



1 **Using kinetic energy measurements from altimetry to detect shifts in the**
2 **positions of fronts in the Southern Ocean**

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7 **Abstract.** A novel analysis is performed utilizing cross-track kinetic energy (CKE) computed
8 from sea surface height anomalies derived from along-track satellite altimetry. The mid-point of
9 enhanced kinetic energy averaged over three-year periods from 1993 to 2015 is determined
10 across the Southern Ocean and examined to detect shifts in frontal positions, based on previous
11 observations that kinetic energy is largest along fronts and jets in the Antarctic Circumpolar
12 Current system. It is demonstrated that although the CKE does not represent the full eddy kinetic
13 energy (computed from crossovers), the shape of the enhanced regions along groundtracks is the
14 same, and CKE has a much finer spatial sampling of 6.9 km. Results indicate no significant shift
15 in the front positions across the Southern Ocean, on average, although there are some localized,
16 large movements. This is consistent with other studies utilizing sea surface temperature gradients
17 and the latitude of mean transport, but inconsistent with studies utilizing the movement of
18 contours of dynamic topography.



19 1. INTRODUCTION

20 There is as much we don't know about the circulation of the Southern Ocean as we do.
21 Although the current system is routinely called the Antarctic Circumpolar Current (ACC), it
22 consists of several fronts with distinct water properties to the north and south of the fronts
23 (Nowlin and Clifford, 1982; Orsi et al., 1995; Belkin and Gordon, 1996). The most significant of
24 these, responsible for the majority of the ACC volume transport (e.g., Cunningham et al., 2003),
25 are the Subantarctic Front (SAF) and the Polar Front (PF). However, we now know the
26 circulation along these fronts is not a well-defined single current, but instead consists of
27 anywhere from three to four narrow jets per front that are highly variable in strength and location
28 (Sokolov and Rintoul, 2007, 2009a, 2009b; Sallee et al., 2008; Thompson et al., 2010;
29 Thompson and Richards, 2011; Langlais et al., 2011; Graham et al., 2012; Gille, 2014; Kim and
30 Orsi, 2014).

31 An open question is how the fronts/jets that comprise the ACC will respond in a warming
32 climate. Analysis of climate models suggests that as the atmosphere warms, the winds that drive
33 the fronts/jets of the ACC will migrate south (e.g., Fyfe and Saenko, 2006; Swart and Fyfe,
34 2012). It should be noted, however, that the mean position of the southern hemisphere westerlies
35 in the models lies significantly equatorward of the true position. Thus, it is not entirely clear
36 whether the model is predicting a true shift in the wind position, or whether the model has not
37 yet reached equilibrium with winds in the proper location.

38 Still, based on these model results, researchers have been testing a hypothesis that as winds
39 in the Southern Ocean shift south, the frontal positions and jets will also migrate south. So far,
40 the results are mixed. Sokolov and Rintoul (2009a, 2009b) proposed a method based on finding
41 the contour of a mean dynamic topography associated with the steepest meridional gradient of



42 dynamic topography (i.e., zonal geostrophic current) and associate this with a front or jet. They
43 could then track the latitude of that contour over time to deduce if the front had moved. When
44 averaged over all longitudes, their analysis indicated that the SAF and PF had both moved south
45 by approximately 60 km over 15 years (Sokolov and Rintoul, 2009b).

46 While this method is nice because it allows one to determine a front position even when
47 gradients are small due to continuity of the contour line and is less contaminated with errors such
48 as those based on methods that examine gradients alone (e.g., Chapman, 2014), it has several
49 issues because of several assumptions made. First, it assumes that the average of the position
50 shift of the contours across all longitudes represents the shift at all longitudes. Kim and Orsi
51 (2014) recently considered this assumption and found that while the average frontal position
52 indicates a strong southward shift, this is due to substantial shifts only in the Indian Ocean sector.
53 There are no significant shifts throughout the Pacific or Atlantic Ocean sectors.

54 The second assumption assumes that a contour of dynamic topography shifting south is
55 unique to a front/jet axis moving south. This is not true. Gille (2014) demonstrated that all
56 contours in the Southern Ocean have shifted south on average, and that this follows from the
57 observed rise in sea level – as the sea surface height rises, the contours will appear to shift south.
58 Gille (2014) used a different measure to determine the position of the ACC fronts, based on the
59 latitude of the mean surface transport of the ACC measured by altimetry, and found no
60 significant shift but considerable interannual variability.

61 The third assumption is that the front/jet position has shifted south but that the magnitude and
62 width of the jet has not changed. The problem with this assumption is demonstrated in **Figure 1**,
63 where we show the mean dynamic topography (MDT) from two jet scenarios: 1) where the peak
64 of two Gaussian shaped jets have shifted south, and 2) where the peak has not shifted, but the



65 magnitude has decreased, the width has broadened, and the shape has become slightly skewed.
66 Although the resulting MDT profiles are not identical, they are similar, and both suggest a
67 southward movement of MDT contours.

68 Researchers using other methods also find little or no southern migration of the fronts or jets
69 in the Southern Ocean as a whole. Graham et al. (2012) used a high-resolution model to show
70 that the Polar Front and Subantarctic Front are constrained by bathymetry, even in increasing and
71 shifting winds. Freeman et al. (2016) used weekly estimates of the Polar Front position
72 determined from satellite sea surface temperature (SST) measurements to show no significant
73 southward shift, except in the Indian Ocean (similar to Kim and Orsi, 2014) and a northward
74 shift in part of the south Pacific.

