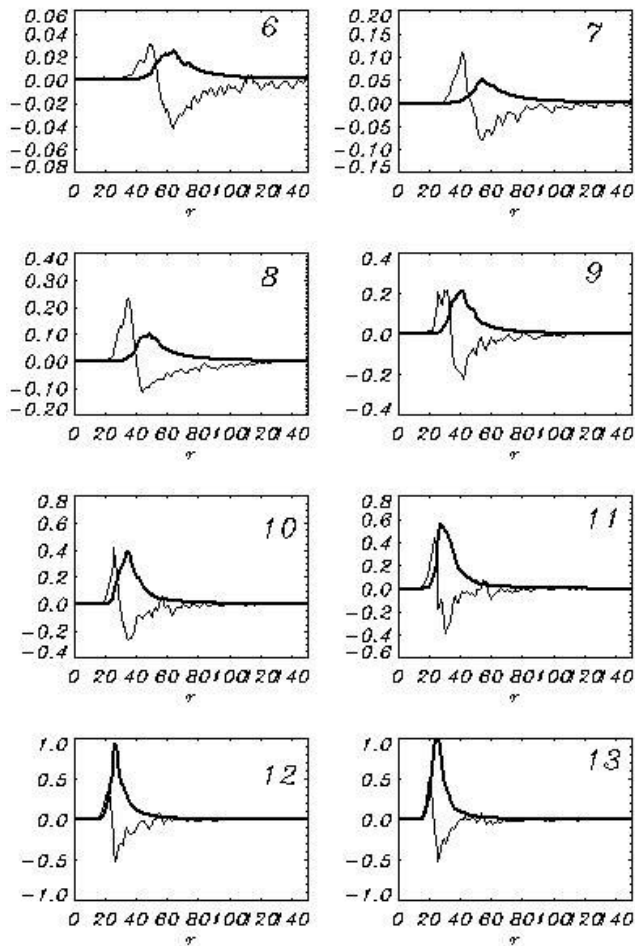


618  
 619 **Figure 10.** Evolution of number of wave breaking events  
 620  $N_b$  expressed in percentage of the number of grid points  
 621  $N_x \times N_y$ .

632 An integral term describing nonlinear interaction  $\overline{N}$  in Eq. (26) is small, but the magnitude of  
 633 spectrum  $N(r)$  is comparable with input  $I(r)$  and dissipation  $D_i(r)$  and  $D_b(r)$  terms. The  
 634 presentation of term  $N(r)$  in a form shown in Figs. (6) – (8) is not clear. This is why the spectra



635  $10^8 N(r)$  averaged over interval  $\Delta t = 100$  are plotted separately in Fig. 11 for the last eight  
 636 intervals (thick curves) together with the wave spectrum  $10^6 S_h(r)$ . In general, the shapes of  
 637 spectrum  $N(r)$  agree with the conclusions of the quasi-linear Hasselmann (1962) theory. At low  
 638 wave number slope of spectrum the nonlinear influx of energy is positive while at the opposite  
 639 slope it is negative. This process produces shifting of spectrum to the lower wave number  
 640 (downshifting). Opposite to the Hasselmann's theory, these results are obtained by solution of  
 641 full three-dimensional equations. It would be interesting to compare our results with the  
 642 calculations of Hasselmann's integral. Unfortunately, neither of the existing programs of such  
 643 type permits  
 644



**Figure 11.** Sequence of wave spectra  $S_h(r)$  (thick curves) and nonlinear input term  $N(r)$  (thin curves) averaged over consequent eight periods of length  $\Delta t = 100$  starting from 6<sup>th</sup> period.

673

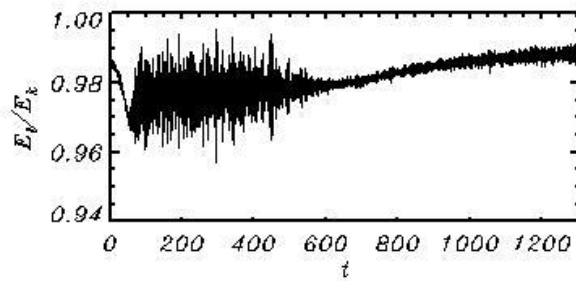
674 doing calculations with such a high resolution that was used in the current model. Note that  
 675 nonlinear interactions also produce widening of spectrum.