75 Here, we will utilize a new method to study the position of the fronts in the Southern Ocean,
76 based on tracking the location of eddy kinetic energy (EKE) measured by altimetry. It is known
77 from modeling studies that the front positions are associated with increased EKE, due to
78 instabilities in the jets and interactions with bathymetry (Thompson et al., 2010; Thompson and
79 Richards, 2011). Thus, it is reasonable to assume that if the front position has shifted that the
80 region of high EKE should also shift by a comparable amount.

81

82 **2. DATA AND METHODS**

83 We will utilize geostrophic surface current anomalies computed from the 23-year record
84 from the TOPEX/Poseidon, Jason-1, and Jason-2 1-Hz sea surface height (SSH) data along each
85 groundtrack in the Southern Ocean (**Figure 2**). We utilize this record rather than the gridded
86 products based on mapping SSH from multiple altimeters (e.g., Ducet et al., 2000), because the
87 along-track data high have a finer resolution in space (6.9 km along the groundtrack) and Hogg



88 et al. (2015) demonstrated that the mapped altimetry data underestimated EKE due to attenuation
89 from the large-scale correlation functions used in the mapping procedure.

90 The altimetry data used are from three separate altimeter missions: TOPEX/Poseidon
91 (January 1993 – January 2002), Jason-1 (February 2002 – July 2008), and Jason-2 (August 2008
92 – December 2015). Because the official TOPEX/Poseidon (T/P) geophysical data records
93 (GDRs) have not been updated since the late 1990s, we utilize the corrected data products from
94 the Integrated Multi-Mission Ocean Altimeter Data for Climate Research provided by Beckley et
95 al. (2010) at the NASA PO.DAAC site ([https://podaac.jpl.nasa.gov/Integrated_Multi-](https://podaac.jpl.nasa.gov/Integrated_Multi-Mission_Ocean_AltimeterData)
96 [Mission_Ocean_AltimeterData](https://podaac.jpl.nasa.gov/Integrated_Mission_Ocean_AltimeterData)). Jason-1 data are from the GDR-C version and were
97 downloaded from the NASA PO.DAAC site in June 2010. Jason-2 are from the GDR-D version
98 and were downloaded from NOAA NODC (<ftp://ftp.nodc.noaa.gov/pub/data.nodc/jason2>)
99 between August 2012 and June 2016. We utilize the 1-Hz along-track data from the three
100 altimeters and interpolate to the same fixed ground track utilizing the DTU10 mean sea surface
101 model (Andersen and Knudsen, 2009;
102 http://www.space.dtu.dk/english/Research/Scientific_data_and_models/downloaddata). All
103 recommended geophysical and surface corrections (e.g., water vapor, ionosphere, sea state bias,
104 ocean tides, inverted barometer, etc) have been applied. SSH anomalies are computed by
105 subtracting the DTU10 mean sea surface model from the interpolated SSH data.

106 The time-varying, anomalous geostrophic current normal to the groundtrack (u_T) can be
107 computed directly from the derivative of the SSH anomaly (η) along the ground-track distance
108 (dr) from

$$109 \quad u_T = \frac{g}{f} \frac{d\eta}{dr}, \quad (1)$$



110 where g is the acceleration due to gravity, and f is the Coriolis parameter. This cross-track
111 current is a projection of both the zonal (u) and meridional (v) components of the full velocity
112 field. However, neither u nor v can be determined unambiguously from u_T without significant
113 simplifying assumptions. Here, we merely examine the variability of u_T without making any
114 assumptions concerning how it may be related to the full velocity, or u and v .

115 Because derivatives have to be computed numerically (here, center-differences are used) and
116 η contains significant noise at the 1 Hz sampling-rate of the altimeters, we chose to optimally
117 interpolate η along-track using a model of the covariance of the signal and error. We used the
118 method of Wunsch (2006, Chapter 3) and a covariance function modeled as a Gaussian with a
119 roll-off of 98 km and random noise of 2 cm, which was determined from the autocovariance of
120 all TOPEX/Poseidon, Jason-1, and Jason-2 SSHA data from 1993-2015 between 40°S and 65°S.

121 Once $u_T(t)$ was computed at each 1-sec bin along the groundtracks in Figure 2 for each 10-
122 day repeat cycle, the cross-track kinetic energy (CKE) was computed as $\text{CKE}(x,t) = 0.5 u_T(x,t)^2$,
123 where x here is used to denote a generic 1-sec bin along the ground track. These CKE values
124 were averaged over the entire 23-year record and examined for each groundtrack segment (both
125 ascending and descending) to judge where CKE was exceptionally high (Figure 3). We also
126 computed CKE using the raw values of η with no optimal interpolation and compared to that
127 computed with optimal interpolation. The locations of high CKE were the same, although values
128 were slightly higher with the unsmoothed data. The quiescent regions of the ocean also showed
129 considerably more noise, making it more difficult to determine boundaries of elevated CKE. For
130 this reason, the values determined from the optimally interpolated data were used.

131 Several criteria were utilized to quantify where the high CKE values were considered to be
132 associated with fronts or jets. First, we constrained the southern boundary to be 5° south of the



133 Orsi et al. (1995) values of the PF and the northern boundary to be 5° north of the SAF.
134 Secondly, we used a lower-limit for CKE of $200 \text{ cm}^2 \text{ s}^{-2}$ for detection; only values higher than
135 this were considered high and likely associated with a front. For the example shown in Figure 3
136 (from a track in the south Indian Ocean), there is a region of CKE greater than $200 \text{ cm}^2 \text{ s}^{-2}$
137 between 55°S and 47°S . The value of $200 \text{ cm}^2 \text{ s}^{-2}$ was chosen over a lower value because it was
138 found that the regions selected with the $200 \text{ cm}^2 \text{ s}^{-2}$ were consistent from one 10-day period to
139 another. Using lower minimum values resulted in areas of high CKE activity being detected in
140 some cycles, but not others. An example is the small rise in mean CKE at 41°S in Figure 3.
141 Although the mean CKE is about $150 \text{ cm}^2 \text{ s}^{-2}$, suggestive of a jet, this all but disappears for some
142 periods of time. This behavior of jets has been observed in models (e.g., Thompson and
143 Richards, 2011). Thus, the minimum CKE was selected to be as conservative as possible.

144 The mean CKE profile pictured in Figure 3 has multiple local maxima, most likely associated
145 with the narrow jets that surround the front. Most of the other profiles examined had similar
146 features. We initially tried tracking each of the maxima, but that quickly became complicated
147 because sometimes the four or local maxima would become five, or even just one. This is likely
148 due to the instability of the jets around the front. In order to ensure consistency and assuming the
149 mean of the region of high CKE followed the front position, the half-power point, or centroid, of
150 the CKE bump was computed. This is similar in principle to the computation made by Gille
151 (2014) of the latitude of the mean ACC transport, except here we are focused on mean of the
152 CKE around a particular front, not over all the fronts of the ACC.

153 To compute the half-power point, a southern and northern boundary of the CKE bump had to
154 be determined. These were computed by first finding the maximum of CKE in the bump, then
155 finding the first value to the north just below 25% of that peak along with the similar value to the



156 south (shown in Figure 3). These were selected as the north (x_{north}) and south (x_{south}) boundaries
157 for the calculation of the half-power point (x_{mid}) so that

158
$$\int_{x_{south}}^{x_{mid}} CKE(x) dx = \frac{1}{2} \int_{x_{south}}^{x_{north}} CKE(x) dx . \quad (2)$$

159 Values other than 25% of the peak were tested. Using value greater than this, up to 50%, resulted
160 in no significant difference in the half-power point. Using values smaller resulted in some
161 boundaries not being defined. Thus, 25% of peak CKE was considered reasonable. If multiple
162 regions of enhanced CKE were found along the same track, this process was carried out for each
163 of them. This was done for all the mean CKE profiles to establish the mean locations of the
164 fronts between 1993 and 2015.

165 A similar procedure was done for CKE averaged over discrete 3-year intervals, starting in
166 1993. This provided 8 distinct samples of CKE for each groundtrack from which to deduce shifts
167 in the half-power point.

168 3. RESULTS AND ANALYSIS

169 The first thing tested was how well CKE represented the full EKE, where $EKE = 0.5(u^2 +$
170 $v^2)$. One can calculate both components of the velocity at crossover points, where the ascending
171 and descending groundtracks cross, under the assumption that the velocity field has not changed
172 significantly between the times of the groundtracks (Parke et al., 1987). At high latitudes of the
173 Southern Ocean, the time separation between ascending and descending passes is less than 3
174 days for 78% of the crossovers, so this is a reasonable assumption.



175 Knowing the groundtrack angle with the north meridian (θ) one can compute the zonal
 176 ($d\eta/dy$) and meridional gradients ($d\eta/dx$) of SSHA directly from the gradients of SSHA for the
 177 ascending pass ($d\eta/dr_{asc}$) and descending pass ($d\eta/dr_{des}$) using simple geometry

$$178 \quad \frac{d\eta}{dy} = \frac{\left[\frac{d\eta}{dr_{asc}} - \frac{d\eta}{dr_{des}} \right]}{2 \sin \theta}, \quad \frac{d\eta}{dx} = \frac{\left[\frac{d\eta}{dr_{asc}} + \frac{d\eta}{dr_{des}} \right]}{2 \cos \theta}, \quad (3)$$

179 noting that this formulation assumes the gradients represent the derivative of the northern SSHA
 180 relative to the southern SSHA (for both the ascending and descending passes). Once this is
 181 computed, the velocities can be computed directly from the zonal and meridional gradients:

$$182 \quad u = -\frac{g}{f} \frac{d\eta}{dy}, \quad v = \frac{g}{f} \frac{d\eta}{dx}. \quad (4)$$

183 Although CKE is lower than EKE along all groundtracks (see Figure 4 for examples), the
 184 pattern of KE rise then fall is virtually identical. CKE, however, has the benefit of higher and
 185 more regular sampling. Thus, we conclude CKE is a reasonable proxy for locating front positions
 186 even though it may not be useful for quantifying the full energy of the anomalous currents.

187 Three general types of enhanced CKE were found (Figure 4). In most regions, the “bump” in
 188 CKE is more or less symmetrical. In several, however, the bump is skewed, with a long rise then
 189 a sharp drop-off. Finally, there were also a few cases where two distinct regions of enhanced
 190 CKE were identified. In all cases, the shape of the CKE closely followed that of EKE, although
 191 the amplitude was attenuated, by anywhere from 25-50%.