676 Obviously, the nonlinearity is quite an important property of surface waves. The  
 677 contribution of nonlinearity can be estimated, for example, by comparison of the kinetic energy



678 of linear component  $E_l = 0.5(\overline{\varphi_x^2} + \overline{\varphi_y^2} + \overline{\varphi_z^2})$  and the total kinetic energy  $E_k$  (Fig. 12). A ratio  
 679  $E_l / E_k$  as a function of time remains very close to 1, which proves that the nonlinear part of  
 680 energy makes up just a small percentage of the total energy. It does not mean that the role of  
 681 nonlinearity is small; its influence can manifest itself over large time scales.

682

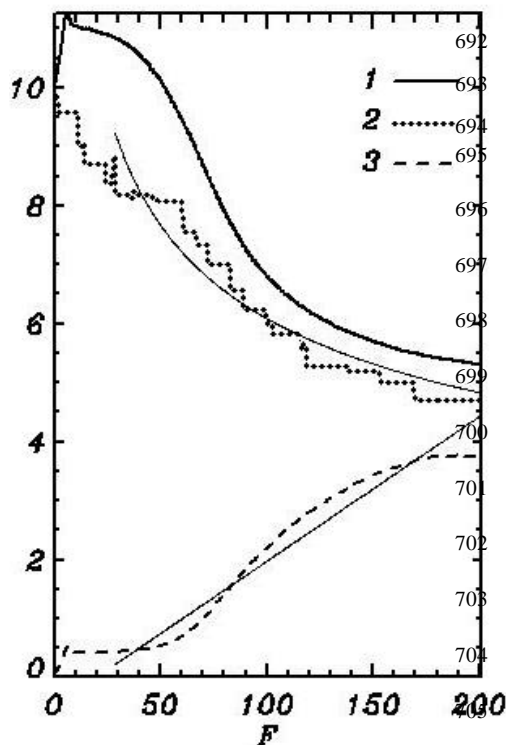


**Figure 12.** Time evolution of ratio  $E_l / E_k$ .

689

690 The time evolution of integral spectral characteristics is presented in Fig. 13.

691



**Figure 13.** Time evolution of: weighted frequency  $\omega_w$  (1) (Eq. 34); spectral peak frequency  $\omega_p$  (2); full energy  $E$  (3) (Eq. 25). Thin curves are empirical 97distance passed by the spectral peak.

706

707 Curve 1 corresponds to the weighted frequency  $\omega_w$ ,



$$\omega_p = \left( \frac{\int k S dk dl}{\int S dk dl} \right)^{1/2}, \quad (34)$$

708 where integrals are taken over the entire Fourier-domain. The value  $k_w$  is not sensitive to the  
 709 details of spectrum, hence, it well characterizes the position of spectrum and its shifting. Curve2  
 710 describes evolution of the spectral maximum. The step shape of curve corresponds to the  
 711 fundamental property of downshifting. Opposite to the common views, development of spectrum  
 712 occurs not monotonically, but by appearance of a new maximum at lower wave number as well  
 713 as by attenuation of the previous maximum. Curve 3 describes the change of total energy  
 714  $E = E_p + E_k$ . As seen all three curves have a tendency for saturation (decrease of evolution rate).  
 715

716 The numerical experiment reproduces the case when development of wave field occurs  
 717 under the action of permanent and uniform wind. This case corresponds to JONSWAP  
 718 experiment. Despite large scatter, the data allow us to construct empirical approximations of  
 719 wave spectrum, as well as to investigate the evolution of spectrum as a function of fetch  $F$ . In  
 720 particular, it is suggested that the frequency of spectral peak changes as  $F^{-1/3}$ , while full energy  
 721 grows linearly with  $F$ . Neither of the dependences can be exact, since they do not take into  
 722 account the approaching to a stationary regime. Besides, the dependence of frequency on fetch is  
 723 singular at  $F = 0$ .

724 The value of fetch in periodic problem can be calculated by integration of peak phase  
 725 velocity  $c_p = |k|^{-1/2}$  over time.

$$F = \int_{t_0}^t c_p dt \quad (35)$$

727 The JONSWAP dependencies for the wave number of spectral peak  $k_p$  and full energy  $E$  are  
 728 shown in Fig 13 by thin curves. Dependence  $\omega_p \propto F^{1/3}$  is qualitatively valid. Dependence of the  
 729 total energy on fetch does not look like a linear one, but it is worth to note that JONSWAP  
 730 dependence is evidently inapplicable to a very small and large fetch.