192 Figure 2 shows the locations of the half-power points determined from the mean CKE
 193 profiles. Although a potential front position is not found along every groundtrack, they are
 194 typically found in areas where currents are strong and close to perpendicular with the ascending



195 or descending groundtracks. It is also clear that there are more sites found along the SAF than the
196 PF. A few are found between the two fronts, but this is likely due to either errors in the older
197 front database of Orsi et al. (1995), changes since the hydrographic data used in that study were
198 collected, or jets between the fronts. One location is found substantially south of the Polar Front,
199 and is likely related to the weaker Southern ACC front.

200 Figure 5 shows the variability in the enhanced CKE along one groundtrack for the three-year
201 intervals compared to the mean, and demonstrates the problem with separating the positions of
202 the highly variable jets from the front. For example, in 1999-2001, there were 6 distinct local
203 maxima of CKE (indicated by blue dots), suggesting jets. In 2002-2004, there were also 6
204 (indicated by orange dots), but in different locations. Does this indicate a movement of the jet, or
205 a new one? That is impossible to say without making considerable assumptions, which is why we
206 only consider the movement of the half-power point of enhanced CKE as defined in Section 2.

207 To quantify potential movement, a linear trend is fit to the 8 estimations of the half-power
208 point from 1993-2015 for each location shown in Figure 2. Analysis of the residuals about the
209 trend indicated they were random (lag-1 autocorrelation < 0.1 for all cases), so standard error
210 was computed by scaling the formal error from ordinary least squares by the standard deviation
211 of the residuals. This was also scaled up to account for the degrees of freedom lost by estimating
212 the trend (making the effective degrees of freedom 6) by scaling by $\sqrt{8/6}$. Finally, the 90%
213 confidence interval was computed by scaling by 1.94 for 6 effective degrees of freedom
214 assuming a normal t-distribution.

215 The results indicate considerable regional variability in the change of the half-power point
216 over 23 years, with large uncertainty bars (Figure 6). This is due to the substantial temporal
217 variability in the position, which can be seen somewhat in Figure 5, where the leading edge of



218 the CKE bump varies by over 1 degree of latitude (over 100 km) between 1993-1995 and 2011-
219 2012. To better see significant changes outside the uncertainty (90% confidence) interval, one
220 can compute the signal to noise ratio (SNR = trend/uncertainty). Examining this (Figure 7), one
221 can see there are some regions where the half-power point has moved southward by a significant
222 distance over the last 23 years (13.6% of points), but there are also points where it has moved
223 north (9.6%). For the majority of points (76.8%), there is no significant change

224

225 **4. DISCUSSION AND CONCLUSIONS**

226 The results from the analysis of the positions of enhanced kinetic energy suggest no overall
227 shift in the front positions across the Southern Ocean, but some large, localized movements. The
228 region indicative of some southward shifts between 90°E and 170°E is in approximately the
229 same area where Kim and Orsi (2014) and Freeman et al. (2016) also reported large shifts.
230 However Freeman et al. only examined the Polar front, and Kim and Orsi (2014) only found
231 large shifts in the PF and the southern ACC front. They found shifts of order 50-100 km in the
232 SAF where the points in this study cluster. Our results suggest considerably larger shifts in some
233 areas between 90°E and 170°E, although the overall average (-29 km per decade, or -66.7 km in
234 23 years), is consistent with what Kim and Orsi (2014) found.

235 Kim and Orsi (2014) and Freeman et al. (2016) also found slight northward shifts in the front
236 positions in the southeast Pacific, between 200°E-270°E. We also find some locations in this
237 region with a significant northward shift in the SAF. Kim and Orsi (2014) found the shift of the
238 SAF was about 30-40 km between 1992 and 2011. Our results suggest larger shifts in some
239 areas; averaged over the area, our results are 46 km per decade to the north, or 106 km from



240 1993-2015, which is consistent with the average over the region computed by Freeman et al.
241 (2016) from sea surface temperature data, but for the Polar Front.

242 Overall, this study supports the recent studies by Kim and Orsi (2014), Gille (2014), and
243 Freeman et al. (2016) that the frontal positions of the ACC are highly variable in time and that
244 there is no statistically significant shift in the fronts to the south on average. This study utilized a
245 novel technique to reach this conclusion, which adds to the robustness of evidence that there has
246 not been a shift in the frontal positions. Thus, while the fronts may eventually shift south in a
247 warming climate, there is no strong evidence that it is happening at the moment.

248 The only evidence supporting this hypothesis comes from mapping the location of contours
249 of constant dynamic topography over time (e.g., Sokolov and Rintoul, 2009b; Kim and Orsi,
250 2014). As Gille (2014) argued and as we have demonstrated based on a simple thought
251 experiment (Figure 1), there are other equally plausible explanations for the apparent southern
252 shift of the contours. Considering that three different techniques – location of mean transport
253 (Gille, 2014), maximum SST gradients (Freeman et al., 2016), and location of enhanced kinetic
254 energy (this study) – all agree that the fronts have not moved significantly on average, one has to
255 conclude that the method of using dynamic topography contours to detect changes in front
256 position is flawed. As Gille (2014) concluded, the most likely explanation for the signal is the
257 observed rise in sea level over the region due to warming of water in the region.

258



259

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263 REFERENCES

- 264 Andersen O B, and Knudsen P: DNSC08 mean sea surface and mean dynamic topography
265 models, *J. Geophys. Res.*, 114, C11001, doi:10.1029/2008JC005179, 2009.
- 266 Beckley, B.D., Zelensky, N.P., Holmes, S.A., Lemoine, F.G., Ray, R.D., Mitchum, G.T., Desai,
267 S., and Brown, S. T.: Assessment of the Jason-2 Extension to the TOPEX/Poseidon, Jason-1
268 Sea-Surface Height Time Series for Global Mean Sea Level Monitoring, *Marine Geodesy*,
269 33(S1): 447-471, Supplemental Issue on OSTM/Jason-2 calibration/validation, Vol. 1, DOI:
270 10.1080/01490419.2010.491029, 2010
- 271 Belkin, I. M., and Gordon, A. L.: Southern Ocean fronts from the Greenwich meridian to
272 Tasmania, *J. Geophys. Res.*, 101, 3675–3696, 1996.
- 273 Chapman, C. C.: Southern Ocean jets and how to find them: Improving and comparing common
274 jet detection methods, *J. Geophys. Res. Oceans*, 119, 4318–4339,
275 doi:10.1002/2014JC009810, 2014.
- 276 Cunningham, S. A., Alderson, S. G. , King, B. A., and Brandon, M. A.: Transport and variability
277 of the Antarctic Circumpolar Current in Drake Passage, *J. Geophys. Res.*, 108(C5), 8084,
278 doi:10.1029/2001JC001147, 2003
- 279 Dong, S., Sprintall, J., and Gille, S. T. : Location of the Antarctic Polar Front from AMSR-E
280 Satellite Sea Surface Temperature measurements, *J. Phys. Oceanogr.*, 36, 2075–2089,
281 doi:10.1175/JPO2973.1, 2006.
- 282 Ducet, N., Le Traon, P.-Y., Reverdin, G.: Global high resolution mapping of ocean circulation
283 from TOPEX/Poseidon and ERS-1 and -2. *Journal of Geophysical Research* 105 (C8),
284 19477–19498, 2000.
- 285 Fyfe, J. C., and Saenko, O. A.: Simulated changes in the extratropical Southern Hemisphere
286 winds and currents, *Geophys. Res. Lett.*, 33, L06701, doi:10.1029/2005GL025332, 2006.
- 287 Gille, S. T.: Meridional displacement of the Antarctic Circumpolar Current, *Philos. Trans. R.*
288 *Soc. A*, 372, 20130273, doi:10.1098/rsta.2013.0273, 2014.
- 289 Graham, R. M., De Boer, A. M., Heywood, K. J., Chapman, M. R., and Stevens, D. P.: Southern
290 Ocean fronts: Controlled by wind or topography?, *J. Geophys. Res. Oceans*, 117, C08018,
291 doi:10.1029/2012JC007887, 2012.
- 292 Hogg, A. McC., Meredith, M. P., Chambers, D. P., Abrahamsen, E. P., Hughes, C. W., and
293 Morrison, A. K.: Recent trends in the Southern Ocean eddy field, *J. Geophys. Res. Oceans*,
294 120, 257–267, doi:10.1002/2014JC010470, 2015.
- 295 Kim, Y. S., and Orsi, A. H.: On the variability of Antarctic Circumpolar Current fronts inferred
296 from 1992-2011 altimetry, *J. Phys. Oceanogr.*, 44, 3054–3071, doi:10.1175/JPO-D-13-
297 0217.1, 2014.



- 298 Langlais, C., Rintoul, S. R., and Schiller, A.: Variability and mesoscale activity of the Southern
299 Ocean fronts: Identification of a circumpolar coordinate system, *Ocean Modell.*, 39, 79–96,
300 doi:10.1016/j.ocemod.2011.04.010, 2011.
- 301 Moore, J. K., Abbott, M. R., and Richman, J. G.: Location and dynamics of the Antarctic Polar
302 Front from satellite sea surface temperature data, *J. Geophys. Res.*, 104, 3059–3073,
303 doi:10.1029/1998JC900032, 1999.
- 304 Nowlin, W. D., and Clifford, M.: The kinematic and thermohaline zonation of the Antarctic
305 Circumpolar current at Drake Passage, *J. Mar. res.*, 40, 481–507, 1995.
- 306 Orsi, A. H., Whitworth III, T., and Nowlin Jr., W. D.: On the meridional extent and fronts of the
307 Antarctic Circumpolar Current, *Deep Sea Res., Part I*, 42(5), 641–673, doi:10.1016/0967-
308 0637(95)00021-W, 1995.
- 309 Parke, M. E., Stewart, R. H., Farless, D. L., and Cartwright, D. E.: On the choice of orbits for an
310 altimetric satellite to study ocean circulation and tides, *J. Geophys. Res.*, 92, 11693–11707,
311 1987.
- 312 Sallee, J. B., Speer, K., and Morrow, R.: Response of the Antarctic Circumpolar Current to
313 atmospheric variability, *J. Clim.*, 21(12), 3020–3039, doi:10.1175/2007JCLI1702.1, 2008.
- 314 Sokolov, S., and Rintoul, S. R.: Multiple jets of the Antarctic Circumpolar Current south of
315 Australia, *J. Phys. Oceanogr.*, 37, 1394–1412, doi:10.1175/JPO3111.1, 2007.
- 316 Sokolov, S., and Rintoul, S. R.: Circumpolar structure and distribution of the Antarctic
317 Circumpolar Current fronts: 1. Mean circumpolar paths, *J. Geophys. Res.*, 114, C11018,
318 doi:10.1029/2008JC005108, 2009a.
- 319 Sokolov, S., and Rintoul, S. R.: Circumpolar structure and distribution of the Antarctic
320 Circumpolar Current fronts: 2. Variability and relationship to sea surface height, *J. Geophys.*
321 *Res.*, 14, C11019, doi:10.1029/2008JC005248, 2009b.
- 322 Swart, N., and Fyfe, J. C.: Observed and simulated changes in the Southern Hemisphere surface
323 westerly wind-stress. *Geophys. Res. Lett.*, 39, L16711, doi:10.1029/2012GL052810, 2012.
- 324 Thompson, A. F., Haynes, P. H., Wilson, C., and Richards, K. J.: Rapid Southern Ocean front
325 transitions in an eddy-resolving ocean GCM, *Geophys. Res. Lett.*, 37, L23602,
326 doi:10.1029/2010GL045386, 2010.
- 327 Thompson, A. F., and Richards, K. J.: Low frequency variability of Southern Ocean jets, *J.*
328 *Geophys. Res.*, 116, C09022, doi:10.1029/2010JC006749, 2011.
- 329 Wunsch, C.: *The Ocean Circulation Inverse Problem*, 458 pp., Cambridge Univ. Press, Cambridge, Mass.