731

## 732 5. Discussion

733

734 A model based on the full three-dimensional equations potential motion with free surface  
 735 was used for simulation of development of wave fields. The model is written in the surface-  
 736 following nonstationary non-orthogonal coordinate system. The details of numerical scheme and  
 737 the results of validation of the model were described in (Chalikov et al, 2014). The main  
 738 difference between the given model and HOS model (Ducroset et al, 2017) is that our model is  
 739 based on direct solution of 3-D equations for velocity potential. This approach is similar to that  
 740 developed at Technical University of Denmark (TUD, see Engsig-Karup et al, 2009). Actually,  
 741 the models developed at TUD are directed to solution of a variety of problems including such  
 742 problems as modeling of wave interaction with submerged objects and simulation of wave  
 743 regime in the basins with real shape and topography.

744 In the current paper a three-dimensional model was used for simulation of development  
 745 of wave field under the action of wind and dissipation. The input energy is described by single



746 term, i.e., surface pressure  $p$  in Eq. (4). It is traditionally assumed that the complex pressure  
747 amplitude in Fourier space is linearly connected with the complex elevation amplitude with a  
748 complex coefficient called  $\beta$  – function. Such simple formulations can be imperfect. Firstly, it is  
749 assumed that wave field is represented by superposition of linear modes with slowly changing  
750 amplitudes and phase velocity obeying the linear dispersive relation. This assumption is valid  
751 only for a low-frequency part of spectrum. In reality, the amplitudes of medium and high-  
752 frequency modes undergo fluctuations created by reversible interactions. A solid dispersion  
753 relation does not connect their phase velocities with wave number. Besides, it is also quite  
754 possible that the suggestion of linearity of the connection between pressure and elevation  
755 amplitudes is not precise, i.e.,  $\beta$  – function can depend on amplitudes of modes.

756 We are not familiar with any observation data that can be used for formulation of a more  
757 sophisticated scheme for calculation of the input energy to waves. The only method that can give  
758 more or less reliable results is mathematical modeling of the statistical structure of turbulent  
759 boundary layer above a curvilinear moving surface, which characteristics satisfy kinematic  
760 conditions. As a whole, the problem of boundary layer seems even more complicated than the  
761 wave problem itself. Some early attempts to solve this problem were made on the basis of the  
762 finite difference two-dimensional model of boundary layer written in the simple surface  
763 following coordinate (see review Chalikov, 1986). Waves were assigned as a superposition of  
764 linear modes with random phases corresponding to the empirical wave spectrum. This approach  
765 was quite accurate since it did not take into consideration the nonlinear properties of surface (for  
766 example, the sharpness of real waves and the absence of dispersive relation for waves of medium  
767 and high frequencies. The next step was formulation of coupled models for boundary layer and  
768 potential waves, both written in the conformal coordinates (Chalikov, Rainchik, 2014). The  
769 calculations showed that pressure field consists mostly of random fluctuations not directly  
770 connected with waves. A small part of these fluctuations is in phase with surface disturbances.  
771 The calculated values of  $\beta$  in Eq. (13) have large dispersion. However, since the volume of data  
772 was very large, the shape of  $\beta$ -function was found with high-level accuracy. Probably,  
773 approximation of  $\beta$  used in the current work can be considered as most adequate. We are  
774 planning additional investigations based on coupled wind-wave models. The next step in  
775 investigations of Wave boundary Layer (WBL) should use a three-dimensional LES approach.  
776 Note that even availability of large volume of data on the structure of WBL does not make the  
777 problem of parameterization of wind input in spectral wave models easily solvable, since the  
778 pressure is characterized by a broad continuous spectrum created by nonlinearity.

779 The wave breaking is obviously even more complicated than the input energy.  
780 Nevertheless, this problem can be simplified, if common ideas used in the numerical fluid  
781 mechanics are accepted. For example, in LES modeling the more or less artificial viscosity is  
782 introduced to prevent too large local velocity gradients. It is a fact that the numerical instability  
783 terminating computations precedes wave breaking. Hence, the scheme should prevent breaking  
784 approach to preserve stability of the numerical scheme. Hence, a wave model should contain the  
785 algorithms preventing appearance of too large slopes. The criterion of breaking is introduced not  
786 for recognizing of the breaking itself, but for the choice of places where it might happen (or,  
787 unfortunately, might not happen). Finally, the algorithm should produce local smoothing of  
788 elevation (and surface potential). The algorithm should be highly selective so that ‘breaking’  
789 would occur within narrow intervals and not affect the entire area. The exact criteria of breaking  
790 events (most evident of them is the breaking itself) cannot be used for parameterization of



791 breaking since in coordinate system (1) the numerical instability occurs long before breaking. In  
792 our opinion, the most sensitive parameter indicating potential instability is the curvilinearity  
793 (second derivative) of elevation.