333 **Figure Captions**

334

335 **Figure 1.** a) Mean dynamic topography in the Southern Ocean along a north-south meridian for
336 three scenarios, and b) the corresponding geostrophic velocity, with positive values indicating
337 eastward flow. The scenarios are: an initial state (dashed black line), a shift of the two fronts
338 south by 60 km with no change in magnitude or shape of the currents (red line), and no shift of
339 the mean of the current, but a change in the magnitude and shape (blue line).

340

341 **Figure 2.** Positions of the T/P, Jason-1, and Jason-2 groundtracks used for this study (black
342 lines), the approximate locations of the Subantarctic Front (red line) and the Polar Front (blue
343 line), as well as locations where enhanced cross-track kinetic energy was found (orange dots).
344 The front positions are from Orsi et al. (1995).

345

346 **Figure 3.** An example profile of mean CKE along a ground track in the southern Indian Ocean,
347 demonstrating the location of the half-power point and the locations of the southern and northern
348 boundaries of the CKE bump. See text for details of the computations.

349

350 **Figure 4.** Examples of the three types of CKE profiles found (black lines), along with the value
351 of the full EKE computed at crossover points.

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353 **Figure 5.** Examples of the time-variable CKE profiles found, averaged over 3-year periods. The
354 ground track is the same shown Figure 3 for the southern Indian Ocean. Colored dots indicate the
355 location of probable jets in 2002-2004 (orange) and in 1999-2001 (blue).

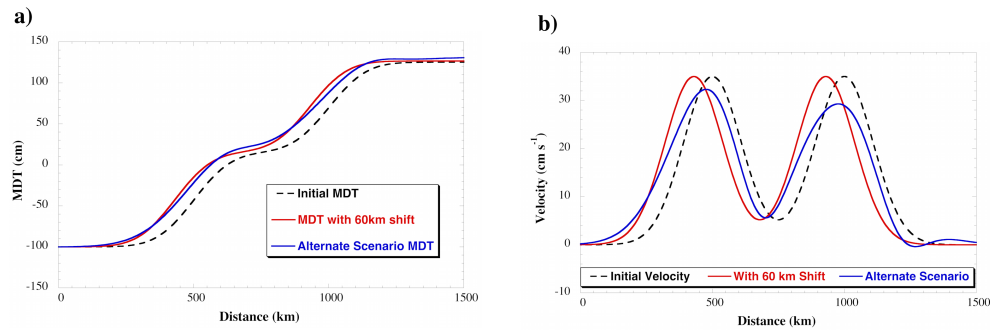
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357 **Figure 6.** Estimated trend in the half-power point of CKE for each location shown in Figure 2, as
358 a function of latitude. Error bars represent the 90% confidence interval.

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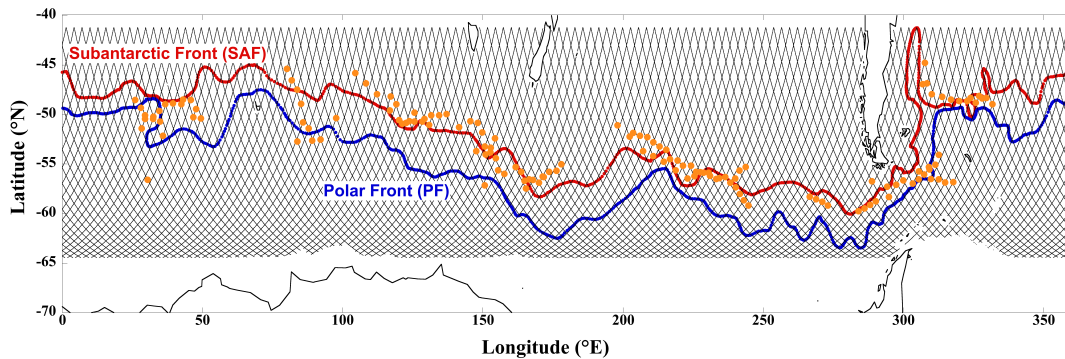
360 **Figure 7.** SNR (trend/error in Figure 6). Values larger than 1 indicate a statistically significant
361 northern shift. Values smaller than -1 indicate a statistically significant southern shift. Values
362 between ± 1 indicate no statistically significant shift.