794 In the current work, the breaking is parameterized by diffusion algorithm with the  
795 nonlinear coefficient diffusion providing high selectivity of smoothing. We admit that such  
796 approach can be realized in many different forms. The same situation is observed in a problem of  
797 turbulence modeling for parameterization of subgrid scales.

798 We can finally conclude that the physics included in the wave model is still based on a  
799 shaky ground. Nevertheless, the result of the calculations looks quite realistic, which convinces  
800 us that the approach deserves further development.

801 The numerical models of waves similar to that considered in the paper have a lot of  
802 important applications. Firstly, they are a perfect tool for development of physical  
803 parameterizations schemes in spectral wave models. Secondly, the direct model can be used in  
804 future for numerical simulation of wave processes in the basins of small and medium size. These  
805 investigations can be based on HOS model (Ducroset et al, 216) or the model used in the current  
806 paper. However, the most universal approach seems to be developed at the Technical University  
807 of Denmark (see Engsig-Karup, 2009). Any model used for a long-term simulation of wave field  
808 evolution should include the algorithms describing transformation of energy, similar to those  
809 considered in the current paper.

#### 810 **Acknowledgements**

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883

884 **Figure captions**

885

886 **Figure 1.** Probability of curvilinearity  $\eta_{\xi\xi}$ . Thick curve calculated with full 3-D model; thin  
 887 curve is a normal probability.

888 **Figure 2.** Evolution of integral characteristics of solution, rate of evolution of integral energy  
 889 multiplied by  $10^7$ ) due to: 1 – tail dissipation  $D_t$  (Eqs. 16-20); 2 – breaking dissipation  $D_b$  (Eqs.  
 890 21-24); 3 – input of energy from wind  $I$  (Eqs. 13-15); 4 – balance of energy  $I + D_t + D_b$ . Curve  
 891 5 shows the evolution of wave energy  $10^5 E$ . Vertical bars of grey color show the instantaneous  
 892 values; thick curve shows the smoothed behavior.

893 **Figure 3.** The wave spectra  $S_h(r)$  integrated over angle  $\psi$  in the polar coordinates and  
 894 averaged over consequent intervals of length about 100 units of nondimensional time  $t$ . The  
 895 spectra are growing and shifting from right to left.

896 **Figure 4** Sequence of 3-D images of  $\lg_{10}(S(k,l))$  where each panel corresponds to single curve  
 897 in Fig. 3. The left side refers to wave number  $l(-M_y \leq l \leq M_y)$  and front side – to  
 898  $k(0 \leq k \leq M_x)$

899 **Figure 5.** Sequence of 2-D images of  $\lg_{10}(S(k,l))$  averaged over consequent seven periods  
 900 length  $\Delta t = 200$ . Numbers indicate the period of averaging (first panel marked 0, refers to initial  
 901 conditions). Horizontal and vertical axes correspond to wave numbers  $k$  and  $l$  correspondingly

902 **Figure 6.** The spectrum of energy input  $I(r)$  integrated over angle  $\psi$  in the polar coordinates  
 903 and averaged over consequent intervals of length about 100 units of nondimensional time  $t$ .

904 **Figure 7.** Tail dissipation spectra  $D_t(r)$  integrated over angle  $\psi$  in the polar coordinates and  
 905 averaged over consequent intervals of length about 100 units of nondimensional time  $t$ .

906 **Figure 8,** Breaking dissipation spectra  $D_b(r)$  integrated over angle  $\psi$  in the polar coordinates  
 907 and averaged over consequent intervals of length about 100 units of nondimensional time  $t$ .

908 **Figure 9.** Example of energy input due to breaking  $D_b(x)$ .

909 **Figure 10.** Evolution of number of wave breaking events  $N_b$  expressed in percentage of the  
 910 number of grid points  $N_x \times N_y$ .

911 **Figure 11.** Sequence of wave spectra  $S_h(r)$  (thick curves) and nonlinear input term  $I(r)$  (thin  
 912 curves) averaged over consequent eight periods of length  $\Delta t = 100$  starting from 6<sup>th</sup> period.

913 **Figure 12.** Time evolution of ratio  $E_l / E_k$ .

914 **Figure 13.** Time evolution of: weighted frequency  $\omega_w$  (1) (Eq. 34); spectral peak frequency  $\omega_p$   
 915 (2); full energy  $E$  (3) (Eq. 25). Thin curves are empirical dependence for peak wave number and  
 916 energy.  $F$  is a distance passed by the spectral peak.

917