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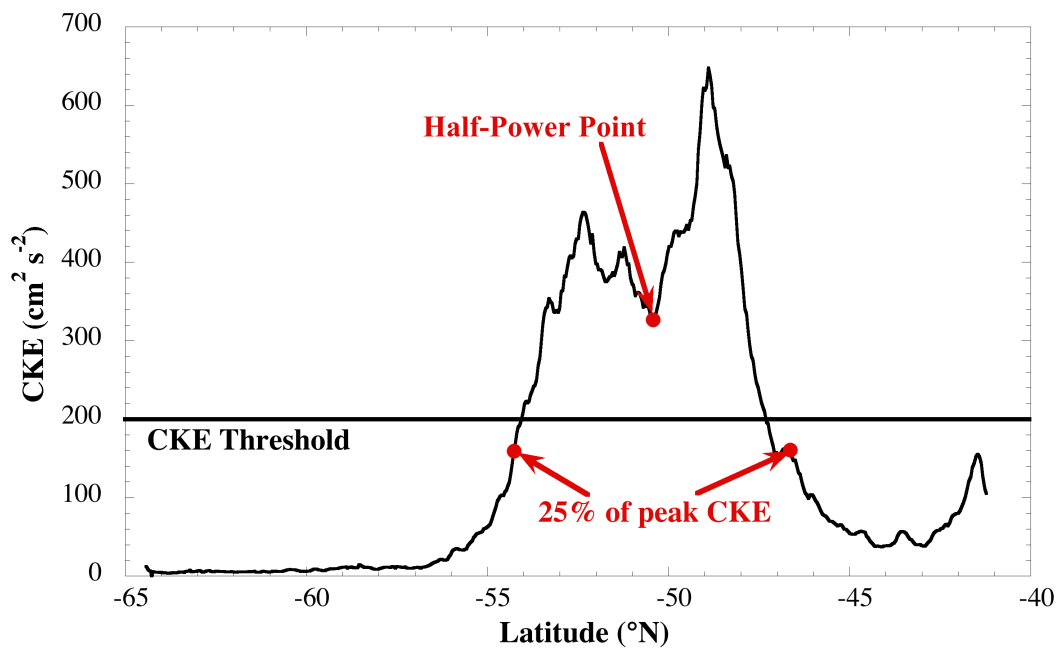
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Figure 1. a) Mean dynamic topography in the Southern Ocean along a north-south meridian for three scenarios, and b) the corresponding geostrophic velocity, with positive values indicating eastward flow. The scenarios are: an initial state (dashed black line), a shift of the two fronts south by 60 km with no change in magnitude or shape of the currents (red line), and no shift of the mean of the current, but a change in the magnitude and shape (blue line).



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Figure 2. Positions of the T/P, Jason-1, and Jason-2 groundtracks used for this study (black lines), the approximate locations of the Subantarctic Front (red line) and the Polar Front (blue line), as well as locations where enhanced cross-track kinetic energy was found (orange dots). The front positions are from Orsi et al. (1995).

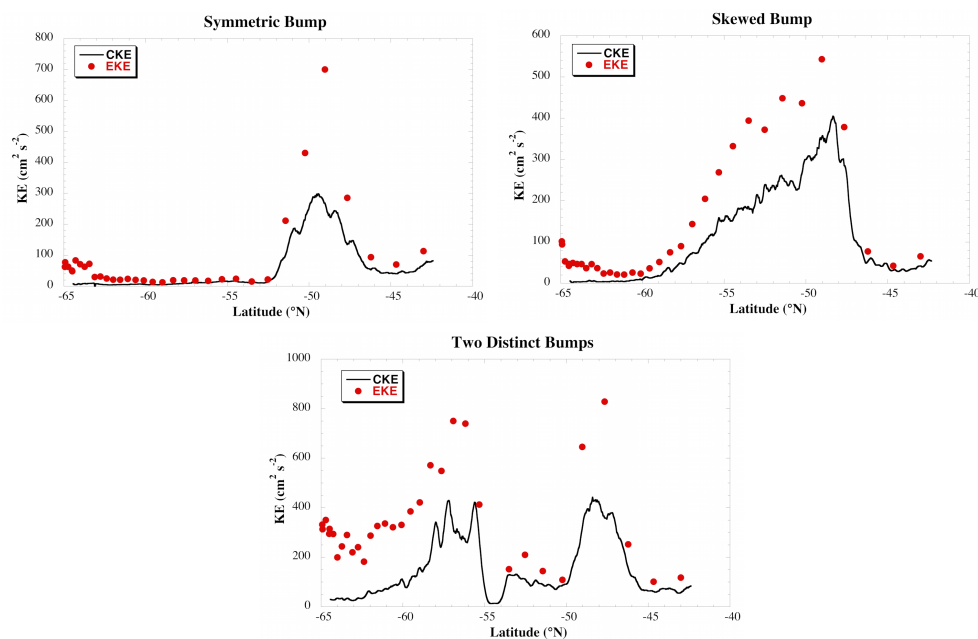


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Figure 3. An example profile of mean CKE along a ground track in the southern Indian Ocean, demonstrating the location of the half-power point and the locations of the southern and northern boundaries of the CKE bump. See text for details of the computations.



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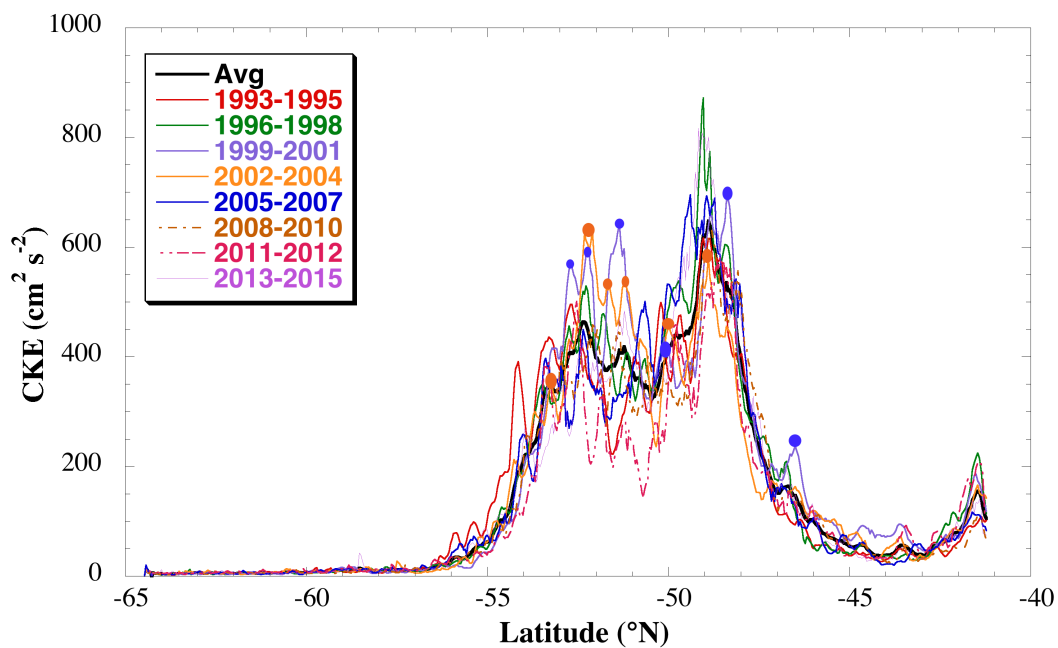


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Figure 4. Examples of the three types of CKE profiles found (black lines), along with the value of the full EKE computed at crossover points.

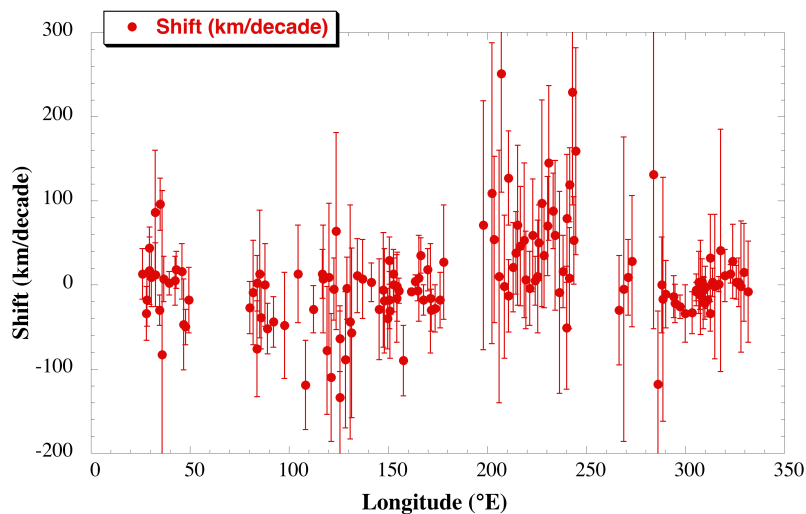


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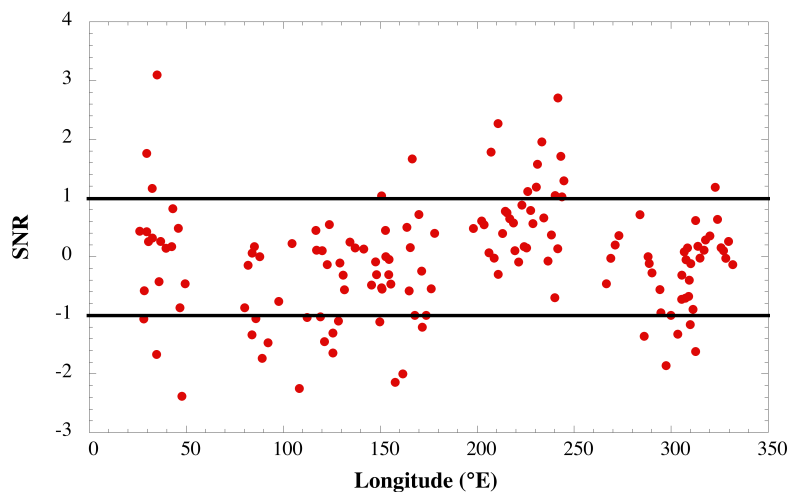
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Figure 6. Estimated trend in the half-power point of CKE for each location shown in Figure 2, as a function of latitude. Error bars represent the 90% confidence interval.



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Figure 7. SNR (trend/error in Figure 6). Values larger than 1 indicate a statistically significant northern shift. Values smaller than -1 indicate a statistically significant southern shift. Values between ± 1 indicate no statistically significant shift.